# SEMICLASSICAL APPROACHES TO QUANTAL ANDREEV BILLIARDS AND 

 OTHER TOPICS IN INHOMOGENEOUS SUPERCONDUCTIVITYBY<br>İNANÇ ADAGIDELI

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## THESIS

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# SEMICLASSICAL APPROACHES TO QUANTAL ANDREEV BILLIARDS AND OTHER TOPICS IN INHOMOGENEOUS SUPERCONDUCTIVITY 

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Semiclassical approaches to inhomogeneous superconductivity are explored within three settings. The first is that of Andreev billiards, which are finite, arbitrarilyshaped, normal-state regions, surrounded by superconductor. A theoretical framework for the investigation of the short-wave quantal properties of quasiparticles in Andreev billiards is developed. Among the central results are two semiclassical trace formulas for the density of states. The first, a lower-resolution formula, corresponds to the well-known quasiclassical approximation. The second, a higher-resolution formula, allows the density of states to be expressed in terms of: (i) An explicit formula, valid in the short-wave limit, for billiards of arbitrary shape and dimensionality; and (ii) Higher-resolution corrections, expressed as a sum over periodic orbits that creep around the billiard boundary. The second involves quasiparticle motion near an extended impurity in a d-wave superconductor. Via an extension of the quasiclassical approximation, the idea that sign-variations in the superconducting pair-potential lead to low-energy states is extended beyond its original setting of boundary scattering to the broader context of scattering by general single-particle potentials. The index-theoretic origin of these states is exhibited via a connection with Witten's supersymmetric quantum-mechanical model. The third concerns an extended impurity in a cuprate superconductor in the pseudogap regime. It is shown that if the d-wave correlations persist into the pseudogap regime, but with pair-potential phase-fluctuations that destroy long-range ordering, then the low-energy states near extended impurities in the superconducting state should persist as resonances in the pseudogap regime. Thus, such states can serve as a probe of d-wave correlations in the pseudogap regime.

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## Contents

Chapter
1 Introduction ..... 1
2 Introduction to Andreev Billiards ..... 5
3 Andreev billiards ..... 10
3.1 Idealization of the physical system ..... 10
3.2 Green function and density of states ..... 14
3.2.1 Green function for the Bogoliubov-De Gennes equation ..... 15
3.2.2 Connection between Bogoliubov-De Gennes Green function and density of states ..... 15
3.2.3 Fundamental Green function for a homogeneous normal region ..... 17
3.2.4 Fundamental Green function for a homogeneous superconduct- ing region ..... 19
3.3 Andreev's quasiclassical aproximation scheme ..... 21
4 Multiple scattering expansion ..... 24
4.1 Matching conditions and boundary integral equations for the Bogoliubov-De Gennes Green function ..... 25
4.2 Reorganized multiple scattering expansion ..... 28
4.3 Effective boundary conditions ..... 32
5 Asymptotics of the multiple scattering expansion ..... 34
5.1 Classical dynamics in Andreev billiards ..... 35
5.2 Asymptotics of the fundamental Green functions and their derivatives ..... 40
5.3 Asymptotic expansion for the quantal amplitude ..... 43
5.4 Integrating out propagation in the superconducting region: Effective reflection ..... 50
6 Density of states oscillations ..... 58
6.1 Scheme A: Andreev approximation ..... 61
6.2 Scheme B: Mesoscale oscillations beyond the resolution of the Andreev approximation ..... 68
6.2.1 Perfectly Charge-Interconverting Model ..... 69
6.2.2 Stationary chords ..... 70
6.2.3 Creeping orbits ..... 74
6.2.4 Incorporating ordinary reflection ..... 81
7 Concluding remarks; Perspectives ..... 89
8 Intermezzo: d-wave superconductivity ..... 93
9 Low-energy quasiparticle states near extended scatterers in d-wave superconductors ..... 95
9.1 Introduction ..... 95
9.2 Bogoliubov-de Gennes eigenproblem ..... 98
9.3 Andreev's approximation for a strong single-particle potential ..... 99
9.4 Zero-energy states ..... 101
9.5 Connection with Witten's model of supersymmetric quantum mechan- ics and index theory ..... 103
9.6 Discussion and outlook ..... 103
10 Probing d-wave pairing correlations in the pseudogap regime of the cuprate superconductors via low-energy states near impurities ..... 106
10.1 Introduction ..... 106
10.2 Model of cuprate superconductors with an extended impurity ..... 110
10.3 Semiclassical approach to the
Bogoliubov-de Gennes eigenproblem ..... 113
10.4 Averaging over phase fluctuations ..... 117
10.4.1 Gaussian model for phase fluctuations ..... 118
10.4.2 Approximate evaluation of trajectory integrals ..... 118
10.4.3 Linewidth near the superconducting phase boundary ..... 120
10.5 Numerical estimate of the linewidth; Concluding remarks ..... 122

## Appendix

A Application of boundary integral technique to the Helmholtz and Bogoliubov-De Gennes wave equations ..... 124
A. 1 Review of elementary ingredients ..... 125
A. 2 Single- and double-layer parametrizations of wave functions; Jump con- ditions ..... 128
A. 3 Parametrizations for scalar Helmholtz wave functions ..... 129
A.3.1 Jump conditions for one-component Helmholtz wave functions ..... 133
A.3.2 Parametrizations for two-component Bogoliubov-De Gennes wave functions ..... 134
A.3.3 Jump conditions for two-component Bogoliubov-De Gennes wave functions ..... 136
B (d-1)-dimensional Fourier transforms ..... 138
C Propagation outside the normal region and nonconvex shapes: can- cellations ..... 140
References ..... 144
Vita ..... 156

## List of Figures

3.1 Two-dimensional example of an Andreev billiard. ..... 11
4.1 Typical term in the MSE for the Green function. ..... 28
5.1 Geometry for Andreev reflection from a planar surface. ..... 36
5.2 The analogy between Andreev reflection and optical refraction ..... 37
5.3 Typical closed classical trajectory in an Andreev billiard. ..... 39
5.4 Approximating the surface by its tangent plane ..... 41
6.1 Geometry for $\boldsymbol{x}$ integration in Eq. (6.10) ..... 62
6.2 Geometry for fluctuations of the reflection points ..... 64
6.3 Jacobian of the transformation from boundary points to scattering parameters ..... 67
6.4 Example of a billiard that is not circular but has degenerate stationary chords. ..... 73
6.5 A closed periodic orbit with imperfect retro-reflection ..... 75
6.6 More realistic closed periodic orbit with imperfect retro-reflection ..... 77
6.7 Density of states oscillations for a circular Andreev billiard ..... 79
6.8 Density of states oscillations for a circular Andreev Billiard ..... 87
10.1 Schematic illustration of a classical trajectory for a quasiparticle scat- tering from an extended impurity. ..... 114
10.2 Spectral linewidth versus (scaled) temperature. ..... 121
C. 1 Example of nonconvex inside and outside regions of a billiard. ..... 141

## Chapter 1

## Introduction

The purpose of this thesis is to investigate the quantal dynamics of elementary electron and hole quasiparticle excitations in inhomogeneous superconductors. We shall be concerned both with situations in which the inhomogeneity lies in the superconducting pair-potential and when it lies in the single-particle potential. In the former case, we shall explore the setting of Andreev billiards. In the latter case we shall focus on the setting, relevant to high-temperature superconductivity, in which there is an anisotropy in the pair-potential. To date, the main motivations for studying inhomogeneous superconductivity have arisen to understand phenomena such as vortex formation and motion, electromagnetic response, response to an impurity, and pairpotential fluctuation effects near the superconducting transition temperature. Our main interest is to study mesoscopic phenomena in inhomogeneous superconductivity, in which the quantum interference effects on the the quasiparticle mechanics plays a major role, depending on the lengthscale characterizing the inhomogeneity.

By mesoscopic phenomena, we mean phenomena that are quantal (rather than classical) in origin and reveal themselves as the system size is reduced [1]. Typical theoretical methods conventionally employed to study mesoscopic phenomena include replica and supersymmetry techniques invoked to study disordered conductors [2], and random matrix and short-wave asymptotic (i.e. semiclassical) methods
from the field of Quantum Chaos [3]. From the point of view of Quantum Chaos, which is predominantly concerned with the question of how classical chaos shows up at the quantum level, mesoscopic systems are a natural experimental framework. Systems conventionally studied in quantum chaos include Schrödinger billiards (i.e. single quantal particles confined to regions of arbitrary shapes), because they are minimal systems showing the necessary complexity in their classical dynamics. It is typically assumed that the billiard size is much larger than the wavelength of the contained particle, and the central theme then is the connection between the density of energy eigenstates and the classical dynamics of the billiard. The semiclassical density of states of a Schrödinger billiard is separated into a smooth part (the Thomas-Fermi or Weyl part), which depends on bulk properties of the billiard, such as volume, surface area etc., and an oscillatory part, which depends on the closed classical periodic orbits. The smooth part can be considered as the remnant of the macroscopic density of states, whereas the oscillatory part can be regarded as the leading mesoscopic correction due to quantum interference effects.

The mesoscopic physics of inhomogeneous superconductivity is different from its conventional counterpart(i.e. mesoscopics of non-superconducting systems) in that there are two energy scales (viz. the Fermi energy $k_{\mathrm{F}}^{2}$ and the superconducting pairpotential $\Delta$, quasiparticle excitation energy being assumed to be of order of $\Delta$ ). The strong separation of these energy scales (viz. $\Delta \ll k_{\mathrm{F}}^{2}$ ) has long been appreciated, and has justified the quasiclassical scheme first introduced by Andreev [4]. We shall, in chapters 1-7, study a superconducting billiard system which we shall call an Andreev billiard, and show that owing to the additional energy scale there are two layers of density of states oscillations. The coarser one can be described by the quasiclassical approximation. The finer one, however, is inaccessible via the quasiclassical approach and its semiclassical description is one of the main results of this work.

This thesis is organized as follows. Following the introduction to Andreev billiards
(Chapter 2), we define Andreev billiards and present the corresponding BogoliubovDe Gennes (BDG) eigenproblem (in Chapter 3). In this chapter we also introduce the Green function for the BDG eigenproblem and provide its connection to the density of states, and review the standard quasiclassical approach to the BDG eigenproblem (due to Andreev). In Chapter 4 we formulate the computation of the Green function in terms of an expansion in which the basic processes are reflection from and transmission through the interface separating the normal ( N ) and superconducting $(\mathrm{S})$ regions. We then make a physically-motivated reorganization of this expansion, which will later allow us to integrate out states in the superconducting region, thus obtaining a description solely in terms of states within the billiard. (Chapter 4 constitutes the technical basis of the first part of this thesis; however, it may safely be omitted by readers wishing to focus on results rather than methods.) After this exact reformulation we proceed, in Chapter 5, to integrate out the states in the superconducting region within an approximation scheme valid for short quasiparticle wavelenghts. In this way, we obtain an effective Multiple Reflection (rather than Scattering) Expansion. We continue, in Chapter 6, by following certain approximation strategies that allow us to compute the DOS at various levels of energy-resolution, in each case obtaining the corresponding trace formula. This section is the heart of this work. In Chapter 7 we make some concluding remarks and hint at some possible applications of the ideas we have presented.

We then turn from Andreev billiards and s-wave superconductivity to the subject of d-wave superconductors. After a brief introduction to d-wave superconductors in Chapter 8, we move on to explore the properties of low-energy electron-hole quasiparticle excitattions near extended scatterers in d-wave superconductors (Chapter 9). We then discuss, in Chapter, possibility of using the low-energy density of states as a probe of d-wave pairing correlations the pseudogap regime.

We have relegated to appendices some background material, including a derivation
of the Bogoliubov-De Gennes wave equation and the rudiments of boundary integral methods, as well as some technical and parenthetical passages.

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## Chapter 2

## Introduction to Andreev Billiards

The purpose of this work is to consider the quantal dynamics of elementary electron and hole quasiparticle excitations existing within and in the vicinity of a normalstate region of matter that is completely surrounded by an essentially infinite region of conventional superconductor. The entire system-normal-state region and superconducting surround - may be envisaged as three-dimensional, although the approach that we shall be developing is applicable in any number of dimensions. Owing to the inability of the surrounding superconductor to support propagating quasiparticle excitations at sufficiently low energies, electron and hole quasiparticle excitations at such energies are bound to the normal-state region and its vicinity, and it is on the properties of such bound states that we shall be focusing our attention.

The process responsible for the confinement of these excitations to the normalstate region and its vicinity is Andreev reflection [4] from the surrounding superconducting condensate; we shall therefore refer to such structures, near to which quasiparticles are confined, as Andreev billiards. Andreev billiards were introduced and certain simple aspects of their classical dynamics were discussed in Ref. [5]. A very brief account of the approach and results contained in the present thesis were reported in Refs. [6, 7]. Certain quantum-mechanical properties of Andreev billiards were studied in Refs. [8].

As we explore the quantal dynamics of quasiparticle excitations of Andreev billiards, our primary focus will be on the relationship between the quasiparticle energy eigenvalue spectrum and the geometrical shape of the normal-state region, i.e., the question of spectral geometry in this novel setting of excitations confined by a superconducting pair-potential. In the setting of conventional billiard systems [9] confinement is, by contrast, accomplished by an infinite (or occasionally finite) single-particle potential [20]. As mentioned above, we shall primarily be concerned with confined quasiparticle states, and therefore in energy eigenvalues lying within the quasiparticle gap of the surrounding superconductor (although our approach is also suited to the study of scattering states). The strategy that we shall develop is inspired by the beautiful work of Balian and Bloch, in which, inter alia, the eigenvalue spectrum of the Laplace operator was investigated for generically-shaped spatial regions and various types of boundary conditions [10-12]. The central theme of the work of Balian and Bloch is the relationship between the boundary shape, the type of boundary conditions, and the the eigenvalue spectrum. We shall refer to Refs. [10-12] respectively as BB-I, BB-II, and BB-III.

As it is so central to the properties of Andreev billiards, let us pause to review the core qualitative features of the Andreev reflection process: to a high degree of accuracy it (i) interconverts electron and hole excitations; and (ii) reverses the velocity of excitations. It is this latter, retro-reflective, character of the Andreev reflection process that endows Andreev billiards with dynamical characteristics quite distinct from those of conventional billiards, in which confinement is caused by specular reflection from a single-particle potential.

We shall take as a model for the normal-state region of an Andreev billiard a Fermi gas, parametrized by the Fermi energy. Thus, in the normal-state region we shall be neglecting the effects of band structure, impurity scattering, and quasiparticle interactions. We shall account for the superconducting nature of the matter surrounding
the normal-state region by asserting that there is a superconducting pair-potential $\Delta$ that varies discontinuously: inside the normal-state region $\Delta=0$; outside the normalstate region $\Delta$ takes on the constant value $\Delta_{0}(\neq 0)$. We shall refer to the surface on which $\Delta$ changes discontinuously as the shape of the Andreev billiard. Thus, we shall not be working self-consistently, but shall benefit from being in a position to develop an interface-scattering approach to the quasiparticle dynamics, in which we are able to focus on processes occurring at the interface. Hence, we can incorporate in a direct and natural manner the impact of the shape of the billiard (i.e. the shape of the interface) on the spectrum of energy eigenvalues of the confined electron-hole quasiparticles.

We see four principal sources of motivation for the present work. First, Andreev billiards provide a novel setting for the exploration of spectrum-shape relationships, a branch of mathematics with a distinguished history [21, 22]. The novelty is fed in by the Andreev reflection process occuring at the normal-to-superconductor boundary. Second, the usual spectral-geometric scenario [in which deviations of the density of modes from its large-system limit become appreciable as the wavelength become comparable to the characteristic linear size of the system] is not the whole story for the case of Andreev billiards. Instead, owing to the presence of a second, much larger, lengthscale, set by the difference between the momenta of incident electrons and the holes they become upon Andreev reflection (and vice versa). As this momentum difference is small (on the scale of the Fermi momentum), the corresponding lengthscale is much larger that the Fermi wavelength. Through this new lengthscale the eigenvalue spectrum can be sensitive to the shape of the billiard even when the characteristic size of the billiard is much larger than the underlying wave length associated with quasiparticle motion. This relevance of the new lengthscale has long been appreciated, showing up, e.g., in the effectively one-dimensional settings of Tomasch [23] and McMillan-Rowell [24] oscillations in the tunneling density of states above the
superconducting gap, and in de Gennes-Saint-James bound states [25] below the superconducting gap Third, as we shall see when we develop a trace formula for the (oscillatory part of the) density of quasiparticle eigenstates (DOS), there turns out to be a novel and useful separation in the scale of periods of the two dominating classes of (primitive, classical, periodic) trajectories that feature. As a consequence, the DOS will comprise: (i) a relatively smooth contribution due to retracings of geometrical chords across the billiard of stationary lengths, dressed by (ii) a more rapidly varying contribution arising from orbits located near the boundary and involving charge-preserving as well as charge-interconverting reflection processes. Thus, from the oscillatory part of the DOS one can "hear" aspects of the shape of the billiard such as the stationary values of the lengths of the chord ). We are not aware of any other spectral-geometric contexts that feature this type of information. Fourth, the quasiparticle energy eigenvalue spectrum, and its sensitivity to the shape of the billiard, should be experimentally accessible, e.g., via tunneling spectroscopy on hybrid superconducting/normal-state structures. The current state of microfabrication technology makes such experiments realizable [26].

We see the following as the principal results of the present work. First, we provide the machinery for computing the Green function for Andreev billiards of arbitrary shape in terms of a multiple scattering expansion that focuses on the influence of the billiard shape. Second, we implement this machinery to construct two semiclassical schemes (resulting in two semiclassical trace formulas) for computing the oscillatory component of the DOS. One, which we shall refer to as Scheme A, simply amounts to an elaboration (to billiards of arbitrary shape) of Andreev's approximation. Thus, it gives a DOS that takes the form of an integral over the chords of the normal-state region with an appropriate weight function. Hence, it realizes the intuitively natural notion that the chords, being the periodic orbits at the Andreev level, determine the energy eigenvalue spectrum, according to Bohr-Sommerfeld quantization condi-
tions. The other, Scheme B, captures certain physical effects that are inaccessible to Scheme A, such as mesoscale oscillations in the DOS.

## Chapter 3

## Andreev billiards

### 3.1 Idealization of the physical system

The physical system of interest in the present thesis is an Andreev billiard of arbitrary shape, i.e., a normal-state region embedded inside an infinite superconducting region, as depicted in Fig. 3.1. Following Gor'kov's mean-field approach to superconductivity [27, 28], we describe the system by the (variable particle-number) Hamiltonian $H$, given by

$$
\begin{align*}
H=\sum_{\alpha= \pm} \int d^{d} x \psi_{\alpha}^{\dagger}(\mathbf{x}) & \left\{-\frac{\hbar^{2}}{2 m} \nabla^{2}-\mu\right\} \psi_{\alpha}(\mathbf{x}) \\
& +\int d^{d} x\left\{\Delta(\mathbf{x})^{*} \psi_{+}(\mathbf{x}) \psi_{-}(\mathbf{x})+\Delta(\mathbf{x}) \psi_{+}^{\dagger}(\mathbf{x}) \psi_{-}^{\dagger}(\mathbf{x})\right\} \tag{3.1}
\end{align*}
$$

Here, $m$ is the effective electron mass, $\mu$ is the electron chemical potential (which we take to be uniform throughout the system), $\Delta(\mathbf{x})$ is a given superconducting pairpotential which characterizes the superconducting condensate, $\psi_{\alpha}^{\dagger}(\mathbf{x})$ and $\psi_{\alpha}(\mathbf{x})$ are creation and annihilation field operators for electron quasiparticles having position $\mathbf{x}$ and spin-projection $\alpha= \pm$, and $d$ is the dimension of space (typically two or three). We shall not go beyond the picture of quasiparticle excitations propagating in the presence of a superconducting condensate implied by this description. Within the Gor'kov description of the consequences of the electron-electron interaction, any


Figure 3.1: Two-dimensional example of an Andreev billiard, showing a normal region (N) surrounded by a superconducting region (R). In this example the billiard is convex.
(Heisenberg-representation) excited state that is arrived at by the addition of a single spin-up electron quasiparticle to the (Heisenberg-representation) ground state $\left|\Phi_{0}\right\rangle$ evolves into a coherent superposition of such a state and a state arrived at by the removal of a spin-down electron quasiparticle from the ground state[29], the Hamiltonian (3.1) maintaining the system in this sector of Fock space. Thus, it is adequate to address states of the form

$$
\begin{equation*}
\left|\Phi_{1}\right\rangle \equiv \int d^{d} x\left(u(\mathbf{x}) \psi_{\uparrow}^{\dagger}(\mathbf{x})+v(\mathbf{x}) \psi_{\downarrow}(\mathbf{x})\right)\left|\Phi_{0}\right\rangle \tag{3.2}
\end{equation*}
$$

which are described by the two-component complex-valued amplitudes $(u(\mathbf{x}), v(\mathbf{x}))$, i.e., the family of one-quasiparticle excited states.

To derive the equation of motion for these one-quasiparticle states (i.e. the socalled time-dependent BDG equation [30]), we follow Andreev [4] and suppose that the system is in some (Heisenberg-representation) one-quasiparticle excited state $\left|\Phi_{1}\right\rangle$. Then the amplitude $u(\mathbf{x}, t)$ for finding the state at time $t$ to be the ground state with an up-spin electron quasiparticle added at position $\mathbf{x}$ (i.e. the Heisenberg-
representation state $\left.\psi_{+}^{\dagger}(\mathbf{x}, t)\left|\Phi_{0}\right\rangle\right)$ is given by

$$
\begin{equation*}
\left\langle\Phi_{0}\right| \psi_{+}(\mathbf{x}, t)\left|\Phi_{1}\right\rangle . \tag{3.3}
\end{equation*}
$$

Similarly, the amplitude $v(\mathbf{x}, t)$ for finding the state at time $t$ to be the ground state with a down spin electron quasiparticle removed at position $\mathbf{x}$ (i.e. the Heisenbergrepresentation state $\left.\psi_{-}(\mathbf{x}, t)\left|\Phi_{0}\right\rangle\right)$ is given by

$$
\begin{equation*}
\left\langle\Phi_{0}\right| \psi_{-}^{\dagger}(\mathbf{x}, t)\left|\Phi_{1}\right\rangle . \tag{3.4}
\end{equation*}
$$

Here, $\psi_{+}(\mathbf{x}, t) \equiv \mathrm{e}^{i H t / \hbar} \psi_{+}(\mathbf{x}) \mathrm{e}^{-i H t / \hbar}$ and $\psi_{-}^{\dagger}(\mathbf{x}, t) \equiv \mathrm{e}^{i H t / \hbar} \psi_{-}^{\dagger}(\mathbf{x}) \mathrm{e}^{-i H t / \hbar}$ are, respectively, Heisenberg-representation field operators. Thus, the wave functions $u(\mathbf{x}, t)$ and $v(\mathbf{x}, t)$ serve as amplitudes for the present up-spin electron quasiparticle and the absent down-spin electron (i.e. up-spin hole) quasiparticle. Then, by virtue of the Heisenberg equation of motion for the field operators (see, e.g., Ref. [28], Sec. 6), together with the Hamiltonian (3.1), it is a straightforward exercise in computing commutators of field operators to show that the amplitudes $u(\mathbf{x}, t)$ and $v(\mathbf{x}, t)$ evolve according to the appropriate time-dependent Schrödinger equation, i.e., the timedependent BDG equation:

$$
i \hbar \frac{\partial}{\partial t}\binom{u(\mathbf{x}, t)}{v(\mathbf{x}, t)}=\left(\begin{array}{cc}
-\frac{\hbar^{2}}{2 m} \nabla^{2}-\mu & \Delta(\mathbf{x})  \tag{3.5}\\
\Delta^{*}(\mathbf{x}) & -\left(-\frac{\hbar^{2}}{2 m} \nabla^{2}-\mu\right)
\end{array}\right)\binom{u(\mathbf{x}, t)}{v(\mathbf{x}, t)} .
$$

Analysis of this equation via the separation of the time variable, appropriate when there is no external time-dependence, leads to the BDG eigenproblem

$$
\left(\begin{array}{cc}
-\frac{\hbar^{2}}{2 m} \nabla^{2}-\mu & \Delta(\mathbf{x})  \tag{3.6}\\
\Delta^{*}(\mathbf{x}) & -\left(-\frac{\hbar^{2}}{2 m} \nabla^{2}-\mu\right)
\end{array}\right)\binom{u_{n}(\mathbf{x})}{v_{n}(\mathbf{x})}=E_{n}\binom{u_{n}(\mathbf{x})}{v_{n}(\mathbf{x})}
$$

where $n$ is a (collective) index for all quantum numbers and $\left\{E_{n}\right\}$ and $\left\{\left(u_{n}(\mathbf{x}), v_{n}(\mathbf{x})\right)\right\}$ are the corresponding energy eigenvalues and (two-component) eigenfunctions.

In general, $\Delta$ may vary spatially. However, we shall consider the situation in which deep inside the superconductor $\Delta$ goes to a constant value $\Delta_{0}$, whereas throughout the
normal metal it vanishes. In the intermediate region (i.e. within a superconducting coherence length outside of the billiard boundary) $\Delta$ is suppressed to a value smaller than $\Delta_{0}$, and falls to zero as the N region is entered. We shall ignore the effects of this suppression and assume that $\Delta$ varies discontinuously between 0 to $\Delta_{0}$ at a surface, which we refer to as the billiard boundary and denote by $\partial \mathcal{V}$, that divides the system into two homogeneous regions, the billard interior (denoted $\mathcal{V}$ ) and the billiard exterior (denoted $\overline{\mathcal{V}}$ ). For the sake of simplicity, we further assume that there are no metallurgical differences between the normal-state and superconducting regions, inasmuch as the only difference between them is the value of the pair potential (the effective mass, e.g., being common).

To ease the notation we shall adopt units in which $\hbar^{2} / 2 m=1$. To recover results in terms of the original physical units, one multiplies the three variables having the dimensions of energy [viz. $\mu, \Delta(\mathbf{r})$ and $E$ ] by the factor $\left(2 m / \hbar^{2}\right)$.

The Bogoliubov-De Gennes eigenproblem plays the same role for Andreev billiards that the Schrödinger eigenproblem plays for conventional billiards. If a conventional billiard is surrounded by a region in which the single-particle potential is infinite then we refer to the billiard as a hard billiard, and the Schrödinger equation outside the billiard is replaced by the homogeneous Dirichlet (i.e. vanishing) boundary condition on the Schrödinger eigenfunction, this boundary condition leading to the quantization of the eigenvalue spectrum. If, on the other hand, a conventional billiard is surrounded by a region in which the single-particle potential is finite then we refer to the billiard as a soft billiard, and the solution to the Schrödinger equation outside must be matched on to the the solution of the Schrödinger equation inside, this matching leading to the quantization of the eigenvalue spectrum. The Andreev billiard is, therefore, analogous to a soft Schrödinger billiard; its hard limit seems difficult to realize because, at least in known superconductors, the pair potential is far smaller than the chemical potential.

The eigenproblem for Andreev billiards, then, is given by Eq. (3.6) with $\Delta(\mathbf{x})=0$ for $\mathbf{x}$ inside the billiard, and $\Delta(\mathbf{x})=\Delta_{0}$ for $\mathbf{x}$ outside the billiard. Thus, we are faced with the task of addressing the BDG eigenproblem for the case in which the system comprises two spatially homogeneous regions (one normal, one superconducting) that meet at a closed surface. Owing to the spatial homogeneity of these regions, the general solution of the BDG equation can readily be obtained in each region. The quantization of the eigenvalue spectrum results from the matching of the solutions and their normal derivatives across this surface, together with the confinement of the eigenfunctions to the vicinity of the Andeev billiard. The resulting spectrum depends, therefore, on the shape of this surface. Exploring this dependence is the central aim of the present work. It would be straightforward to extend the present framework to handle issues such as Josephson coupling between superconducting regions, scattering from Andreev billiards, etc.

### 3.2 Green function and density of states

The spectrum of eigenvalues of the BDG wave equation $\left\{E_{n}\right\}$ is assembled into the DOS $\rho(E)$, which is defined as

$$
\begin{equation*}
\rho(E) \equiv \sum_{n} \delta\left(E-E_{n}\right) \tag{3.7}
\end{equation*}
$$

As is commonly the case, it is convenient to approach $\rho(E)$ via a Green function for the BDG wave equation, which we now introduce.

### 3.2.1 Green function for the Bogoliubov-De Gennes equation

The ( $2 \times 2$ matrix) Green function $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right)$ for the BDG wave equation is defined by the following matrix partial differential equation:

$$
\left(\begin{array}{cc}
-\nabla^{2}-\mu-z & \Delta(\mathbf{x})  \tag{3.8}\\
\Delta^{*}(\mathbf{x}) & \nabla^{2}+\mu-z
\end{array}\right) \mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right)=\mathbf{I} \delta\left(\mathbf{x}-\mathbf{x}^{\prime}\right),
$$

where $z$ is the complex energy and $\mathbf{I}$ is the $(2 \times 2)$ identity matrix. Under the far-field boundary condition $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right) \rightarrow \mathbf{0}$ as $\left|\mathbf{x}-\mathbf{x}^{\prime}\right| \rightarrow \infty$ the Green function $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right)$ is unique [31]. For the case of the Andreev billiard we have $\Delta(\mathbf{x})=0$ for $\mathbf{x}$ in $\mathcal{V}$ and $\Delta(\mathbf{x})=\Delta_{0}$ for x in $\overline{\mathcal{V}}$.

It is useful to express $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right)$ in terms of the eigenfunctions $\left(u_{n}, v_{n}\right)$ of the BDG Hamiltonian, i.e.,

$$
\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right)=\sum_{n} \frac{1}{E_{n}-z}\left(\begin{array}{ll}
u_{n}(\mathbf{x}) u_{n}^{*}\left(\mathbf{x}^{\prime}\right) & u_{n}(\mathbf{x}) v_{n}^{*}\left(\mathbf{x}^{\prime}\right)  \tag{3.9}\\
v_{n}(\mathbf{x}) u_{n}^{*}\left(\mathbf{x}^{\prime}\right) & v_{n}(\mathbf{x}) v_{n}^{*}\left(\mathbf{x}^{\prime}\right)
\end{array}\right)
$$

In order to see that this form does indeed satisfy Eq. (3.8), one may substitute this expression into Eq. (3.8) and make use of Eq. (3.6) and the completeness of the eigenfunctions, i.e.,

$$
\sum_{n}\left(\begin{array}{ll}
u_{n}(\mathbf{x}) u_{n}^{*}\left(\mathbf{x}^{\prime}\right) & u_{n}(\mathbf{x}) v_{n}^{*}\left(\mathbf{x}^{\prime}\right)  \tag{3.10}\\
v_{n}(\mathbf{x}) u_{n}^{*}\left(\mathbf{x}^{\prime}\right) & v_{n}(\mathbf{x}) v_{n}^{*}\left(\mathbf{x}^{\prime}\right)
\end{array}\right)=\mathbf{I} \delta\left(\mathbf{x}-\mathbf{x}^{\prime}\right)
$$

### 3.2.2 Connection between Bogoliubov-De Gennes Green function and density of states

We now follow the standard practice of expressing a DOS in terms of the corresponding Green function by making use of the identity

$$
\begin{equation*}
\frac{1}{\pi} \operatorname{Im} \lim _{\epsilon \rightarrow+0} \frac{1}{E_{n}-E-i \epsilon}=\delta\left(E_{n}-E\right) \tag{3.11}
\end{equation*}
$$

Together with Eqs. (3.9) and (3.7), this allows us to see that

$$
\begin{aligned}
\rho(E) & =\frac{1}{\pi} \int d^{d} x\left\{\operatorname{Im} \operatorname{Tr} \mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime} ; E+i \epsilon\right)\right\}_{\mathbf{x}^{\prime}=\mathbf{x}} \\
& =\int d^{d} x \sum_{n} \frac{1}{\pi} \operatorname{Im} \frac{\left|u_{n}(\mathbf{x})\right|^{2}+\left|v_{n}(\mathbf{x})\right|^{2}}{E_{n}-E-i \epsilon} \\
& =\sum_{n} \frac{1}{\pi} \operatorname{Im} \frac{1}{E_{n}-E-i \epsilon}=\sum_{n} \delta\left(E-E_{n}\right) .
\end{aligned}
$$

Here, $\operatorname{Tr} \mathbf{G}$ denotes the (matrix) trace over the diagonal components of $\mathbf{G}$.
Equation (3.12) is an expression for $\rho(E)$ in terms of $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime} ; E+i \epsilon\right)$, and any approximation to $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime} ; E+i \epsilon\right)$ thus furnishes an approximation to $\rho(E)$. However, being a sum of delta functions, $\rho(E)$ is not a smooth function and can, therefore, be extremely awkward to approximate. In order to find approximations to $\rho(E)$ it is preferable to seek a continuous function that carries essentially the same information as it. One candidate is the smoothed DOS $\rho_{\gamma}(E)$, defined via

$$
\begin{equation*}
\rho_{\gamma}(E) \equiv \int_{-\infty}^{\infty} d E^{\prime} f\left(E-E^{\prime} ; \gamma\right) \rho\left(E^{\prime}\right) \tag{3.12}
\end{equation*}
$$

where $f\left(E-E^{\prime} ; \gamma\right)$ is some smoothing function and $\gamma$ is the (real) smoothing width. We remark, parenthetically, that any DOS derived from experiment will be smoothed to some extent, depending on the resolving power of the apparatus and/or the lifetime of the single-particle excitations.

There are several possible choices for the smoothing fuction $f\left(E-E^{\prime} ; \gamma\right)$, including, e.g., Lorentzian, Gaussian and logarithmic-Gaussian. It is also possible to define a continuous integral transform of $\rho(E)$ that is, itself, a physical property, and for which we can derive some approximation. For example, under the Lambert transform of $\rho(E)$, which exchanges energy for temperature, $\rho(E)$ is transformed into the total (equilibrium internal) energy as a function of temperature [32]. Which smoothing procedure is the best choice depends on the method used to approximate $\rho(E)$. For Green-function-based methods, such as the one we shall adopt, the Lorentzian
smoothing function,

$$
\begin{equation*}
f\left(E-E^{\prime} ; \gamma\right) \equiv \frac{1}{\pi} \frac{\gamma}{\left(E-E^{\prime}\right)^{2}+\gamma^{2}}, \tag{3.13}
\end{equation*}
$$

is the most appropriate, for reasons that should become clear below.
We now give the analogue of Eq. (3.12) for relating the Lorentzian-smoothed DOS $\rho_{\gamma}(E)$ and the Green function at complex energy $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right)$. Following essentially the procedure that lead to Eq. (3.12), we have the following identity for arbitrary (real) $\gamma$ :

$$
\begin{align*}
\rho_{\gamma}(E) & =\left.\frac{1}{\pi} \int d^{d} x \operatorname{Im} \operatorname{Tr} \mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime} ; E+i \gamma\right)\right|_{\mathbf{x}^{\prime}=\mathbf{x}}=\int d^{d} x \sum_{n} \frac{1}{\pi} \operatorname{Im} \frac{\left|u_{n}(\mathbf{x})\right|^{2}+\left|v_{n}(\mathbf{x})\right|^{2}}{E_{n}-E-i \gamma} \\
& =\sum_{n} \frac{1}{\pi} \frac{\gamma}{\left(E-E_{n}\right)^{2}+\gamma^{2}}=\int_{-\infty}^{\infty} d E^{\prime} \frac{1}{\pi} \frac{\gamma}{\left(E-E^{\prime}\right)^{2}+\gamma^{2}} \sum_{n} \delta\left(E^{\prime}-E_{n}\right) .(3.14) \tag{3.14}
\end{align*}
$$

Naturally, in the limit $\gamma \rightarrow 0^{+}, \rho_{\gamma}(E)$ passes to $\rho(E)$.

### 3.2.3 Fundamental Green function for a homogeneous normal region

In this section we derive the fundamental (i.e. homogeneous) normal-state Green function, the first of two Green functions that are central to the construction of the Green function for an Andreev billiard. Along the way, we introduce some convenient notation that we shall subsequently make use of. This fundamental Green function is the translationally-invariant solution $\mathbf{G}^{\mathrm{N}}\left(\mathrm{x}, \mathrm{x}^{\prime} ; z\right)$ of the equation

$$
\begin{align*}
& (\mathbf{H}-z \mathbf{I}) \mathbf{G}_{0}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right)=\delta^{(d)}\left(\mathbf{x}-\mathbf{x}^{\prime}\right) \mathbf{I},  \tag{3.15}\\
& \mathbf{H}=\left(\begin{array}{cc}
\hat{\mathbf{p}}^{2}-\mu & 0 \\
0 & -\hat{\mathbf{p}}^{2}+\mu
\end{array}\right) \equiv\left(-\hat{\mathbf{p}}^{2}+\mu\right) \boldsymbol{\sigma}_{3}, \tag{3.16}
\end{align*}
$$

where $\boldsymbol{\sigma}_{1,2,3}$ are the three Pauli matrices, together with the boundary condition that the Green function vanishes at infinity:

$$
\begin{equation*}
\mathrm{G}^{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right) \rightarrow \mathbf{0} \quad \text { as } \quad\left|\mathbf{x}-\mathbf{x}^{\prime}\right| \rightarrow \infty \tag{3.17}
\end{equation*}
$$

The solution for this Green function is given by the (spatial) matrix element of the operator $(z \mathbf{I}-\mathbf{H})^{-1}$, i.e.,

$$
\begin{align*}
\mathbf{G}^{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right)= & \langle\mathbf{x}| \frac{1}{z \mathbf{I}-\mathbf{H}}\left|\mathbf{x}^{\prime}\right\rangle=\langle\mathbf{x}| \frac{z \mathbf{I}+\left(\hat{\mathbf{p}}^{2}-\mu\right) \boldsymbol{\sigma}_{3}}{z^{2}-\left(\hat{\mathbf{p}}^{2}-\mu\right)^{2}}\left|\mathbf{x}^{\prime}\right\rangle \\
= & \left(z \mathbf{I}-\left(\nabla_{x}^{2}+\mu\right) \boldsymbol{\sigma}_{3}\right) / 2 E \\
& \quad \times\left(\langle\mathbf{x}|\left(z-\hat{\mathbf{p}}^{2}+\mu\right)^{-1}\left|\mathbf{x}^{\prime}\right\rangle+\langle\mathbf{x}|\left(z+\hat{\mathbf{p}}^{2}-\mu\right)^{-1}\left|\mathbf{x}^{\prime}\right\rangle\right), \tag{3.18}
\end{align*}
$$

which is, of course, a matrix in electron/hole space. The two terms inside the parantheses are quite familiar: they are the usual Green functions of the Helmholtz wave equation, and can be evaluated in the standard way. In three dimensions, e.g., one has

$$
\begin{equation*}
\langle\mathbf{x}|\left(z \mp \hat{\mathbf{p}}^{2} \pm \mu\right)^{-1}\left|\mathbf{x}^{\prime}\right\rangle= \pm \frac{1}{4 \pi} \frac{\mathrm{e}^{ \pm i k_{ \pm}\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}, \tag{3.19}
\end{equation*}
$$

The symbols $k_{ \pm}$, which will be used throughout this thesis, denote particle and hole wave numbers, and are given by the expressions

$$
\begin{equation*}
k_{ \pm}=\sqrt{\mu \pm z} \tag{3.20}
\end{equation*}
$$

where it is understood that the roots having positive real parts are the ones that are adopted. Then, for the three-dimensional case, one arrives at the the fundamental Green function for the normal state:

$$
\begin{align*}
\mathbf{G}^{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right) & =\frac{z \mathbf{I}-\left(\nabla^{2}+\mu\right) \boldsymbol{\sigma}_{3}}{2 z} \frac{1}{4 \pi}\left(\frac{\mathrm{e}^{i k_{+}\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}-\frac{\mathrm{e}^{-i k_{-}\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}\right) \\
& =\frac{1}{4 \pi\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}\left(\begin{array}{cc}
\mathrm{e}^{i k_{+}\left|\mathbf{x}-\mathbf{x}^{\prime}\right|} & 0 \\
0 & -\mathrm{e}^{-i k_{-}\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}
\end{array}\right) . \tag{3.21}
\end{align*}
$$

This Green function has the expected form: an outgoing spherical wave in the particle component (i.e. the particle Green function) and an incoming spherical wave in the hole component (i.e. the hole Green function), the latter representing an outgoing hole. Moreover, as we expect in the normal state, there is no mixing between the particle and hole components, the off-diagonal elements being zero. Next, we introduce
some convenient notation for the components of this fundamental Green function, viz.,

$$
\begin{equation*}
g_{ \pm}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \equiv \frac{\mathrm{e}^{ \pm i k_{ \pm}\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|} \tag{3.22}
\end{equation*}
$$

in terms of which $\mathbf{G}^{\mathrm{N}}$ becomes

$$
\mathbf{G}^{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime} ; z\right)=\left(\begin{array}{cc}
g_{+}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) & 0  \tag{3.23}\\
0 & -g_{-}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)
\end{array}\right) .
$$

### 3.2.4 Fundamental Green function for a homogeneous superconducting region

The fundamental Green function for a homogeneous superconducuting region is the translationally invariant solution $\mathbf{G}^{\mathrm{S}}\left(\mathbf{x}, \mathrm{x}^{\prime} ; z\right)$ of the equation:

$$
\begin{equation*}
\left\{z \mathbf{I}+\left(\nabla^{2}+\mu\right) \boldsymbol{\sigma}_{3}+\Delta_{1} \boldsymbol{\sigma}_{1}+\Delta_{2} \boldsymbol{\sigma}_{2}\right\} \mathbf{G}^{\mathrm{S}}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)=\mathbf{I} \delta^{(d)}\left(\mathbf{x}-\mathrm{x}^{\prime}\right) \tag{3.24}
\end{equation*}
$$

together with the boundary condition that it vanishes as $\left|x-x^{\prime}\right|$ goes to infinity. Here $\Delta_{1}$ and $\Delta_{2}$ represent the (constant) real and imaginary parts of the complex pair-potential. Note that we shall henceforth take the complex energy $z$ to be the real energy $E$. We shall be concerned with situations in which the quasiparticles are bound to the billiard and shall, threfore, assume that $|E|<\Delta$, where the magnitude $\Delta$ of the pair potential obeys $\Delta^{2} \equiv \Delta_{1}^{2}+\Delta_{2}^{2}$. In the present work, it is convenient to choose a specific gauge, which we do by setting $\Delta_{1}=0$ and $\Delta_{2}=\Delta$. However, for extensions of the present work to settings, such as SNS junctions, in which there are physical implications of phase differences it is necessary to ksow the Green function for arbitrary gauges, and it is therefore for this case that we provide the Green function.

To obtain the Green function one first formally inverts Eq. (3.24) to obtain

$$
\begin{equation*}
\mathbf{G}^{\mathbf{S}}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)=\left\{E \mathbf{I}+\left(\nabla^{2}+\mu\right) \boldsymbol{\sigma}_{3}+\Delta_{1} \boldsymbol{\sigma}_{1}+\Delta_{2} \boldsymbol{\sigma}_{2}\right\}^{-1} \delta^{(d)}\left(\mathbf{x}-\mathbf{x}^{\prime}\right) \tag{3.25}
\end{equation*}
$$

By manipulating this equation so as to separate the matrix and partial-differential
aspects of the inversion operation one then arrives at

$$
\begin{align*}
\mathbf{G}^{\mathrm{S}}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)= & \left\{E \mathbf{I}+\left(\nabla^{2}+\mu\right) \boldsymbol{\sigma}_{3}+\Delta_{1} \boldsymbol{\sigma}_{1}+\Delta_{2} \boldsymbol{\sigma}_{2}\right\}^{-1} \delta^{3}\left(\mathbf{x}-\mathbf{x}^{\prime}\right) \\
= & \left\{E \mathbf{I}-\left(\nabla^{2}+\mu\right) \boldsymbol{\sigma}_{3}-\Delta_{1} \boldsymbol{\sigma}_{1}-\Delta_{2} \boldsymbol{\sigma}_{2}\right\} \\
& \times\left\{E^{2}-\Delta^{2}-\left(\nabla^{2}+\mu\right)^{2}\right\}^{-1} \delta^{(d)}\left(\mathbf{x}-\mathbf{x}^{\prime}\right) \\
= & \left\{E \mathbf{I}-\left(\nabla^{2}+\mu\right) \boldsymbol{\sigma}_{3}-\Delta_{1} \boldsymbol{\sigma}_{1}-\Delta_{2} \boldsymbol{\sigma}_{2}\right\} \frac{1}{2 i \sqrt{\Delta^{2}-E^{2}}} \\
& \times\left\{\left(\nabla^{2}+\left(k_{+}^{\mathrm{S}}\right)^{2}\right)^{-1}-\left(\nabla^{2}+\left(k_{-}^{\mathrm{S}}\right)^{2}\right)^{-1}\right\} \delta^{(d)}\left(\mathbf{x}-\mathbf{x}^{\prime}\right) \tag{3.26}
\end{align*}
$$

where we identify the wave vectors $k_{+}^{\mathrm{S}}$ and $k_{-}^{\mathrm{S}}$ for the electron-like and hole-like components, respectively, as follows:

$$
\begin{equation*}
k_{ \pm}^{S} \equiv \sqrt{\mu \pm i \sqrt{\Delta^{2}-E^{2}}}, \quad \operatorname{Re} k_{ \pm}^{S}>0 \tag{3.27}
\end{equation*}
$$

As the partial differential operators $\left(\nabla^{2}+\left(k_{ \pm}^{S}\right)^{2}\right)$ are partial differential operators of the Helmholtz wave equation for wave vectors of length $k_{ \pm}^{\mathrm{S}}$, their inversion is well known,

$$
\begin{equation*}
g_{ \pm}^{\mathrm{S}}\left(\mathbf{x}-\mathbf{x}^{\prime}\right) \equiv-\left(\nabla^{2}+\left(k_{ \pm}^{\mathrm{S}}\right)^{2}\right)^{-1} \delta^{3}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)=\frac{\mathrm{e}^{ \pm i k_{ \pm}^{\mathrm{S}}\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{4 \pi\left|\mathbf{x}-\mathbf{x}^{\prime}\right|} \tag{3.28}
\end{equation*}
$$

where the final form on the right hand side holds for the three-dimensional case, and the minus sign in the exponent for holes ensures the proper decay at large distances. Upon using Eq. (3.28), $\mathbf{G}^{\text {S }}$ becomes

$$
\begin{align*}
\mathbf{G}^{\mathrm{S}}= & \left\{E \mathbf{I}-\left(\nabla^{2}+\mu\right) \boldsymbol{\sigma}_{3}-\Delta_{1} \boldsymbol{\sigma}_{1}-\Delta_{2} \boldsymbol{\sigma}_{2}\right\} \frac{1}{2 i \sqrt{\Delta^{2}-E^{2}}}\left(g_{+}^{\mathrm{S}}-g_{-}^{\mathrm{S}}\right)  \tag{3.29}\\
= & \frac{1}{2}\left\{\frac{E}{i \sqrt{\Delta^{2}-E^{2}}} \mathbf{I}-\frac{\Delta_{1}}{i \sqrt{\Delta^{2}-E^{2}}} \boldsymbol{\sigma}_{1}-\frac{\Delta_{2}}{i \sqrt{\Delta^{2}-E^{2}}} \boldsymbol{\sigma}_{2}\right\}\left(g_{+}^{\mathrm{S}}-g_{-}^{\mathrm{S}}\right) \\
& +\frac{1}{2} \boldsymbol{\sigma}_{3}\left(g_{+}^{\mathrm{S}}+g_{-}^{\mathrm{S}}\right) . \tag{3.30}
\end{align*}
$$

To get to the second line, the action of $\left(\nabla^{2}+\mu\right)$ on $g_{ \pm}^{\mathrm{S}}$ is calculated with the help of Eq. (3.28).

### 3.3 Andreev's quasiclassical aproximation scheme

We now review the conventional approximation scheme for studying N-S hybrid systems, such as Andreev billiards. This approximation scheme was put forth by Andreev [4], and we shall refer to it as Andreev's approximation. It takes advantage of the fact that for typical N-S systems the dimensionless parameter $\Delta / \mu$ is much less than unity. The scheme consists of the separation of rapid and slow oscillations in the wavefunction. It can be motivated in the following way: consider an arbitrary N-S structure (i.e. arbitrary $\Delta(\mathbf{x}) \ll \mu$ ). Suppose that the electronic properties of the system are probed during a short time $\Delta t$, so that there is insufficient energy resolution around $E=0$ to resolve $\Delta(\mathbf{x})$. Then no measurement obtained via this probe is capable of distinguishing between the original system and a system described by the same BDG equation but with $\Delta(\mathbf{x})$ set to zero. An energy resolution of $\Delta E$ around $E=0$ corresponds to a momentum resolution of $\Delta p=m \Delta E / k_{\mathrm{F}}$ which, via the Heisenberg uncertainty relation, corresponds to a spatial resolution of $\Delta x=\hbar^{2} k_{\mathrm{F}} / m \Delta E$. Therefore, in order to resolve any effects of the pair-potential on the electronic states, the system must be probed on lengthscales larger than $\xi \equiv \hbar^{2} k_{\mathrm{F}} / m \Delta$, i.e., the superconducting coherence length. Thus, the single-particle wavefunctions must have the form of plane-waves at the Fermi momentum with an envelope that varies on the lengthscale $\xi$, i.e.,

$$
\begin{equation*}
\binom{u(\mathbf{x})}{v(\mathbf{x})}=\mathrm{e}^{i k_{F} \mathbf{n} \cdot \mathbf{x}}\binom{\bar{u}(\mathbf{x})}{\bar{v}(\mathbf{x})}, \tag{3.31}
\end{equation*}
$$

where the unit vector $\mathbf{n}$ defines the orientation of the wavevector of the plane wave and $\bar{u}$ and $\bar{v}$ are the slowly varying envelope amplitudes. By substituting this form into the BDG equation (3.6) one obtains

$$
\left(\begin{array}{cc}
-\left(\nabla^{2}+2 i k_{F} \mathbf{n} \cdot \boldsymbol{\nabla}\right) & \Delta(\mathbf{x})  \tag{3.32}\\
\Delta^{*}(\mathbf{x}) & \left(\nabla^{2}+2 i k_{F} \mathbf{n} \cdot \boldsymbol{\nabla}\right)
\end{array}\right)\binom{\bar{u}(\mathbf{x})}{\bar{v}(\mathbf{x})}=E\binom{\bar{u}(\mathbf{x})}{\bar{v}(\mathbf{x})} .
$$

As $\bar{u}$ and $\bar{v}$ vary on the lengthscale $\xi$, one has that

$$
\begin{equation*}
\frac{\left|\nabla^{2} \bar{u}\right|}{\left|k_{F} \mathbf{n} \cdot \boldsymbol{\nabla} \bar{u}\right|} \sim \mathcal{O}\left(\frac{1}{k_{F} \xi}\right)=\mathcal{O}\left(\frac{\Delta}{\mu}\right) \ll 1 . \tag{3.33}
\end{equation*}
$$

Thus, to leading order in $\Delta / \mu$ it is permissible to ignore the $\nabla^{2}$ term in Eq. (3.32), a singular approximation because it involves changing the order of the system of partial differential equations. This approximation to the BDG Hamiltonian is Andreev's approximation, and leads to the Andreev eigenproblem

$$
\left(\begin{array}{cc}
-2 i k_{F} \mathbf{n} \cdot \boldsymbol{\nabla} & \Delta(\mathbf{x})  \tag{3.34}\\
\Delta^{*}(\mathbf{x}) & 2 i k_{F} \mathbf{n} \cdot \boldsymbol{\nabla}
\end{array}\right)\binom{\bar{u}(\mathbf{x})}{\bar{v}(\mathbf{x})}=E\binom{\bar{u}(\mathbf{x})}{\bar{v}(\mathbf{x})} .
$$

Let us now apply Eq. (3.34) to Andreev billiards. First, it is useful to express the variable $\mathbf{x}$ in terms of $\mathbf{b}$ (i.e. the impact parameter or, equivalently, the transverse parameter), and $s$ (i.e. the longitudinal parameter) such that

$$
\begin{equation*}
\mathbf{x}=\mathbf{b}+\mathbf{n} s \tag{3.35}
\end{equation*}
$$

Notice that $\mathbf{b}$ represents the transverse degree(s) of freedom whereas $s$ represents the longitudinal degree of freedom of the excitation. As, in Eq. (3.34), there is no differential operator acting on the variable $\mathbf{b}$, the wavefunctions take the form

$$
\begin{equation*}
\binom{\bar{u}(\mathbf{b}, s)}{\bar{v}(\mathbf{b}, s)}=\delta\left(\mathbf{b}-\mathbf{b}_{0}\right)\binom{u\left(s ; \mathbf{b}_{0}\right)}{v\left(s ; \mathbf{b}_{0}\right)}, \tag{3.36}
\end{equation*}
$$

where $\mathbf{b}_{0}$ can be interpreted as the transverse quantum number. By substituting this form into Eq. (3.34) one obtains

$$
\left(\begin{array}{cc}
-2 i k_{F} \partial / \partial s & \Delta\left(\mathbf{b}_{0}, s\right)  \tag{3.37}\\
\Delta^{*}\left(\mathbf{b}_{0}, s\right) & 2 i k_{F} \partial / \partial s
\end{array}\right)\binom{u\left(s ; \mathbf{b}_{0}\right)}{v\left(s ; \mathbf{b}_{0}\right)}=E\binom{u\left(s ; \mathbf{b}_{0}\right)}{v\left(s ; \mathbf{b}_{0}\right)} .
$$

Thus, one has reduced the partial differential eigenvalue equation (3.6) to a family of approximate ordinary differential eigenvalue equations parametrized by $\left(\mathbf{n}, \mathbf{b}_{0}\right)$.

Now, for Andreev billiards $\Delta$ is real and piecewise constant. In this case, the solution of Eq. (3.34) proceeds as follows. The parameters ( $\mathbf{n}, \mathbf{b}_{0}$ ) fix a line, which determines $\Delta\left(\mathbf{b}_{0}, s\right)$, in the sense that $\Delta\left(\mathbf{b}_{0}, s\right)=\Delta\left(\mathbf{b}_{0}+\mathbf{n} s\right)$. As $\Delta\left(\mathbf{b}_{0}, s\right)$ is piecewise constant, the task of solving Eq. (3.37) is straightforward. The quantization condition depends on the chord length $\ell\left(\mathbf{n}, \mathbf{b}_{0}\right)$ [i.e. the length of the part of the line (specified by $\left(\mathbf{n}, \mathbf{b}_{0}\right)$ ) lying inside N$]$ :

$$
\begin{equation*}
E \ell\left(\mathbf{n}, \mathbf{b}_{0}\right) / k_{\mathrm{F}}-2 \varphi=2 \pi m \tag{3.38}
\end{equation*}
$$

where $m=0, \pm 1, \pm 2, \ldots$ and $\varphi \equiv \cos ^{-1}(E / \Delta)$. Thus, for each chord there is a ladder of energy eigenvalues. In order to obtain the DOS one must first obtain the DOS for a single chord, and then sum over all chords. However, it is not a priori completely clear what weight should be assigned to each chord in performing the continuous summation. In fact, as we shall see in Sec. 6.1, at the Andreev level of approximation the DOS for a (convex, $d$-dimensional) Andreev billiard is given by

$$
\begin{equation*}
\rho(E) \approx \sum_{m=-\infty}^{\infty} \frac{k_{\mathrm{F}}^{d-2}}{2(2 \pi)^{d-1}} \int d \mathbf{n} d \mathbf{b} \ell\left(\mathbf{n}, \mathbf{b}_{0}\right) \delta\left(\frac{E}{k_{\mathrm{F}}} \ell\left(\mathbf{n}, \mathbf{b}_{0}\right)-2 \varphi-2 \pi m\right), \tag{3.39}
\end{equation*}
$$

the two ( $d-1$ )-dimensional integrations, one over $\mathbf{n}$ and one over $\mathbf{b}$, implement the summation over chords.

## Chapter 4

## Multiple scattering expansion

We now construct a Multiple Scattering Expansion capable of yielding the Green function $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$ associated with the BDG equation (3.8), in settings in which $\Delta(\mathbf{x})$ is piecewise constant (i.e. takes on certain constant values in various regions of space that are delineated by $d$ - 1 -dimensional surfaces). Although the construction is applicable to a wider range of settings, we shall have in mind the application to an Andreev billiard, for which $\Delta(\mathbf{x})$ vanishes inside the billiard and has a constant nonzero value $\Delta_{0}$ outside it. The spirit of our approach very much parallels that of BB-I, although there there are significant differences arising from (i) the form of the eigenproblem (BDG rather than Laplace), and (ii) our need to employ matching (rather than boundary) conditions.

In this section we set up the general formalism for the exact BDG Green function for the case in which the system is divided, via a change in the pair potential, into at least two distinct regions. As we shall see, the exact Green function can be expressed as a sum, generated by an interation scheme, over all possible scatterings from the boundary dividing these regions.

Readers not familiar with the elements of potential theory that we shall be using (which are sometimes referred to as boundary integral techniques) may wish to pause to read App. A, in which we give a self-contained introduction to this subject
and develop the elaborations necessary for application to the BDG eigenproblem. Specifically, we shall need to handle two-component wave functions in the context of matching (rather than boundary) conditions.

### 4.1 Matching conditions and boundary integral equations for the Bogoliubov-De Gennes Green function

We now introduce a convenient parametrization of the Green function $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$. We do this by decomposing $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$ into a particular integral and a complementary function. The particular integral, which yields the delta function under the action of the BDG operator, is built from the fundamental Green functions $G^{N}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)$ and $\mathbf{G}^{\mathrm{S}}\left(\mathrm{x}-\mathrm{x}^{\prime}\right)$. The complementary function, which obeys the BDG wave equation [i.e. the homogeneous version of Eq. (3.8], is specified in terms of as-yet undetermined single and double layers $\boldsymbol{\mu}^{\mathrm{ii}}\left(\boldsymbol{\alpha}, \mathrm{x}^{\prime}\right)$, $\boldsymbol{\nu}^{\mathrm{oi}}\left(\boldsymbol{\alpha}, \mathrm{x}^{\prime}\right), \boldsymbol{\mu}^{\text {io }}\left(\boldsymbol{\alpha}, \mathrm{x}^{\prime}\right)$, and $\boldsymbol{\nu}^{\mathrm{oo}}\left(\boldsymbol{\alpha}, \mathrm{x}^{\prime}\right)$. Thus we write
$\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)= \begin{cases}\mathbf{G}^{\mathrm{ii}} \equiv \mathbf{G}^{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\mathbf{x}, \boldsymbol{\alpha}) \boldsymbol{\mu}^{\mathrm{ii}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right), & \text { if } \mathbf{x} \in \mathcal{V} \text { and } \mathbf{x}^{\prime} \in \mathcal{V} ; \\ \mathbf{G}^{\mathrm{oi}} \equiv \quad \int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{G}^{\mathrm{S}}(\mathbf{x}, \boldsymbol{\alpha}) \boldsymbol{\nu}^{\mathrm{oi}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right), & \text { if } \mathbf{x} \in \overline{\mathcal{V}} \text { and } \mathbf{x}^{\prime} \in \mathcal{V} ; \\ \mathbf{G}^{\mathrm{io}} \equiv \quad \int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\mathbf{x}, \boldsymbol{\alpha}) \boldsymbol{\mu}^{\mathrm{io}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right), & \text { if } \mathbf{x} \in \mathcal{V} \text { and } \mathbf{x}^{\prime} \in \overline{\mathcal{V}} ; \\ \mathbf{G}^{\mathrm{oo}} \equiv \mathbf{G}^{\mathrm{S}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{G}^{\mathrm{S}}(\mathbf{x}, \boldsymbol{\alpha}) \boldsymbol{\nu}^{\mathrm{oo}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right), & \text { if } \mathbf{x} \in \overline{\mathcal{V}} \text { and } \mathbf{x}^{\prime} \in \overline{\mathcal{V}} .\end{cases}$

We are employing the notation $\partial_{\alpha} f(\boldsymbol{\alpha})$ to indicate the value of the component of the gradient of $f(\mathbf{x})$ with respect to $\mathbf{x}$, evaluated at the surface point $\boldsymbol{\alpha}$, directed along the inward normal direction at $\boldsymbol{\alpha}$. Similarly, $d \sigma_{\alpha}$ indicates the (scalar) surface element at the point $\boldsymbol{\alpha}$. We shall find it useful to decorate $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$ with labels (such as, e.g., ii and io) according to the region of space in which the variables $\mathbf{x}$ and $\mathbf{x}^{\prime}$ lie.

For example, if $\mathbf{x} \in \mathcal{V}$ and $\mathbf{x}^{\prime} \in \overline{\mathcal{V}}$ then we write $\mathbf{G}^{\mathrm{io}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$ for $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$, conveying the idea that x lies inside $\mathcal{V}$ whereas $\mathrm{x}^{\prime}$ lies outside $\mathcal{V}$.

Next, we focus on the surface $\partial \mathcal{V}$ across which the pair-potential is discontinuous. From the Green function equation (3.8) it is evident that both $\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$ and its normal derivative $\mathbf{n} \cdot \nabla_{x} \mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$ are continuous as $\mathbf{x}$ varies across $\partial \mathcal{V}$. What this means is that

$$
\begin{align*}
\lim _{\mathbf{x} \in \mathcal{V} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) & =\lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) ;  \tag{4.2}\\
\lim _{\mathbf{x} \in \boldsymbol{\mathcal { V }} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \mathbf{n}_{\beta} \nabla_{x} \mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) & =\lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \mathbf{n}_{\beta} \nabla_{x} \mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) . \tag{4.3}
\end{align*}
$$

We apply this pair of matching conditions across $\partial \mathcal{V}$ to the parametrizations (4.1), once for $\mathrm{x}^{\prime} \in \mathcal{V}$ and once for $\mathrm{x}^{\prime} \in \overline{\mathcal{V}}$. To evaluate the necessary limits we make use of the using the continuity conditions (A.36) and (A.35), as well as the jump conditions (A.34) and (A.37), thus arriving at the following conditions on the single layers $\boldsymbol{\nu}^{\text {oi }}$ and $\boldsymbol{\nu}^{\text {oo }}$ and double layers $\boldsymbol{\mu}^{\text {ii }}$ and $\boldsymbol{\mu}^{\text {io }}$ :

$$
\begin{align*}
& \frac{1}{2} \boldsymbol{\sigma}_{3} \boldsymbol{\mu}^{\mathrm{ii}}\left(\boldsymbol{\beta}, \mathrm{x}^{\prime}\right)=-\mathrm{G}^{\mathrm{N}}\left(\boldsymbol{\beta}, \mathrm{x}^{\prime}\right)-\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\boldsymbol{\beta}, \boldsymbol{\alpha}) \boldsymbol{\mu}^{\mathrm{ii}}\left(\boldsymbol{\alpha}, \mathrm{x}^{\prime}\right)  \tag{4.4}\\
& +\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathrm{G}^{\mathrm{S}}(\boldsymbol{\beta}, \boldsymbol{\alpha}) \boldsymbol{\nu}^{\mathrm{oi}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) ; \\
& \frac{1}{2} \boldsymbol{\sigma}_{3} \boldsymbol{\nu}^{\mathrm{oi}}\left(\boldsymbol{\gamma}, \mathrm{x}^{\prime}\right)=\partial_{\gamma} \mathbf{G}^{\mathrm{N}}\left(\boldsymbol{\gamma}, \mathrm{x}^{\prime}\right)+\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\gamma}^{+} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\boldsymbol{\gamma}, \boldsymbol{\alpha}) \boldsymbol{\mu}^{\mathrm{ii}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right)  \tag{4.5}\\
& -\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\gamma} \mathrm{G}^{\mathrm{S}}(\boldsymbol{\gamma}, \boldsymbol{\alpha}) \boldsymbol{\nu}^{\mathrm{oi}}\left(\boldsymbol{\alpha}, \mathrm{x}^{\prime}\right) ; \\
& \frac{1}{2} \boldsymbol{\sigma}_{3} \boldsymbol{\mu}^{\mathrm{io}}\left(\boldsymbol{\beta}, \mathrm{x}^{\prime}\right)=\mathrm{G}^{\mathrm{S}}\left(\boldsymbol{\gamma}, \mathrm{x}^{\prime}\right)+\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathrm{G}^{\mathrm{S}}(\boldsymbol{\beta}, \boldsymbol{\alpha}) \boldsymbol{\nu}^{\mathrm{oo}}\left(\boldsymbol{\alpha}, \mathrm{x}^{\prime}\right)  \tag{4.6}\\
& -\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\boldsymbol{\beta}, \boldsymbol{\alpha}) \boldsymbol{\mu}^{\mathrm{io}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) ; \\
& \frac{1}{2} \boldsymbol{\sigma}_{3} \boldsymbol{\nu}^{\mathrm{oo}}\left(\boldsymbol{\gamma}, \mathrm{x}^{\prime}\right)=-\partial_{\gamma} \mathrm{G}^{\mathrm{S}}\left(\boldsymbol{\gamma}, \mathrm{x}^{\prime}\right)-\int_{\partial \nu} d \sigma_{\alpha} \partial_{\gamma} \mathrm{G}^{\mathrm{S}}(\boldsymbol{\gamma}, \boldsymbol{\alpha}) \boldsymbol{\nu}^{\mathrm{oo}}\left(\boldsymbol{\alpha}, \mathrm{x}^{\prime}\right)  \tag{4.7}\\
& +\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\gamma}^{+} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\boldsymbol{\gamma}, \boldsymbol{\alpha}) \boldsymbol{\mu}^{\mathrm{io}}\left(\boldsymbol{\alpha}, \mathrm{x}^{\prime}\right) .
\end{align*}
$$

These four matching conditions constitute a system of coupled integral equations for the single layers $\boldsymbol{\nu}^{\mathrm{oi}}$ and $\boldsymbol{\nu}^{\mathrm{oo}}$ and the double layers $\boldsymbol{\mu}^{\mathrm{ii}}$ and $\boldsymbol{\mu}^{\mathrm{io}}$. It is convenient to
collect together the single and double layers into a $4 \times 4$ matrix $\mathbb{M}\left(\boldsymbol{\gamma}, \mathbf{x}^{\prime}\right)$, defined via

$$
\mathbb{M}\left(\gamma, \mathrm{x}^{\prime}\right) \equiv\left(\begin{array}{ll}
\boldsymbol{\mu}^{\mathrm{ii}}\left(\boldsymbol{\gamma}, \mathrm{x}^{\prime}\right) & \boldsymbol{\mu}^{\mathrm{io}}\left(\boldsymbol{\gamma}, \mathrm{x}^{\prime}\right)  \tag{4.8}\\
\boldsymbol{\nu}^{\mathrm{oi}}\left(\boldsymbol{\gamma}, \mathrm{x}^{\prime}\right) & \boldsymbol{\nu}^{\mathrm{oo}}\left(\boldsymbol{\gamma}, \mathrm{x}^{\prime}\right)
\end{array}\right) .
$$

In terms of $\mathbb{M}\left(\boldsymbol{\gamma}, \mathbf{x}^{\prime}\right)$ the system becomes

$$
\begin{equation*}
\mathbb{M}\left(\boldsymbol{\gamma}, \mathbf{x}^{\prime}\right)=2 \mathbb{M}^{0}\left(\gamma-\mathbf{x}^{\prime}\right)+2 \int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbb{G}(\boldsymbol{\gamma}, \boldsymbol{\alpha}) \mathbb{M}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) \tag{4.9}
\end{equation*}
$$

where the inhomogeneity $\mathbb{M}^{0}$ and the kernel $\mathbb{G}$ are defined via

$$
\begin{align*}
\mathbb{M}^{0}\left(\boldsymbol{\gamma}-\mathbf{x}^{\prime}\right) & \equiv\left(\begin{array}{cc}
-\sigma_{3} \mathbf{G}^{\mathrm{N}}\left(\boldsymbol{\gamma}-\mathbf{x}^{\prime}\right) & \boldsymbol{\sigma}_{3} \mathrm{G}^{\mathrm{S}}\left(\boldsymbol{\gamma}-\mathbf{x}^{\prime}\right) \\
\boldsymbol{\sigma}_{3} \partial_{\gamma} \mathbf{G}^{\mathrm{N}}\left(\boldsymbol{\gamma}-\mathrm{x}^{\prime}\right) & -\boldsymbol{\sigma}_{3} \partial_{\gamma} \mathbf{G}^{\mathrm{S}}\left(\boldsymbol{\gamma}-\mathrm{x}^{\prime}\right)
\end{array}\right) ;  \tag{4.10}\\
\mathbb{G}(\boldsymbol{\gamma}, \boldsymbol{\alpha}) & \equiv\left(\begin{array}{cc}
-\boldsymbol{\sigma}_{3} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\gamma-\boldsymbol{\alpha}) & \boldsymbol{\sigma}_{3} \mathrm{G}^{\mathrm{S}}(\boldsymbol{\gamma}-\boldsymbol{\alpha}) \\
\boldsymbol{\sigma}_{3} \partial_{\gamma} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\boldsymbol{\gamma}-\boldsymbol{\alpha}) & -\boldsymbol{\sigma}_{3} \partial_{\gamma} \mathrm{G}^{\mathrm{S}}(\boldsymbol{\gamma}-\boldsymbol{\alpha})
\end{array}\right) . \tag{4.11}
\end{align*}
$$

Equation (4.9) is one of the central elements of this Paper. Its output ( $\boldsymbol{\mu}^{\mathrm{ii}}$ etc.) can be fed into Eq. (4.1) to obtain the ingredients ( $\mathbf{G}^{\mathrm{ii}}$, etc.) of $\mathbf{G}$ In terms of our $4 \times 4$ notation, this connection becomes

$$
\begin{align*}
\left(\begin{array}{cc}
\mathrm{G}^{\mathrm{ii}}\left(\mathrm{x}, \mathrm{x}^{\prime}\right) & \mathrm{G}^{\mathrm{io}}\left(\mathrm{x}, \mathrm{x}^{\prime}\right) \\
\mathrm{G}^{\mathrm{oi}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) & \mathrm{G}^{\mathrm{oo}}\left(\mathbf{x}, \mathrm{x}^{\prime}\right)
\end{array}\right)= & \left(\begin{array}{cc}
\mathrm{G}^{\mathrm{N}}\left(\mathrm{x}, \mathrm{x}^{\prime}\right) & 0 \\
0 & \mathrm{G}^{\mathrm{S}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)
\end{array}\right)  \tag{4.12}\\
& +\int_{\partial \mathcal{V}} d \sigma_{\gamma}\left(\begin{array}{cc}
\partial_{\gamma} \mathbf{G}^{\mathrm{N}}(\mathbf{x}, \gamma) & \mathbf{0} \\
\mathbf{0} & \mathrm{G}^{\mathrm{S}}(\mathbf{x}, \gamma)
\end{array}\right) \mathbb{M}\left(\gamma, \mathbf{x}^{\prime}\right) .
\end{align*}
$$

When the iterative solution of Eq. (4.9) is fed into Eq. (4.12) for the Green function, the resulting expansion is called a Multiple Scattering Expansion (MSE) for the Green function. A typical term in this expansion is shown diagrammatically in Fig. 4.1. The physical content of the MSE is this: the iterations generate terms that correct the free Green function by accounting for multiple scatterings from the superconducting condensate surrounding the billiard. However, the expansion is not simply a perturbation expansion in powers of the pair-potential; instead, terms involving $n$ transmissions (i.e. terms with $\mathbf{G}^{\mathrm{N}}$ followed by $\mathbf{G}^{S}$ and vice versa) account nonperturbatively for all


Figure 4.1: Typical term in the MSE for the Green function. Lines running internally (externally) to the billiard represent homogeneous-region normal (superconducting) Green functions $\mathbf{G}^{\mathrm{N}}\left(\mathbf{G}^{\mathrm{S}}\right)$; each point on the boundary ( $\boldsymbol{\alpha}, \boldsymbol{\beta}$, etc.) at which a scattering event occurs is to be integrated over the complete boundary.

Feynman trajectories that traverse the boundary $n$ times, thus spending intervals in the superconducting region. (The reason such a re-organization of the simple perturbation expansion in powers of the pair-potential is possible is that one knows fully the fundamental Green functions that describe propagation in homogeneous N or S regions.)

### 4.2 Reorganized multiple scattering expansion

The multiple scattering expansion for the Green function was constructed by iterating a four-by-four matrix integral equation, viz., Eq. (4.9). In fact, this $4 \times 4$ structure consists of two substructures: (i) the electron-hole structure, which is an essential ingredient in Andreev billiards. (In fact it is an essential ingredient for any system involving superconductivity.) (ii) the inside-outside structure: this structure is specific to Andreev billiards, as they consist of two distinct regions separated by a boundary. As for the inside-outside structure, it is possible to diagonalize this (as we shall soon show), and thus to obtain a $2 \times 2$ matrix integral equation, whose iteration produces exactly the same MSE as the $4 \times 4$ matrix integral equation does. In this way, we reduce the problem of determining the full-space Green function to an effective, but nonetheless exact interior (or, if one wishes, exterior) problem. (In-
deed, in the following section we shall obtain effective boundary conditions for the interior Green function problem, the solution of which coincides with the that for the full-space problem.)

The advantages of this formulation are two-fold: First it allows us to easily obtain the (Feynman) rules for evaluating a generic term in the MSE. Second, as we shall see when we develop approximation schemes for the Green function, it is especially well suited for the task of integrating out the outside-propagation processes and, thus, obtaining an effective Multiple Reflection Expansion. In this way, the physics of Andreev reflection as well as the corrections associated with charge-preserving reflection will become evident.

To make this reorganization, let us introduce the operators $\mathbb{D}$ (diagonal) and $\mathbb{O}$ (off-diagonal), defined as follows by their action on 4-component functions:

$$
\begin{align*}
& \mathbb{D} \boldsymbol{\Psi}(\boldsymbol{\gamma}) \equiv 2 \int_{\partial \mathcal{V}} d \sigma_{\beta}\left(\begin{array}{cc}
-\boldsymbol{\sigma}_{3} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\boldsymbol{\gamma}-\boldsymbol{\beta}) & \mathbf{0} \\
\mathbf{0} & -\boldsymbol{\sigma}_{3} \partial_{\gamma} \mathbf{G}^{\mathrm{S}}(\boldsymbol{\gamma}-\boldsymbol{\beta})
\end{array}\right) \boldsymbol{\Psi}(\boldsymbol{\beta}),(  \tag{4.13}\\
& \mathbb{O} \boldsymbol{\Psi}(\boldsymbol{\gamma}) \equiv 2 \int_{\partial \mathcal{V}} d \sigma_{\beta}\left(\begin{array}{cc}
\boldsymbol{0} & \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}}(\boldsymbol{\gamma}-\boldsymbol{\beta}) \\
\boldsymbol{\sigma}_{3} \partial_{\gamma} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\boldsymbol{\gamma}-\boldsymbol{\beta}) & \mathbf{0}
\end{array}\right) \boldsymbol{\Psi}(\boldsymbol{\beta}) \tag{4.14}
\end{align*}
$$

These operators constitute the (two-by-two block) diagonal and off-diagonal operator elements of the kernel $\mathbb{G}$ in the integral equation (4.9), the iterative solution of which generates the MSE. Note that we write the four-by-four identity as $\mathbb{I}$. In terms of $\mathbb{D}$ and $\mathbb{O}$ Eq. (4.9) can be written symbolically (i.e. using an obvious condensed notation) as

$$
\begin{equation*}
\mathbb{M}=2 \mathbb{M}^{0}+2(\mathbb{D}+\mathbb{O}) \mathbb{M} \tag{4.15}
\end{equation*}
$$

To obtain the reorganized MSE, first we rewrite Eq. (4.15) as

$$
\begin{equation*}
(\mathbb{I}-2 \mathbb{D}) \mathbb{M}=2 \mathbb{M}^{0}+2 \mathbb{O M} \tag{4.16}
\end{equation*}
$$

and then, by inverting the operator $(\mathbb{I}-2 \mathbb{D})$, we obtain

$$
\begin{equation*}
\mathbb{M}=(\mathbb{I}-2 \mathbb{D})^{-1} 2 \mathbb{M}^{0}+(\mathbb{I}-2 \mathbb{D})^{-1} 2 \mathbb{O M} \tag{4.17}
\end{equation*}
$$

in which the kernel has become (block) off-diagonal. Next, in order to obtain a (block) diagonal structure, we iterate this equation once, thus arriving at
$\mathbb{M}=2(\mathbb{I}-2 \mathbb{D})^{-1} \mathbb{M}^{0}+4(\mathbb{I}-2 \mathbb{D})^{-1} \mathbb{O}(\mathbb{I}-2 \mathbb{D})^{-1} \mathbb{M}^{0}+4(\mathbb{I}-2 \mathbb{D})^{-1} \mathbb{O}(\mathbb{I}-2 \mathbb{D})^{-1} \mathbb{O M}$.

At this stage it is useful to define the kernel $\mathbb{K}$ obeying

$$
\begin{equation*}
(\mathbb{I}-2 \mathbb{D}) \mathbb{K}=\mathbb{I} \tag{4.19}
\end{equation*}
$$

This kernel is block diagonal,

$$
\mathbb{K}=\left(\begin{array}{cc}
\mathbf{K}^{\mathrm{ii}} & \mathbf{0}  \tag{4.20}\\
\mathbf{0} & \mathbf{K}^{\mathrm{oo}}
\end{array}\right)
$$

and the diagonal (two-by-two) blocks obey

$$
\begin{align*}
\mathbf{K}^{\mathrm{ii}}+2 \boldsymbol{\sigma}_{3} \partial \mathbf{G}^{\mathrm{N}} \mathbf{K}^{\mathrm{ii}} & =\mathbf{I}  \tag{4.21}\\
\mathbf{K}^{\mathrm{oo}}+2 \boldsymbol{\sigma}_{3} \delta \mathbf{G}^{\mathrm{S}} \mathbf{K}^{\mathrm{oo}} & =\mathbf{I} . \tag{4.22}
\end{align*}
$$

Here and elsewhere we shall use $\delta$ (resp. $\partial$ ) without a subscript to indicate an inward normal derivative with respect to the first (resp. second) argument of the succeeding Green function, i.e.,

$$
\begin{align*}
\delta \mathbf{G}(\boldsymbol{\alpha}, \boldsymbol{\beta}) & \equiv \partial_{\alpha} \mathbf{G}(\boldsymbol{\alpha}, \boldsymbol{\beta})  \tag{4.23}\\
\partial \mathbf{G}(\boldsymbol{\alpha}, \boldsymbol{\beta}) & \equiv \partial_{\beta} \mathbf{G}(\boldsymbol{\alpha}, \boldsymbol{\beta}) \tag{4.24}
\end{align*}
$$

In terms of $\mathbb{K}$, Eq. (4.18), when pre-multiplied by $(\mathbb{I}-2 \mathbb{D})$, becomes

$$
\begin{equation*}
(\mathbb{I}-2 \mathbb{D}) \mathbb{M}=2 \mathbb{M}^{0}+4 \mathbb{O} \mathbb{K} \mathbb{M}^{0}+4 \mathbb{O} \mathbb{K} \mathbb{O} \mathbb{M} \tag{4.25}
\end{equation*}
$$

the upper-left (two-by-two) block of which can be rearranged to read

$$
\begin{equation*}
\boldsymbol{\mu}^{\mathrm{ii}}=-2 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{N}}+4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \mathbf{K}^{\mathrm{oo}} \boldsymbol{\sigma}_{3} \delta \mathbf{G}^{\mathrm{N}}+\left\{-2 \boldsymbol{\sigma}_{3} \partial \mathbf{G}^{\mathrm{N}}+4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \mathbf{K}^{\mathrm{oo}} \boldsymbol{\sigma}_{3} \partial \delta \mathbf{G}^{\mathrm{N}}\right\} \boldsymbol{\mu}^{\mathrm{ii}} \tag{4.26}
\end{equation*}
$$

What we have accomplished via these transformations is the construction of a closed $(2 \times 2)$ system of integral equations for the boundary layer $\boldsymbol{\mu}^{\mathrm{ii}}$, which is all that is needed to complete the computation of the Green function $\mathrm{G}^{\mathrm{ii}}$. The virtue of this transformation is that it facilitates the subsequent elimination (via the integrating out of processes involving virtual propagation in the superconducting region) of the states located in the superconducting region. This elimination can now be made straightforwardly, owing to the fact that all superconducting Green functions are now conveniently located in the kernel in Eq. (4.26).

The reorganization just described also allows one to identify the following rules for the construction of all possible contributions at any order $n(=1,2,3, \ldots)$ to the inside-to-inside Green function $\mathbf{G}^{\mathrm{ii}}$ :

1. Write down all possible permutations of $\mathbf{G}^{\mathbf{N}}$ and $\mathbf{G}^{\mathrm{S}}$ (having a total of $n+1$ Green functions), subject to the restriction that the first and last Green functions are $\mathbf{G}^{\mathrm{N}}$. 2. Associate to each permutation a factor $(-1)^{i_{\mathrm{N}}+1}$, where $i_{\mathrm{N}}$ is the number of $\mathbf{G}^{\mathrm{N}}$ factors.
2. Furnish all $\mathbf{G}^{\mathrm{N}}$ factors (except the last) with normal derivatives acting on their second arguments; all $\mathbf{G}^{\text {S }}$ factors carry no normal derivatives.
3. Furnish any Green function factor that follows a $\mathbf{G}^{S}$ factor with an additional normal derivative acting on its first argument.
4. Insert a Pauli-matrix factor $\boldsymbol{\sigma}_{3}$ before every $\mathbf{G}^{\mathrm{N}}$ and $\mathbf{G}^{\mathrm{S}}$ except the first.

In this way one can construct $\mathrm{G}^{\mathrm{ii}}$. An example is provided by the process depicted in Fig. 4.1, for which the corresponding amplitude is

$$
\int_{\partial \mathcal{V}} d \sigma_{\alpha} d \sigma_{\beta} d \sigma_{\gamma} d \sigma_{\delta} \partial \mathbf{G}^{\mathrm{N}}(\mathbf{x}, \boldsymbol{\alpha}) \boldsymbol{\sigma}_{3} \partial \mathbf{G}^{\mathrm{N}}(\boldsymbol{\alpha}, \boldsymbol{\beta}) \boldsymbol{\sigma}_{3} \partial \mathbf{G}^{\mathrm{N}}(\boldsymbol{\beta}, \boldsymbol{\gamma}) \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}}(\boldsymbol{\gamma}, \boldsymbol{\delta}) \boldsymbol{\sigma}_{3} \delta \mathbf{G}^{\mathrm{N}}\left(\boldsymbol{\delta}, \mathbf{x}^{\prime}\right)
$$

It is worth noting that the resulting series features terms containing two or more consecutive factors of $\mathbf{G}^{\mathrm{N}}$. As $\mathbf{G}^{N}$ is diagonal, such terms correspond to electron-to-electron and hole-to-hole reflection processes. At first sight, the presence of such terms might be disconcerting, given the charge-interconverting character of Andreev
reflection. However, it should be recalled that not only does the superconducting surround interconvert electrons and holes, but also it confines these quasiparticles to the normal region. For example, consider the series of terms that contain no superconducting Green functions $\mathbf{G}^{S}$. This series is precisely the Dirichlet series obtained in BB-I. Furthermore, this series can be embedded in any term of the MSE, which amounts to replacing the fundamental (i.e. unconfined) Green function by a suitably confined Green function. Therefore, terms involving consecutive factors of $\mathbf{G}^{\mathrm{N}}$, rather than being disconcerting, are necessary to correct the free-propagation term, doing so by cancelling the Feynman paths that venture into the superconductor.

### 4.3 Effective boundary conditions

Before proceeding with our main issues (viz. the computation of the BDG Green function inside the billiard), we pause to pose and answer two questions: (i) Is there any boundary condition that can be imposed on $\mathbf{G}^{\mathrm{ii}}$ so that $\mathrm{G}^{\mathrm{ii}}$ can be computed by solving the BDG Green function equation solely in $\mathcal{V}$, i.e., without any reference to the region $\overline{\mathcal{V}}$. And if so, (ii) what is the precise form of this boundary condition? (If such an approach turns out to possible then one could dispense with the cumbersome task of dealing with Green functions having arguments outside $\mathcal{V}$, as well as the concomitant need to match Green functions across the boundary.)

To see that such a boundary condition does indeed exist, and to determine its explicit form, we substitute into Eq. (4.26) the parametrization of $\mathrm{G}^{\mathrm{ii}}$ given in Eq. (4.1) in terms of $\boldsymbol{\mu}^{\mathrm{ii}}$. Then all reference to $\boldsymbol{\mu}^{\mathrm{ii}}$ cancels, and we arrive at a nonlocal and billiard-shape-dependent effective boundary condition obeyed by $\mathrm{G}^{\mathrm{ii}}$, viz.,

$$
\begin{equation*}
\mathbf{G}^{\mathrm{ii}}=-2 \mathbf{G}^{\mathrm{S}} \mathbf{K}^{\mathrm{oo}} \boldsymbol{\sigma}_{3} \delta \mathbf{G}^{\mathrm{ii}} \tag{4.27}
\end{equation*}
$$

The reason that the boundary condition is nonlocal is that there exists the possibility of virtual propagation within the superconductor surrounding the billiard. The
reason that the boundary condition is shape-dependent is that this virtual propagation outside the billiard is modified (from the value it would have in a homogeneous superconductor) due to the presence of the normal region.

## Chapter 5

## Asymptotics of the multiple scattering expansion

Up to the present point, our investigation of the Green function for the BDG wave equation has been exact, and we have developed the machinery - the MSE-for computing this Green function in terms of the fundamental ( N and S ) Green functions and the shape of the billiard. The construction, however, is in terms of an infinite series, each term in which involves repeated integration over the boundary of the billiard. Thus, the direct computation of an arbitrary term in this series is prohibitively difficult, unless the shape of the billiard is exceptionally simple. To make progress we therefore need to invoke some approximation scheme and, as the Fermi wavelength is taken to be much smaller than the characteristic linear dimension of the billiard $L$, a very natural one to consider is the semiclassical approximation. In the present setting, this involves the evaluating of the repeated boundary integrals via a shortwave asymptotic approximation scheme. What we mean by this is that we seek an asymptotic approximation for every term in the MSE, the expansion parameter being $1 / k_{\mathrm{F}} L$, where $k_{\mathrm{F}}$ is the Fermi wave vector; having invoked such an approximation, we shall re-sum the MSE.

By following this scheme we shall be able to obtain, inter alia, the Green function
for single-particle excitations, as well as a trace formula for the oscillatory part of the density of energy eigenvalues, valid in the semiclassical regime. What we mean by a trace formula is an explicit formula for the oscillatory part of the density of energy eigenvalues, expressed in terms of a sum over all closed semiclassical particle orbits. As we shall see, by virtue of the retro-reflective character of Andreev reflection, this sum over particle orbits is quite distinct from that arising in the setting of conventional billiards.

Our approach has the virtue of delivering results not only for the DOS at the coarsest of energy resolutions (i.e. the Andreev level of approximation, in which motion is confined to chords traversing the billiard) but also at the finer level, thus revealing the mesoscale oscillations due to the quantal particle motion transverse to each chord.

### 5.1 Classical dynamics in Andreev billiards

The purpose of this subsection is to make a brief intermezzo in which discuss the physics of Andreev reflection from the point of view of geometrical optics. To this end, we develop stationary phase arguments aimed at elucidating the origin of the retro-reflective character of Andreev reflection. Along the way, we shall see that owing to the difference in the wavelengths of the incoming and reflected quasiparticles there is imperfectness in the this retro-reflection, i.e., the reflected excitation does not, in general, precisely retrace the path of the incoming one. Arguments of this type will be useful, subsequently, when we come to incorporate quantum fluctuations around the classical trajectories associated with Andreev reflection.

It is well known that Eq. (3.6) gives rise to the Andreev reflection phenomenon, in which the electrons arriving from the normal metal are converted into holes (and vice versa) at the superconductor boundary, and are retro-reflected (i.e. have the


Figure 5.1: Geometry for Andreev reflection from a planar surface. Initial and final points (i.e. the off-boundary points) are kept fixed, and the point of reflection from the boundary is then determined via the stationarity of the corresponding phase.
excitation velocity reversed). In the present section we discuss the classical limit of this reflection process by making use of the method of stationary phase (i.e. via the principle of least action). Throughout this section we make the (physical-optics-type) assumption that an electron wave having energy $E$ and traveling a distance $r$ acquires a phase

$$
\begin{equation*}
\mathrm{e}^{i k_{+} r} \tag{5.1}
\end{equation*}
$$

whereas a hole traveling the same direction acquires the phase

$$
\begin{equation*}
\mathrm{e}^{-i k-r}, \tag{5.2}
\end{equation*}
$$

where $k_{+}$and $k_{-}$are, respectively, the wavevectors appropriate for for particle and hole motion in the normal region. The energy dependence of these wavevectors is given by

$$
\begin{equation*}
k_{ \pm}=\sqrt{\mu \pm E} \tag{5.3}
\end{equation*}
$$

Following the standard optics-type approach, we envision some process (for an example see Fig. 5.1) and then, by using Eqs. (5.1) and (5.2), we calculate the total phase acquired during this process. In the classical limit, the dominant process is the one
(or ones) that make stationary this total phase, and hence determines information such as relationships between angles of incidence and reflection. Thus, in effect we are finding the rules of classical dynamics for Andreev billiards.

At this point we have to introduce the physics of Andreev reflection "by hand,' and do so by requiring that after one reflection from the billiard boundary an electron is converted into a hole (and vice versa). Thus we are demanding that there is no electron-to-electron (or hole-to-hole) scattering (due to a single reflection). Under these conditions, any scattering process can be analyzed in terms of the basic electron-to-hole and hole-to-electron processes. We need only examine one of these because the corresponding phases are identical and thus, at stationarity the two processes have the same geometry. In other words, the stationary path describing an incoming electron and scattered hole may be reversed (by reversing the direction of propagation of each particle) to give the stationary path of an incoming hole and the scattered electron.


Figure 5.2: The analogy between Andreev reflection and optical refraction

It might be useful to note the similarity between the phenomena of Andreev reflection and the the refraction of light. The common feature is that, in both settings, before and after scattering the waves have the same frequency but differing wavevector
magnitudes. In the case of the refraction of light, the wavevector is changed because the wave enters a medium with a distinct index of refraction. In the case of Andreev reflection, the wavevector is instead changed because the electron wave is converted into a hole wave, the latter having a distinct dispersion relation. In fact, by reversing the sign of the phase for a hole wave, as well as the direction of propagation of the hole wave, one transforms the Andreev reflection process into the familiar optical refraction process. Thus, one has an electron-to-hole reflection law that is essentially identical to Snell refraction (rather than specular reflection). Thus, Andreev reflection looks like optical refraction but with the outgoing direction reversed with respect to the scattering point (see Fig. 5.2).

To quantify these remarks, consider the problem depicted in Fig. 5.1, in which an electron arriving from a fixed point (1) is reflected and converted at the variable point $x$ into a hole, which then propagates to another fixed point (2). The classical path corresponds to the value of $x$ at which the total phase for the process is stationary with respect to variations of $x$. For the process at hand, total phase is given by

$$
\begin{equation*}
k_{+} \sqrt{h_{+}^{2}+x^{2}}-k_{-} \sqrt{h_{-}^{2}+(x-l)^{2}} \tag{5.4}
\end{equation*}
$$

for which the stationarity condition reads

$$
\begin{equation*}
k_{+} \frac{x}{\sqrt{h_{+}^{2}+x^{2}}}-k_{-} \frac{x-l}{\sqrt{h_{-}^{2}+(x-l)^{2}}}=0 . \tag{5.5}
\end{equation*}
$$

By rewriting this condition in terms of the angles of incidence and reflection (i.e. $\theta_{+}$ and $\theta_{-}$shown in Fig. 5.1) one recovers the Snell's law form:

$$
\begin{equation*}
k_{+} \sin \theta_{+}=k_{-} \sin \theta_{-} . \tag{5.6}
\end{equation*}
$$

By using Eq. (5.6) one can construct the stationary paths for an Andreev billiard, just as one does in the case of geometrical optics. When there is more than one reflection, Eq. (5.6) must be satisfied at each one.

One feature of Eq. (5.6) is that it makes evident the fact that the reflected particle is not, in general, perfectly retro-reflected (i.e. $\theta_{+} \neq \theta_{-}$). However, when $k_{+}$and $k_{-}$
are very close to each other, $\theta_{+}$and $\theta_{-}$will be, too. Let us now calculate this small deflection angle $\theta_{-}-\theta_{+}$in terms of $\mu$ and $E$. To do this, let us assume that $\theta_{+}$and $\theta$ - are close and that $E / \mu \ll 1$, and expand Eq. (5.6) to obtain

$$
\begin{equation*}
\theta_{-}-\theta_{+} \approx(E / \mu) \tan \theta_{+} . \tag{5.7}
\end{equation*}
$$

As expected, the deflection angle is $\mathcal{O}(E / \mu)$, unless the incident direction grazes the boundary. This qualification divides the space of incoming trajectories into two classes: (i) a large fraction, occupying most of the phase space, in which $\tan \theta_{+}$is of order unity, and (ii) the rest, in which the deflection angle $\theta_{-}-\theta_{+}$is not small. Equation (5.7), although approximate, provides a guide for addressing whether or not deflections (i.e. imperfectness in retro-reflection) needs to be taken into account.


Figure 5.3: Typical closed classical trajectory in an Andreev billiard. Black lines depict electron paths; gray lines depict hole paths. The imperfectness of the retroreflection is exaggerated.

We are now at the point where we can construct classical dynamics in an Andreev billiard. For any trajectory, the reflection rule in Eq. (5.6) has to be satisfied at every reflection point. This will generate a type of dynamics that differs from that generated by the usual specular reflection rule (in which the outgoing angle equals
the incoming one). An example of a closed trajectory in an Andreev billiard is shown in Fig. 5.3.

### 5.2 Asymptotics of the fundamental Green functions and their derivatives

In the present section we investigate the asymptotic behavior of the fundamental Green functions $\mathbf{G}^{\mathrm{N}}$ and $\mathbf{G}^{S}$ for both small and large values of their (position) arguments. This investigation will allow us to estimate the relative dominance of various processes, and thus to organize the multiple scattering expansion for the exact BDG Green function into a form suitable for establishing its approximate behavior at large $k_{F} L$.

The asymptotic behavior of $\mathbf{G}^{\mathrm{N}, \mathrm{S}}$ is related to the corresponding Helmholtz Green function $g_{ \pm}^{\mathrm{N}, \mathrm{S}}$, which can be represented through the Fourier integral

$$
\begin{equation*}
g_{ \pm}^{\mathrm{N}, \mathrm{~S}}(\mathbf{l})=\int \frac{d^{d} p}{(2 \pi)^{d}} \frac{\mathrm{e}^{i \mathbf{p} \cdot \mathbf{l}}}{p^{2}-k^{2}} \tag{5.8}
\end{equation*}
$$

where $k=k_{ \pm}^{\mathrm{N}, \mathrm{S}}$, depending on the Green function in question. Then the asymptotic behavior of $g_{ \pm}^{\mathrm{N}, \mathrm{S}}$ for large $k l$ can be obtained from the asymptotic evaluation of this Fourier integral, which gives

$$
\begin{equation*}
g_{ \pm}^{\mathrm{N}, \mathrm{~S}}(l) \approx \pm i\left(\frac{k_{ \pm}^{\mathrm{N}, \mathrm{~S}}}{2 \pi l}\right)^{\frac{d-1}{2}} \frac{\exp \left( \pm i k_{ \pm}^{\mathrm{N}, \mathrm{~S}} l \mp i \pi(d-1) / 4\right)}{2 k_{ \pm}^{\mathrm{N}, \mathrm{~S}}} \tag{5.9}
\end{equation*}
$$

The derivatives of the Green functions for large $k l$ can be obtained by differentiating this asymptotic expression.

Determining the small $k l$ asymptotics of $g_{ \pm}^{\mathrm{N}, \mathrm{S}}(l)$ is more tricky. By scaling $p$ with $l$ we get

$$
\begin{equation*}
g(l)=l^{(2-d)} \int \frac{d^{d} a}{(2 \pi)^{d}} \frac{\mathrm{e}^{i \boldsymbol{a} \cdot \hat{\mathrm{I}}}}{a^{2}-(k l)^{2}}, \tag{5.10}
\end{equation*}
$$

where $\boldsymbol{a} \equiv \mathbf{p} l$, and $g$ is shorthand for any of the four $g_{ \pm}^{\mathrm{N}, \mathrm{S}}$. Thus for small $k l$ we have

$$
g(\mathbf{l}) \sim \begin{cases}l^{(2-d)}, & \text { if } d \neq 2  \tag{5.11}\\ \ln (k l), & \text { if } d=2\end{cases}
$$



Figure 5.4: Approximating the surface by its tangent plane
Having determined the form of $g$ for small $k l$, we now investigate what can be said about $g(l), \partial g(l)$ and $\partial \delta g(l)$ when $\mathbf{l}$ is a vector connecting two nearby points on the surface $\partial \mathcal{V}$. The geometry (for the $d=3$ case ) is illustrated in Fig. 5.4. In this figure, $\mathbf{l}=\boldsymbol{\beta}-\boldsymbol{\alpha}$, where $\boldsymbol{\beta}$ and $\boldsymbol{\alpha}$ are points on the surface. In the following, we shall work in $d=3$, although results obtained will be applicable for all $d$. Without loss of generality, we choose $\boldsymbol{\alpha}$ to be the origin of our coordinate system, with the $z$ direction coinciding with the inward normal direction of the surface at $\boldsymbol{\alpha}$. The remaining directions are chosen arbitrarily (at least for the time being), the only constraint being that the coordinate-system be right handed. In this coordinate system the surface may be defined locally through an equation of the form $z=f(x, y)$, where $x$ and $y$ span the tangent plane, and $\left.\partial_{x} f\right|_{(x, y)=(0,0)}=\left.\partial_{y} f\right|_{(x, y)=(0,0)}=0$. Then a point $\boldsymbol{\beta}$ (near $\boldsymbol{\alpha}$ ) on the surface will approximately have the coordinates

$$
\begin{equation*}
\boldsymbol{\beta}=\left(x, y,\left.\frac{\partial^{2} f}{\partial x^{2}}\right|_{0} x^{2}+\left.\frac{1}{2} \frac{\partial^{2} f}{\partial x \partial y}\right|_{0} x y+\left.\frac{\partial^{2} f}{\partial y^{2}}\right|_{0} y^{2}\right) . \tag{5.12}
\end{equation*}
$$

However, by a suitable rotation within the tangent plane the equation for $\boldsymbol{\beta}$ may be written as

$$
\begin{equation*}
\boldsymbol{\beta}=\left(x, y, \frac{x^{2}}{2 R_{1}}+\frac{y^{2}}{2 R_{2}}\right), \tag{5.13}
\end{equation*}
$$

where $R_{1}$ and $R_{2}$ are the two principal radii of curvature of the surface at the point $\boldsymbol{\alpha}$. For reasons that should soon become clear, we define the vector $\boldsymbol{\beta}_{/ /}$to be the projection of $\boldsymbol{\beta}$ on to the tangent plane at $\boldsymbol{\alpha}$. Next, we make two assumptions:

1. the radii of curvature are on the order of the linear size of the billiard, i.e.,

$$
R_{1,2} / L=\mathcal{O}(1)
$$

2. the region we of interest around $\boldsymbol{\alpha}$ has a linear size on the order of $k_{F}^{-1}$.

Under these assumptions, we have that

$$
\begin{equation*}
l \equiv|\boldsymbol{\beta}-\boldsymbol{\alpha}|=\left|\boldsymbol{\beta}_{/ /}\right|\left\{1+\mathcal{O}\left(\left(k_{F} L\right)^{-2}\right)\right\} \tag{5.14}
\end{equation*}
$$

Let us now turn to $\partial g$. In the coordinate system specified above, the normal vector $\mathbf{n}_{\beta}$ at the surface point $\boldsymbol{\beta}$ is given by

$$
\begin{equation*}
\mathbf{n}_{\beta}=\left(\frac{x}{R_{1}}, \frac{y}{R_{2}}, 1\right)+\mathcal{O}\left(\left(k_{F} L\right)^{-2}\right) . \tag{5.15}
\end{equation*}
$$

Then the quantity $\partial l$ is given by

$$
\begin{equation*}
\partial l \equiv \mathbf{n}_{\beta} \cdot \boldsymbol{\nabla}_{\beta}|\boldsymbol{\beta}-\boldsymbol{\alpha}|=\frac{1}{\mid \boldsymbol{\beta}_{/ / \mid}}\left(\frac{x^{2}}{R_{1}}+\frac{y^{2}}{R_{2}}\right)+\mathcal{O}\left(\left(k_{F} L\right)^{-2}\right)=\mathcal{O}\left(\frac{l}{L}\right) . \tag{5.16}
\end{equation*}
$$

Hence, we have that the normal derivative of the Green function is given by

$$
\begin{equation*}
\partial g(l) \approx \frac{\partial g(l)}{\partial l} \frac{1}{l}\left(\frac{x^{2}}{R_{1}}+\frac{y^{2}}{R_{2}}\right) . \tag{5.17}
\end{equation*}
$$

By generalizing to arbitrary dimensionality $d$ we obtain

$$
\begin{equation*}
\partial g(l)=\mathcal{O}\left(L^{-1} l^{(2-d)}\right) \tag{5.18}
\end{equation*}
$$

In order to evaluate $\partial \delta g$, we consider a third point on $\partial \mathcal{V}$, which we denote by $\gamma$, focus on the quantity $\partial \delta l^{\prime}$ (where $l^{\prime} \equiv|\boldsymbol{\beta}-\gamma|$ ) and, at the end of our calculation, let
$\boldsymbol{\gamma}$ tend to $\boldsymbol{\alpha}$. We shall use primed coordinates $x^{\prime}$ and $y^{\prime}$ for $\boldsymbol{\gamma}$. In this way, we find that

$$
\begin{align*}
\partial \delta l^{\prime} & \equiv\left(\mathbf{n}_{\beta} \cdot \boldsymbol{\nabla}_{\beta}\right)\left(\mathbf{n}_{\gamma} \cdot \boldsymbol{\nabla}_{\gamma}\right)|\boldsymbol{\beta}-\gamma| \\
& \approx\left(\frac{x}{R_{1}} \frac{\partial}{\partial x}+\frac{y}{R_{2}} \frac{\partial}{\partial y}+\frac{\partial}{\partial z}\right)\left(\frac{x^{\prime}}{R_{1}} \frac{\partial}{\partial x^{\prime}}+\frac{y^{\prime}}{R_{2}} \frac{\partial}{\partial y^{\prime}}+\frac{\partial}{\partial z^{\prime}}\right)|\boldsymbol{\beta}-\gamma| \\
& =\left(\frac{\partial^{2}}{\partial z \partial z^{\prime}} \sqrt{\left(x-x^{\prime}\right)^{2}+\left(y-y^{\prime}\right)^{2}+\left(z-z^{\prime}\right)^{2}}\right)_{z=\frac{x^{2}}{2 R_{1}}+\frac{y^{2}}{2 R_{2}}, z^{\prime}=\frac{x^{\prime 2}}{2 R_{1}}+\frac{y^{\prime 2}}{2 R_{2}}} \\
& =-\frac{1}{l^{\prime 3}}\left(\frac{x^{2}-x^{\prime 2}}{2 R_{1}}+\frac{y^{2}-y^{\prime 2}}{2 R_{2}}\right)^{2}-\frac{1}{l^{\prime}}=-\frac{1}{l^{\prime}}\left(1+\mathcal{O}\left(\frac{l^{\prime 2}}{L^{2}}\right)\right) . \tag{5.19}
\end{align*}
$$

We are now in a position to evaluate $\partial \delta g$, for which we find

$$
\begin{equation*}
\partial \delta g(l)=\frac{\partial g}{\partial l} \partial \delta l+\frac{\partial^{2} g}{\partial l^{2}}(\partial l)(\delta l)=-\frac{1}{l} \frac{\partial g}{\partial l}\left(1+\mathcal{O}\left(\frac{l^{\prime 2}}{L^{2}}\right)\right) . \tag{5.20}
\end{equation*}
$$

In the MSE, for each term that includes the product $g \partial \delta g$ there is a corresponding term in which $g \partial \delta g$ is replaced by $\partial g \partial g$. It is therefore desirable to estimate relative size of these terms for small values of their arguments. The asymptotic formulas given in the pressent section are useful for this comparison, giving

$$
|\partial g(l) \partial g(l)|=|g(l) \partial \delta g(l)| \times \begin{cases}\mathcal{O}\left(l^{2} / L^{2}\right), & \text { for } d \neq 2  \tag{5.21}\\ \mathcal{O}\left(l^{2} / L^{2} \ln k l\right), & \text { for } d=2\end{cases}
$$

### 5.3 Asymptotic expansion for the quantal amplitude

The MSE of Sec. 4 has provided us with a series expansion for the Green function (i.e. the full quantum-mechanical amplitude for the propagation of a quasiparticle excitation from one point $\mathbf{x}$ to another $\mathbf{x}^{\prime}$ ). This series expresses the correction to the free-space propagation of quasiparticles caused by multiple scattering processes of the quasiparticles from the boundary that separates the normal and superconducting regions of the billiard. A generic term features the following possible processes: (i) inner reflections (which are marked by the occurrence of an adjacent pair of normal-state

Green functions in the algebraic expression for the contribution), (ii) transmissions (marked by the occurrence of an adjacent pair of Green functions, one normal and one superconducting), (iii) and outer reflections (marked by the occurrence of an adjacent pair of superconducting-state Green functions). Throughout the present section, we shall assume that $d=3$. (The extension of the following discussion to general $d$ is straightforward.)

The generic contribution to the amplitude involving a total of $n$ reflections and scatterings can be written as

$$
\begin{equation*}
\mathcal{A}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \equiv \int \mathcal{M}\left(\mathbf{x}, \boldsymbol{\alpha}_{1}, \boldsymbol{\alpha}_{2}, \ldots, \boldsymbol{\alpha}_{n}, \mathbf{x}^{\prime}\right) \exp i k_{\mathrm{F}} \mathcal{S}\left(\mathbf{x}, \boldsymbol{\alpha}_{1}, \boldsymbol{\alpha}_{2}, \ldots, \boldsymbol{\alpha}_{n}, \mathbf{x}^{\prime}\right) \tag{5.22}
\end{equation*}
$$

where the integral is taken over all values of $\left\{\boldsymbol{\alpha}_{1}, \boldsymbol{\alpha}_{2}, \ldots, \boldsymbol{\alpha}_{n-1}, \boldsymbol{\alpha}_{n}\right\}$, each element ranging over the surface $\partial \mathcal{V}$. The modulus function $\mathcal{M}\left(\mathbf{x}, \boldsymbol{\alpha}_{1}, \boldsymbol{\alpha}_{2}, \boldsymbol{\alpha}_{3}, \ldots, \boldsymbol{\alpha}_{n}, \mathbf{x}^{\prime}\right)$ is real and, as can be seen from the iterative solution of Eqs. (4.1-4.7), is a sum of products of functions such as $\left|\boldsymbol{\alpha}_{i}-\boldsymbol{\alpha}_{i+1}\right|^{-1}$, first and second normal derivatives of this function, as well as a polynomial in $k_{\mathrm{F}}$ and $k_{ \pm}^{\mathrm{N}, \mathrm{S}}$. The phase function $\mathcal{S}\left(\mathbf{x}, \boldsymbol{\alpha}_{1}, \boldsymbol{\alpha}_{2}, \boldsymbol{\alpha}_{3}, \ldots, \boldsymbol{\alpha}_{n}, \mathbf{x}^{\prime}\right)$ is, in general, complex and, as can also be seen from the iterative solution of Eqs. (4.14.7), is a sum of terms each of the form $\left(k_{ \pm}^{\mathrm{N}, \mathrm{S}} / k_{\mathrm{F}}\right)\left|\boldsymbol{\alpha}_{i}-\boldsymbol{\alpha}_{i+1}\right|$.

The approximation scheme that we shall invoke involves the short-wave asymptotic expansion of this quantal amplitude $\mathcal{A}$, valid for large $k_{\mathrm{F}} L$. The method used for the construction of this expansion is the asymptotic expansion of the multiple integrals appearing in the terms in the MSE. (See, e.g., Ref. [33] for a discussion of the asymptotic expansion of multiple Fourier integrals.) This method allows one to approximate $\mathcal{A}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$ as a sum over critical points (c.p.) of the domain of integration (which, in this case, is the $2 n$-dimensional manifold $\mathfrak{M}=\partial \mathcal{V} \times \partial \mathcal{V} \times \cdots \times \partial \mathcal{V}$ ):

$$
\begin{equation*}
\mathcal{A}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \approx \sum_{\text {c.p. }} \mathcal{A}_{\text {c.p. }}\left(\mathrm{x}, \mathrm{x}^{\prime}\right) \tag{5.23}
\end{equation*}
$$

where $\mathcal{A}_{\text {c.p. }}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$ depends solely on the local properties of $\mathcal{M}$ and $\mathcal{S}$ at the critical point. As $k_{\mathrm{F}} L \rightarrow \infty$, corrections to this formula vanish faster than any power of
$\left(k_{\mathrm{F}} L\right)^{-1}$ [34]. In other words, for large $k_{\mathrm{F}} L$ contributions from the critical points dominate the total amplitude. A critical point $\left(\boldsymbol{\alpha}_{1}^{\mathrm{c}}, \boldsymbol{\alpha}_{2}^{\mathrm{c}}, \boldsymbol{\alpha}_{3}^{\mathrm{c}}, \cdots, \boldsymbol{\alpha}_{n}^{\mathrm{c}}\right)$ can be interpreted as a trajectory (although not necessarily a classical one) in which an excitation travels from point $\mathbf{x}$ to $\mathbf{x}^{\prime}$, along the way scattering at the points $\boldsymbol{\alpha}_{1}^{\mathrm{c}}, \boldsymbol{\alpha}_{2}^{\mathrm{c}}, \ldots, \boldsymbol{\alpha}_{n}^{\mathrm{c}}$, so that $\mathcal{A}_{\text {c.p. }}\left(\mathrm{x}, \mathrm{x}^{\prime}\right)$ can be interpreted as the amplitude corresponding to this trajectory.

Before actually proceeding with the construction of the expansion of the multiple integrals appearing in the terms in the MSE, we first classify the critical points of $\mathfrak{M}$ : these are

1. Points at which the gradient of the phase, i.e., $\nabla_{\boldsymbol{\alpha}_{i}} S\left(\mathbf{x}, \boldsymbol{\alpha}_{1}, \boldsymbol{\alpha}_{2}, \boldsymbol{\alpha}_{3}, \cdots, \boldsymbol{\alpha}_{n}, \mathbf{x}^{\prime}\right)$ vanishes for all $i$.
2. Points at which $\mathcal{M}\left(\mathbf{x}, \boldsymbol{\alpha}_{1}, \boldsymbol{\alpha}_{2}, \cdots, \boldsymbol{\alpha}_{n-1}, \boldsymbol{\alpha}_{n}, \mathbf{x}^{\prime}\right)$ has a singularity.
3. Points at which $\mathcal{M}$ or $\mathcal{S}$ fail to be infinitely differentiable.
4. All points on the boundary of the manifold $\mathfrak{M}$.
5. Points satisfying criteria (1-4) in a mixed sense, i.e., points satisfying criterion (1) within a submanifold of points satisfying criterion (2) within a submanifold of points satisfying criterion (3) within a submanifold of points satisfying criterion (4).

In the present setting, $\mathfrak{M}$ has no boundaries and, thus, there are no Type 4 critical points. As for Type 3 critical points, when $\partial \mathcal{V}$ is infinitely differentiable, so are $\mathcal{M}$ and $\mathcal{S}$ and, hence, there are no Type 3 critical points either [35]. Thus, the only possible types of critical point are (1), (2) and (5). In present case, $\mathcal{M}$ consists of products of functions such as $\left|\boldsymbol{\alpha}_{i}-\boldsymbol{\alpha}_{i+1}\right|^{-1}$ and its first and second normal derivatives. Thus, critical points of types 2 and 5 occur whenever one or more of the propagation distances $\left|\boldsymbol{\alpha}_{i}-\boldsymbol{\alpha}_{i+1}\right|$ vanish. In accordance with the trajectory interpretation of critical points, in which the sequence $\left\{\boldsymbol{\alpha}_{1}, \boldsymbol{\alpha}_{2}, \cdots, \boldsymbol{\alpha}_{n}\right\}$ defines the trajectory, we call the part of the critical trajectory having vanishing propagation distance a zero-length path. Then, all Type 5 critical points can be generated from Type 1 critical points by the insertion of zero-length paths. Stated technically, if
$\left(\boldsymbol{\alpha}_{1}, \boldsymbol{\alpha}_{2}, \cdots, \boldsymbol{\alpha}_{i}, \boldsymbol{\alpha}_{i+1}, \cdots, \boldsymbol{\alpha}_{n-1}\right)$ is a critical point of the [(2n-2)-dimensional] manifold $\mathfrak{M}^{\prime}$ then the point $\left(\boldsymbol{\alpha}_{1}, \boldsymbol{\alpha}_{2}, \cdots \boldsymbol{\alpha}_{i}, \boldsymbol{\alpha}_{i}, \boldsymbol{\alpha}_{i+1}, \cdots, \boldsymbol{\alpha}_{n-1}\right)$ will be a critical point of Type 5 in the [2n-dimensional] manifold $\mathfrak{M}$. Criterion (1) amounts to the familiar stationary-phase approximation for the amplitude function, because the points at which all gradients of $\mathcal{S}$ vanish are the stationary phase points.

Throughout this Paper we are interested in leading-order contributions to quantal amplitudes. Thus, it is useful to determine whether or not a process contributes to the full amplitude at leading order. To do this, we must be able to estimate the order-of-magnitude of contributions from different types of critical point. To this end, let us consider a type 1 critical point (i.e. a trajectory defined via the principle of stationary phase), and a Type 5 critical point constructed from this Type 1 critical point via the insertion of a zero-length path. For the sake of simplicity, let us consider as our Type 1 critical point a very simple amplitude, i.e., one having just one reflection:

$$
\begin{equation*}
\mathcal{A}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\int \partial g(\mathbf{x}, \boldsymbol{\alpha}) g\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) \tag{5.24}
\end{equation*}
$$

where $g$ is a generic Helmholtz Green function. [For the purposes of determining the order-of-magnitude of the contribution from various critical points, whether the Green function is $g^{\mathrm{N}}$ or $g^{\mathrm{S}}$ is irrelevant.] By using the asymptotic formulas for $g$ presented in Sec. 5.2, it is possible to write asymptotically (up to from numerical factors)

$$
\begin{equation*}
\mathcal{A}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \sim \int \frac{i k}{|\mathbf{x}-\boldsymbol{\alpha}|\left|\boldsymbol{\alpha}-\mathbf{x}^{\prime}\right|} \exp \left(i k|\mathbf{x}-\boldsymbol{\alpha}|+i k^{\prime}\left|\boldsymbol{\alpha}-\mathbf{x}^{\prime}\right|\right) \tag{5.25}
\end{equation*}
$$

Here, $k$ and $k^{\prime}$ can be $\pm k_{+,-}^{\mathrm{N}, \mathrm{S}}$. The position $\left(\boldsymbol{\alpha}_{c}\right)$ of the critical point will depend on the chosen values of $k$ and $k^{\prime}$ via the stationarity condition. However, owing to the facts that $k$ and $k^{\prime}$ are both $\mathcal{O}\left(k_{\mathrm{F}}\right)$ and that $\left|\mathbf{x}-\boldsymbol{\alpha}_{c}\right|$ and $\left|\boldsymbol{\alpha}_{c}-\mathbf{x}^{\prime}\right|$ are both $\mathcal{O}(L)$, it is possible to estimate the order-of-magnitude of the contribution associated with $\left(\boldsymbol{\alpha}_{c}\right)$ to be

$$
\begin{align*}
\mathcal{A}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\mathcal{O}\left(k_{\mathrm{F}} / L^{2}\right) & \exp \left(i k\left|\mathbf{x}-\boldsymbol{\alpha}_{c}\right|+i k^{\prime}\left|\boldsymbol{\alpha}_{c}-\mathbf{x}^{\prime}\right|\right) \\
& \times \int d x d y \exp \left(i k_{\mathrm{F}} L\left(A x^{2}+B x y+C y^{2}\right)\right) \tag{5.26}
\end{align*}
$$

where we have expanded the phase to second order in deviations from $\left(\boldsymbol{\alpha}_{c}\right)$. From dimensional considerations we know that $A, B$ and $C$ are all $\mathcal{O}\left(L^{-2}\right)$ and, thus, that

$$
\begin{equation*}
\mathcal{A}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\mathcal{O}(1 / L) \exp \left(i k\left|\mathbf{x}-\boldsymbol{\alpha}_{c}\right|+i k^{\prime}\left|\boldsymbol{\alpha}_{c}-\mathbf{x}^{\prime}\right|\right) \tag{5.27}
\end{equation*}
$$

Now let us consider the insertion of a zero-length path. The rules described in Sec. 4.2 for constructing a term in the MSE allow three possible types of such insertions: (i) the insertion of $g$, (ii) the insertion of $\partial \delta g$, and (iii) the insertion of $\partial g$. More specifically, we are interested in the amplitudes

$$
\begin{align*}
\mathcal{A}_{\mathrm{i}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) & \equiv \int \partial g(\mathbf{x}, \boldsymbol{\alpha}) g(\boldsymbol{\alpha}, \boldsymbol{\beta}) \delta g\left(\boldsymbol{\beta}, \boldsymbol{x}^{\prime}\right)  \tag{5.28}\\
\mathcal{A}_{\mathrm{ii}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) & \equiv \int g(\mathbf{x}, \boldsymbol{\alpha}) \partial \delta g(\boldsymbol{\alpha}, \boldsymbol{\beta}) g\left(\boldsymbol{\beta}, \boldsymbol{x}^{\prime}\right)  \tag{5.29}\\
\mathcal{A}_{\mathrm{iii}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) & \equiv \int \partial g(\mathbf{x}, \boldsymbol{\alpha}) \partial g(\boldsymbol{\alpha}, \boldsymbol{\beta}) g\left(\boldsymbol{\beta}, \boldsymbol{x}^{\prime}\right) \tag{5.30}
\end{align*}
$$

and their leading-order asymptotic contribution due to the critical point $(\boldsymbol{\alpha}, \boldsymbol{\beta})=$ $\left(\boldsymbol{\alpha}_{c}, \boldsymbol{\alpha}_{c}\right)$. Let us start with $\mathcal{A}_{\mathrm{i}}$. By using a coordinate system centered at $\boldsymbol{\alpha}_{c}$ and the short-distance asymptotics of $g$, and expanding $\mathcal{S}$ to second order around $\boldsymbol{\alpha}_{c}$, we find that

$$
\begin{align*}
\mathcal{A}_{\mathbf{i}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)= & \mathcal{O}\left(\frac{k_{\mathrm{F}}^{2}}{L^{2}}\right) \mathrm{e}^{i S_{c}} \int d \sigma_{\alpha} d \sigma_{\beta} \frac{\mathrm{e}^{i k_{F}\left(D\left(x_{\alpha}-x_{\beta}\right)+E\left(y_{\alpha}-y_{\beta}\right)\right)}}{\sqrt{\left(x_{\alpha}-x_{\beta}\right)^{2}+\left(y_{\alpha}-y_{\beta}\right)^{2}}} \\
& \times \exp \left(i k_{\mathrm{F}} L\left(A x_{\alpha}^{2}+B x_{\alpha} y_{\alpha}+C y_{\alpha}^{2}+A^{\prime} x_{\beta}^{2}+B^{\prime} x_{\beta} y_{\beta}+C^{\prime} y_{\beta}^{2}\right)\right) \\
= & \mathcal{O}\left(\frac{k_{\mathrm{F}}^{2}}{L^{2}}\right) \mathcal{O}\left(\frac{1}{k_{\mathrm{F}}}\right) \mathrm{e}^{i S_{c}} \int d x d y \exp \left(i k_{\mathrm{F}} L\left(A^{\prime \prime} x^{2}+B^{\prime \prime} x y+C^{\prime \prime} y^{2}\right)\right) \\
= & \mathcal{O}\left(\frac{1}{L}\right) \mathrm{e}^{i S_{c}}, \tag{5.31}
\end{align*}
$$

where $A, A^{\prime}, A^{\prime \prime}, B, B^{\prime}, B^{\prime \prime}, C, C^{\prime}$ and $C^{\prime \prime}$ are $\mathcal{O}\left(1 / L^{2}\right), D$ and $E$ are $\mathcal{O}(1)$, and $S_{c}=k\left|\mathbf{x}-\boldsymbol{\alpha}_{c}\right|+k^{\prime}\left|\boldsymbol{\alpha}_{c}-\mathbf{x}^{\prime}\right|$. Thus, $\mathcal{A}_{\mathrm{i}}$ contributes at the same order as $\mathcal{A}$. A similar calculation provides an estimate of the order-of-magnitude of $\mathcal{A}_{\mathrm{ii}}$, and show it to be of the same order as $\mathcal{A}$ and $\mathcal{A}_{\mathrm{i}}$. Next, let us consider $\mathcal{A}_{\mathrm{iii}}$. By using the short-distance behavior

$$
\begin{equation*}
\partial g(l) \approx \mathcal{O}\left(\frac{1}{L}\right) \frac{\partial g(l)}{\partial l} l \propto \mathcal{O}\left(\frac{1}{L}\right) g(l), \tag{5.32}
\end{equation*}
$$

the order of magnitude of $\mathcal{A}_{\mathrm{iii}}$ can be deduced from the estimate of $\mathcal{A}_{\mathrm{i}}$ :

$$
\begin{equation*}
\mathcal{A}_{\mathrm{iii}}=\mathcal{O}\left(\frac{1}{k_{\mathrm{F}} L^{2}}\right) \mathrm{e}^{i S_{c}} . \tag{5.33}
\end{equation*}
$$

Thus, we see that $\mathcal{A}_{\mathrm{iii}}$ contributes to the full amplitude only at subleading order.
In the usual case of a Schrödinger billiard with hard walls (i.e. Dirichlet boundary conditions), the asymptotic contribution coming from the critical points of Type 2 and Type 5 have been shown to be smaller than the stationary-phase (i.e. Type 1 critical points) contributions, and by a factor of $k_{\mathrm{F}} L$ [12]. The reason for this is that in the hard wall case all the singularities of $\mathcal{M}$ are due to $\partial g$ 's. Thus, for the Schrödinger billiard with hard walls, as far as the leading-order contribution is concerned the Type 2 and Type 5 critical points are irrelevant. Thus, in such billiards, the leading asymptotic contribution comes from the stationary-phase points.

In constrast with the case of Schrödinger billiards with hard walls, in Andreev billiards (and Schrödinger billiards with soft walls, i.e., with a finite bounding potential) the Type 2 and Type 5 critical points do not necessarily give only subleading-order contributions. More specifically, in Andreev billiards, in addition to its $\partial g$ singularities, $\mathcal{M}$ can have additional singularities due to $g$ and $\partial \delta g$. As shown above, both of these singularities contribute at leading order in the $\left(k_{\mathrm{F}} L\right)^{-1}$ expansion [36]. Thus, Type 2 and Type 5 critical points are relevant for Andreev billiards.

Having determined the significance of Type 5 critical points for Andreev billiards, we now examine these critical points more closely. The first important observation is that the insertion of a zero-length path does not change the value of the phase function $\mathcal{S}$. As all Type 5 critical points can be regarded as originating from Type 1 critical points via the insertion of a suitable number of zero-length paths, the phase of any Type 5 critical point is equal to that of the originating Type 1 critical point. The second important observation is that, because the phase is not changed by the insertion of zero-length paths, all amplitudes originating (via insertions) from a given Type 1 critical point carry a common phase, and therefore add coherently. Thus,
the effect of the Type 5 critical points is to modify (but not necessarily increase) the amplitude of the contribution to the originating Type 1 critical point.

In order to make less abstract the issue discussed in the previous paragraph, consider the example of reflection from an infinite plane boundary in the short-wave asymptotic limit. For the case of the hard Schrödinger billiard there is a single critical point, which is of the stationary-phase type: it is the classical reflection point (for which the angle of incidence equals the angle of reflection). For the case of the Andreev billiard there are two possible electron reflections: electron-to-hole and electron-to-electron. These two processes have differing phases and, correspondingly, differing stationary-phase (i.e. classical reflection) points. If one were to take into account only the stationary-phase (i.e. Type 1) points then one would find that the amplitude for electron-to-hole reflection would vanish, whereas that for electron-toelectron reflection would be of order unity. However, this finding would be misleading, owing to the fact that the set of critical points that contribute at leading order (in the the short-wave asymptotic limit) is much larger. To see this, focus on the case of electron-to-electron reflection. Let us label the classical reflection point by $\boldsymbol{\alpha}_{\mathrm{c}}$. Then set of critical points is $\left(\boldsymbol{\alpha}_{\mathrm{c}}\right),\left(\boldsymbol{\alpha}_{\mathrm{c}}, \boldsymbol{\alpha}_{\mathrm{c}}\right),\left(\boldsymbol{\alpha}_{\mathrm{c}}, \boldsymbol{\alpha}_{\mathrm{c}}, \boldsymbol{\alpha}_{\mathrm{c}}\right),\left(\boldsymbol{\alpha}_{\mathrm{c}}, \boldsymbol{\alpha}_{\mathrm{c}}, \boldsymbol{\alpha}_{\mathrm{c}}, \boldsymbol{\alpha}_{\mathrm{c}}\right)$, etc., i.e., there is the possibility of multiple scatterings from the boundary, all taking place in the vicinity of the classical reflection point. These additional critical points correct the amplitude for this scattering process, and yield the expected result, namely that the net electron-to-electron amplitude is very small. The origin of this correction, then, is that multiple virtual propagations, inside the superconductor but near to the classical reflection point, decrease the amplitude of the electron-to-electron reflection process. Mutatis mutandis, this mechanism increases the electron-to-hole reflection amplitude, leading to the familiar physics of Andreev reflection.

There is a simple physical explanation for this mechanism. Quasiparticle propagation in the bulk of the normal region is not affected by the superconducting surround,
except via those Feynman paths that pass nearby the boundary. In the short-wave asymptotic limit, this occurs near reflection points. Thus, quantum mechanically, there is an effective volume around the boundary in which propagation is modified due to the amplitude for electron-to-hole conversion (and vice versa). Hence, there is an effective volume around the classical reflection points, and in this volume multiple scatterings convert electrons arriving from the interior of the normal region into holes departing for the interior. Classically, the volume for such processes is zero, i.e., the conversion takes place precisely at the reflection point. Thus, zero-length propagation at the boundary is responsible for the electron-hole interconversion aspect of Andreev reflection, whereas the requirement of phase-stationarity, applied to propagation in the interior of the normal region, is responsible for the retro-reflection aspect of Andreev reflection.

### 5.4 Integrating out propagation in the superconducting region: Effective reflection

We now actually evaluate the contribution to the short-wave asymptotic approximation to the Green function that arises from all critical points involving zero-length propagation. In doing this, we collect contributions from Type 5 critical points, and arrive at the expected result that reflection leads to almost complete electron-hole interconversion.

We start with the expression (4.26) for $\boldsymbol{\mu}^{\mathrm{ii}}$ which, for the sake of convenience, we rewrite here along with the explicit form of $\mathbf{K}^{\mathrm{oo}}$ obtained from definition (4.22):

$$
\begin{aligned}
\boldsymbol{\mu}^{\mathrm{ii}} & =-2 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{N}}+4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \mathbf{K}^{\mathrm{oo}} \boldsymbol{\sigma}_{3} \delta \mathbf{G}^{\mathrm{N}}+\left\{-2 \boldsymbol{\sigma}_{3} \partial \mathbf{G}^{\mathrm{N}}+4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \mathbf{K}^{\mathrm{oo}} \boldsymbol{\sigma}_{3} \partial \delta \mathbf{G}^{\mathrm{N}}\right\} \boldsymbol{\mu}^{\mathrm{ii}}, \\
\mathbf{K}^{\mathrm{oo}} & \equiv\left(\mathbf{I}+2 \boldsymbol{\sigma}_{3} \cdot \delta \mathbf{G}^{\mathrm{S}}\right)^{-1}
\end{aligned}
$$

As we have shown in Sec. 5.3, in the short-wave asymptotic limit, critical trajectories
that have been obtained from a Type 1 critical trajectory (i.e. a pure stationary-phase trajectory) by the insertion of zero-length propagations of $\partial \delta \mathbf{G}^{N}$ and $\mathbf{G}^{S}$ contribute at the same order as the original Type 1 critical trajectory. Thus, such contributions should be summed to all orders. On the other hand, critical trajectories obtained from Type 1 critical trajectories by the insertion of $\partial \mathbf{G}^{N}$ and $\delta \mathbf{G}^{S}$ contribute only at sub-leading orders. Thus, it is appropriate to ignore such contributions.

Moreover, we are considering situations in which the range of $G^{S}$ is much smaller than the size of the billiard. Thus, all critical points that include finite-range superconducting propagation are suppressed exponentially (in the size of the billiard), despite their being formally of leading order. We shall therefore neglect them, at least for the time being. Such contributions constitute the single-particle tunneling amplitude through the classically-inaccessible S region. Below, in App. C, we shall study the consequences of relaxing the condition that the range of $\mathbf{G}^{S}$ be much smaller than the size of the billiard. We shall then show that in settings involving convex billiards (i.e. billiards for which all chords lie inside the billiard) such contributions cancel each other at leading asymptotic order. Thus, for the purposes of a leading-order calculation we can make the approximation $\delta \mathbf{G}^{S} \approx 0$. It follows that

$$
\begin{equation*}
\mathbf{K}^{\mathrm{oo}} \approx \mathbf{I} \tag{5.34}
\end{equation*}
$$

The only remaining appearances of the superconducting Green function (i.e. $\mathbf{G}^{S}$ with no derivatives) in the MSE generate critical points having zero-length superconducting propagation. Moreover, the only Green function that contributes at leading order to both zero-length and nonzero-length propagation is $\partial \delta \mathbf{G}^{\mathrm{N}}$. In order to re-sum all possible zero-length propagations it is natural to separate the operator $\partial \delta \mathbf{G}^{\mathrm{N}}$ into two parts: one solely generating zero-length propagation; the other solely generating finite-length propagation [37]. A convenient way to do this is to define the following
operators:

$$
\begin{align*}
\left(\partial \delta \mathbf{G}_{\mathrm{z}}^{\mathrm{N}} \mathbf{F}\right)(\boldsymbol{\alpha}) & \equiv \int_{\partial \mathcal{V}} d \sigma_{\beta} \partial_{\alpha} \partial_{\beta} \mathbf{G}^{\mathrm{N}}(\boldsymbol{\alpha}, \boldsymbol{\beta}) \mathbf{F}(\boldsymbol{\beta}) w(\boldsymbol{\alpha}-\boldsymbol{\beta}),  \tag{5.35}\\
\left(\partial \delta \mathbf{G}_{\mathrm{f}}^{\mathrm{N}} \mathbf{F}\right)(\boldsymbol{\alpha}) & \equiv \int_{\partial \mathcal{V}} d \sigma_{\beta} \partial_{\alpha} \partial_{\beta} \mathbf{G}^{\mathrm{N}}(\boldsymbol{\alpha}, \boldsymbol{\beta}) \mathbf{F}(\boldsymbol{\beta})(1-w(\boldsymbol{\alpha}-\boldsymbol{\beta})), \tag{5.36}
\end{align*}
$$

where $w(\boldsymbol{\alpha}-\boldsymbol{\beta})$ is a smooth function that equals unity whenever $\boldsymbol{\alpha}$ and $\boldsymbol{\beta}$ are close to one another and vanishes whenever $\boldsymbol{\alpha}$ and $\boldsymbol{\beta}$ are far away from one another. The effect of this function is to isolate the critical point at $\boldsymbol{\beta}=\boldsymbol{\alpha}$ from the remaining critical points, the latter having finite-length propagation involving $\partial \delta \mathbf{G}^{\mathrm{N}}$. For the purposes of our asymptotic expansion, the isolating function $w$ is only a convenience, and its particular form does not affect the final results (as long as the range of $w$ precludes its enveloping simultaneously any pairs of critical points). Then the equation for $\boldsymbol{\mu}^{\mathrm{ii}}$ becomes

$$
\begin{align*}
\boldsymbol{\mu}^{\mathrm{ii}} \approx-2 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{N}}+4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \boldsymbol{\sigma}_{3} & \delta \mathbf{G}^{\mathrm{N}}+\left\{4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \boldsymbol{\sigma}_{3} \partial \delta \mathbf{G}_{\mathrm{z}}^{\mathrm{N}}\right. \\
& \left.+\left(-2 \boldsymbol{\sigma}_{3} \partial \mathbf{G}^{\mathrm{N}}+4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \boldsymbol{\sigma}_{3} \partial \delta \mathbf{G}_{\mathrm{f}}^{\mathrm{N}}\right)\right\} \boldsymbol{\mu}^{\mathrm{ii}} . \tag{5.37}
\end{align*}
$$

Having decomposed the kernel of this equation into two pieces (the first consisting of critical points having zero-length propagation and the second consisting of critical points having finite-length propagation) we invert this equation with respect to the former piece, obtaining

$$
\begin{align*}
& \boldsymbol{\mu}^{\mathrm{ii}} \approx\left(\mathbf{I}-4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \boldsymbol{\sigma}_{3} \partial \delta \mathbf{G}_{\mathrm{z}}^{\mathrm{N}}\right)^{-1}\left\{-2 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{N}}+4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \boldsymbol{\sigma}_{3} \delta \mathbf{G}^{\mathrm{N}}\right. \\
&\left.+\left(-2 \boldsymbol{\sigma}_{3} \partial \mathbf{G}^{\mathrm{N}}+4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \boldsymbol{\sigma}_{3} \partial \delta \mathbf{G}_{\mathrm{f}}^{\mathrm{N}}\right) \boldsymbol{\mu}^{\mathrm{ii}}\right\} . \tag{5.38}
\end{align*}
$$

We now define the renormalized Green function $\mathbf{G}^{\mathrm{R}}\left(\boldsymbol{\alpha}, \mathrm{x}^{\prime}\right)$ :

$$
\begin{equation*}
\mathbf{G}^{\mathrm{R}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) \equiv\left(\mathbf{I}-4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \boldsymbol{\sigma}_{3} \partial \delta \mathbf{G}_{\mathbf{z}}^{\mathrm{N}}\right)^{-1}\left\{-\boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{N}}+2 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}} \boldsymbol{\sigma}_{3} \delta \mathbf{G}^{\mathrm{N}}\right\} \tag{5.39}
\end{equation*}
$$

In terms of $\mathbf{G}^{\mathrm{R}}$, Eq. (5.38) becomes

$$
\begin{equation*}
\boldsymbol{\mu}^{\mathrm{ii}} \approx 2 \mathbf{G}^{\mathrm{R}}+2 \partial \mathbf{G}^{\mathrm{R}} \boldsymbol{\mu}^{\mathrm{ii}} \tag{5.40}
\end{equation*}
$$

We note that this equation, no critical points containing zero-length propagation contribute to $\boldsymbol{\mu}^{\mathrm{ii}}$ at leading order. Thus, the summation of short-range critical orbits is achieved via the calculation of $\mathrm{G}^{\mathrm{R}}$ in the short-wave asymptotic approximation.

The main contributions to the integrals implied in Eq. (5.39) come from the neighborhood of $\boldsymbol{\alpha}$. This follows from: (i) the fact that isolating function is short-ranged; (i) the fact that $\mathrm{G}^{\mathrm{S}}$ is finite-ranged; and (iii) the assumption that the billiard is large enough to exponentially suppress any finite-range critical points produced by $\mathbf{G}^{S}$. Therefore, it is adequate to approximate the boundary surface $\partial \mathcal{V}$ around $\boldsymbol{\alpha}$. The lowest-order approximation to $\partial \mathcal{V}$ is the tangent plane at $\boldsymbol{\alpha}$. The corrections to this approximation are smaller by a factor of $\left(k_{\mathrm{F}} R\right)^{-1}$, where $R$ is the radius of curvature at the point $\boldsymbol{\alpha}$. Throughout this thesis we are assuming that the surface $\partial \mathcal{V}$ is sufficiently smooth that $R$ is of order $L$, i.e., the radius of curvature is comparable to the billiard size. Thus, corrections due to the curvature of the surface do not contribute at leading order. Having replaced $\partial \mathcal{V}$ by a tangent plane, the integral equation for $\mathbf{G}^{\mathrm{R}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right)$ becomes solvable, owing to the resulting translational invariance in all directions parallel to the tangent plane. Thus, by introducing the two-dimensional Fourier transform (2DFT; see App. B) of all Green functions appearing in the integral equation (5.39), it is straightforward to obtain the following algebraic result for the 2DFT of the renormalized Green function:

$$
\begin{align*}
\mathbf{G}^{\mathrm{R}}\left(p, z^{\prime}\right)=\{\mathbf{I}-4 & \left.\boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}}(p) \boldsymbol{\sigma}_{3} \partial \delta \mathbf{G}^{\mathrm{N}}(p)\right\}^{-1} \\
& \times\left\{-\boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{N}}\left(p, z^{\prime}\right)+2 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}}(p) \boldsymbol{\sigma}_{3} \delta \mathbf{G}^{\mathrm{N}}\left(p, z^{\prime}\right)\right\}, \tag{5.41}
\end{align*}
$$

to which there are corrections of order $\left(k_{\mathrm{F}} R\right)^{-1}$. (This Grenn function is exactly the Green function for a planar boundary.) Here and elsewhere, $p$ denotes the magnitude of the 2 D vector $\mathbf{p}$ conjugate to the position-vector in the tangent plane.

We now embark on the task of inverting the $2 \mathrm{DFT} \mathbf{G}^{\mathrm{R}}\left(p, z^{\prime}\right)$ in order to obtain the approximate real-space renormalized Green function $\mathbf{G}^{\mathrm{R}}(\boldsymbol{\alpha}, \mathbf{x})$. Thus, we need to
evaluate the integral

$$
\begin{equation*}
\mathbf{G}^{\mathrm{R}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right)=\int \frac{d^{2} p}{(2 \pi)^{2}} \mathbf{G}^{\mathrm{R}}\left(p, z^{\prime}\right) \exp i \mathbf{p} \cdot(\boldsymbol{\alpha}-\boldsymbol{\beta}) \tag{5.42}
\end{equation*}
$$

where $\boldsymbol{\beta}$ is the component of $\mathbf{x}$ parallel to the plane and $z^{\prime}$ is the perpendicular component. The quantities $\mathbf{G}^{\mathrm{N}}\left(p, z^{\prime}\right)$ and $\delta \mathbf{G}^{\mathrm{N}}\left(p, z^{\prime}\right)$, needed to construct $\mathbf{G}^{\mathrm{R}}\left(p, z^{\prime}\right)$, are derived in App. B, where they are found to be given by

$$
\begin{align*}
\mathbf{G}^{\mathrm{N}}\left(p, z^{\prime}\right) & =\left(\begin{array}{cc}
\frac{1}{2 a_{+}(p)} \mathrm{e}^{-a_{+}(p)\left|z^{\prime}\right|} & 0 \\
0 & -\frac{1}{2 a_{-}(p)} \mathrm{e}^{-a_{-}(p)\left|z^{\prime}\right|}
\end{array}\right) \\
& =\left(\begin{array}{cc}
\frac{1}{2 a_{+}(p)} & 0 \\
0 & -\frac{1}{2 a_{-}(p)}
\end{array}\right)\left(\begin{array}{cc}
\mathrm{e}^{-a_{+}(p)\left|z^{\prime}\right|} & 0 \\
0 & \mathrm{e}^{-a_{-}(p)\left|z^{\prime}\right|}
\end{array}\right)  \tag{5.43}\\
\delta \mathbf{G}^{\mathrm{N}}\left(p, z^{\prime}\right) & =\left(\begin{array}{cc}
\frac{1}{2} \mathrm{e}^{-a_{+}(p)\left|z^{\prime}\right|} & 0 \\
0 & -\frac{1}{2} \mathrm{e}^{-a_{-}(p)\left|z^{\prime}\right|}
\end{array}\right)  \tag{5.44}\\
& =\left(\begin{array}{cc}
\frac{1}{2} & 0 \\
0 & -\frac{1}{2}
\end{array}\right)\left(\begin{array}{cc}
\mathrm{e}^{-a_{+}(p)\left|z^{\prime}\right|} & 0 \\
0 & \mathrm{e}^{-a_{-}(p)\left|z^{\prime}\right|}
\end{array}\right) \tag{5.45}
\end{align*}
$$

in which $a_{ \pm}(p) \equiv \sqrt{p^{2}-k_{ \pm}^{2}}$, the square roots being taken such that their real parts are always positive. Note that the only $z^{\prime}$ dependence in $\mathbf{G}^{\mathrm{R}}\left(p, z^{\prime}\right)$ comes from $\mathbf{G}^{\mathrm{N}}\left(p, z^{\prime}\right)$ and $\delta \mathbf{G}^{\mathrm{N}}\left(p, z^{\prime}\right)$, and it is only in these terms that there is exponential dependence on p. By inserting Eqs. (5.43) and (5.45) into Eq. (5.41) we find that $\mathbf{G}^{\mathrm{R}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right)$ is given by

$$
\mathbf{G}^{\mathrm{R}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right)=\int \frac{d^{2} p}{(2 \pi)^{2}} \mathbf{R}(p)\left(\begin{array}{cc}
\frac{1}{2 a_{+}(p)} & \mathrm{e}^{-a_{+}(p)\left|z^{\prime}\right|} \tag{5.46}
\end{array} 0^{0} \quad-\frac{1}{2 a_{-}(p)} \mathrm{e}^{-a_{-}(p)\left|z^{\prime}\right|}\right) \exp i \mathbf{p} \cdot(\boldsymbol{\alpha}-\boldsymbol{\beta})
$$

Here, $\mathbf{R}(p)$ is a certain algebraic function of $p$, and is defined by

$$
\begin{align*}
& \mathbf{R}(p) \equiv\left\{\mathbf{I}-4 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathbf{S}}(p) \boldsymbol{\sigma}_{3} \partial \delta \mathbf{G}^{\mathrm{N}}(p)\right\}^{-1} \\
& \times\left\{\left(\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right)+2 \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}}(p)\left(\begin{array}{cc}
a_{+}(p) & 0 \\
0 & -a_{-}(p)
\end{array}\right)\right\} \tag{5.47}
\end{align*}
$$

The analytic expression for $\mathbf{R}(p)$ is obtained using the 2 DFTs of $\mathbf{G}^{\mathbf{S}}(p)$ and $\partial \delta \mathbf{G}^{\mathrm{N}}(p)$ :

$$
\mathbf{G}^{\mathrm{S}}(p)=\frac{1}{2}\left\{\frac{E}{i \sqrt{\Delta^{2}-E^{2}}} \mathbf{I}-\frac{\Delta}{i \sqrt{\Delta^{2}-E^{2}}} \boldsymbol{\sigma}_{2}\right\}\left(\frac{1}{2 a_{+}^{\mathrm{S}}(p)}-\frac{1}{2 a_{-}^{\mathrm{S}}(p)}\right)
$$

$$
\begin{array}{r}
+\frac{1}{2} \boldsymbol{\sigma}_{3}\left(\frac{1}{2 a_{+}^{\mathrm{S}}(p)}+\frac{1}{2 a_{-}^{\mathrm{S}}(p)}\right), \\
\partial \delta \mathbf{G}^{\mathrm{N}}(p)=-\frac{1}{4} \mathbf{I}\left(a_{+}(p)-a_{-}(p)\right)-\frac{1}{4} \boldsymbol{\sigma}_{3}\left(a_{+}(p)+a_{-}(p)\right) . \tag{5.49}
\end{array}
$$

We evaluate the integral in Eq. (5.46) via the method of stationary phase, which becomes exact in the limit $k_{\mathrm{F}}\left|\boldsymbol{\alpha}-\mathbf{x}^{\prime}\right| \rightarrow \infty$. From the form of $a_{ \pm}(p)$ we see that for values of $p$ for which $a_{ \pm}(p)$ is essentially imaginary (i.e. for $p<\operatorname{Ree} k_{ \pm}$) there exists the possibility of a stationary-phase point. Note, however, that $a_{+}$behaves differently from $a_{-}$, due to the fact that the imaginary parts of $k_{ \pm}^{2}$ have opposing signs. For $p<\operatorname{Re} k_{-}$(note that $k_{-}$is always smaller than $k_{+}$) we have

$$
\begin{equation*}
a_{ \pm}(p)=\mp i \sqrt{k_{ \pm}^{2}-p^{2}} \tag{5.50}
\end{equation*}
$$

The stationary-phase point is defined by the condition

$$
\begin{equation*}
\frac{\partial}{\partial \mathbf{p}}\left(\mathbf{p} \cdot(\boldsymbol{\alpha}-\boldsymbol{\beta}) \pm z^{\prime} \sqrt{k_{ \pm}^{2}-|\mathbf{p}|^{2}}\right)=\mathbf{0} \tag{5.51}
\end{equation*}
$$

from which we see that stationary-phase point $\mathbf{p}_{\mathrm{c}}$ satisfies

$$
\begin{equation*}
\frac{(\boldsymbol{\alpha}-\boldsymbol{\beta})}{z^{\prime}}= \pm \frac{\mathbf{p}_{\mathrm{c}}}{\sqrt{k_{ \pm}^{2}-\left|\mathbf{p}_{\mathrm{c}}\right|^{2}}} \tag{5.52}
\end{equation*}
$$

and that $\left|\mathbf{p}_{\mathrm{c}}\right|^{2}$ has the value $k_{ \pm}^{2} \sin ^{2} \theta_{\alpha x^{\prime}}$, where $\theta_{\boldsymbol{\alpha} x^{\prime}}$ is defined to be the angle between the normal vector at the surface point $\boldsymbol{\alpha}$ and the vector $\mathrm{x}^{\prime}-\boldsymbol{\alpha}$. Then the effective Green function can be asymptotically approximated as

$$
\begin{align*}
\mathbf{G}^{\mathrm{R}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) \approx & \left(\begin{array}{cc}
R_{++}\left(k_{+}^{2} \sin ^{2} \theta_{\alpha x^{\prime}}\right) & R_{+-}\left(k_{-}^{2} \sin ^{2} \theta_{\alpha x^{\prime}}\right) \\
R_{-+}\left(k_{+}^{2} \sin ^{2} \theta_{\alpha x^{\prime}}\right) & R_{--}\left(k_{-}^{2} \sin ^{2} \theta_{\alpha x^{\prime}}\right)
\end{array}\right) \\
& \times \int \frac{d^{2} p}{(2 \pi)^{2}}\left(\begin{array}{cc}
\frac{\mathrm{e}^{-a_{+}(p)\left|z^{\prime}\right|}}{2 a_{+}(p)} & 0 \\
0 & -\frac{\mathrm{e}^{-a_{-}(p)\left|z^{\prime}\right|}}{2 a_{-}(p)}
\end{array}\right) \mathrm{e}^{i \mathbf{p} \cdot(\boldsymbol{\alpha}-\boldsymbol{\beta})} \\
= & \left(\begin{array}{cc}
R_{++} & R_{+-} \\
R_{-+} & R_{--}
\end{array}\right)\left(\begin{array}{cc}
g_{+}\left(\boldsymbol{\alpha}-\mathbf{x}^{\prime}\right) & 0 \\
0 & -g_{-}\left(\boldsymbol{\alpha}-\mathbf{x}^{\prime}\right)
\end{array}\right) \tag{5.53}
\end{align*}
$$

The second line is obtained by noting that the integral is, in fact, the 2 DFT of $\mathbf{G}^{\mathrm{N}}$ (see App. B). The amplitudes $R_{++}, R_{+-}, R_{-+}$and $R_{--}$can now be respectively interpreted as the electron-electron, electron-hole, hole-electron and hole-hole reflection
amplitudes, and can be obtained from Eq. (5.47). These amplitudes are, in general, nonvanishing. However the charge-preserving amplitudes (i.e. $R_{++}$and $R_{--}$) are smaller than the charge interconverting amplitudes (i.e. $R_{+-}$and $R_{-+}$) by a factor of $\Delta / \mu \cos ^{2} \theta$. In order to evaluate $\mathbf{R}$ to leading order in $\Delta / \mu \cos ^{2} \theta$, we first note that for $\theta_{\alpha x^{\prime}}$ not near $\pi / 2$ (i.e. not near grazing) one has the following approximations for $a(p)$ :

$$
\begin{align*}
a_{ \pm}\left(k_{\mathrm{F}} \sin \theta_{\alpha x^{\prime}}\right) & =\mp i k_{\mathrm{F}} \cos \theta_{\alpha x^{\prime}}+\mathcal{O}(E / \mu),  \tag{5.54}\\
a_{ \pm}^{\mathrm{S}}\left(k_{\mathrm{F}} \sin \theta_{\alpha x^{\prime}}\right) & =\mp i k_{\mathrm{F}} \cos \theta_{\alpha x^{\prime}}+\mathcal{O}\left(\Delta_{0} / \mu\right) . \tag{5.55}
\end{align*}
$$

By applying these approximations to Eqs. (5.48-5.49) we find

$$
\begin{align*}
\mathbf{G}^{\mathrm{S}}\left(k_{\mathrm{F}} \sin \theta_{\alpha x^{\prime}}\right) & \simeq\left(\frac{E}{2 i \sqrt{\Delta_{0}^{2}-E^{2}}} \mathbf{I}-\frac{\Delta_{0}}{2 i \sqrt{\Delta_{0}^{2}-E^{2}}} \boldsymbol{\sigma}_{2}\right) \frac{1}{-i k_{\mathrm{F}} \cos \theta_{\alpha x^{\prime}}}  \tag{5.56}\\
\partial \delta \mathbf{G}^{\mathrm{N}}\left(k_{\mathrm{F}} \sin \theta_{\alpha x^{\prime}}\right) & \simeq \frac{1}{2} \mathbf{I} i k_{\mathrm{F}} \cos \theta_{\alpha x^{\prime}} \tag{5.57}
\end{align*}
$$

By using these expressions in Eq. (5.47) we obtain $\mathbf{R}\left(k_{\mathrm{F}} \sin \theta_{\alpha x^{\prime}}\right) \approx \mathrm{e}^{-i \varphi} \boldsymbol{\sigma}_{1}$ and, thus, from Eq. (5.53) we obtain

$$
\mathbf{G}^{\mathrm{R}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) \approx \exp \left(-i \varphi+i \frac{\pi}{2}\right)\left(\begin{array}{cc}
0 & -g_{-}^{\mathrm{N}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right)  \tag{5.58}\\
g_{+}^{\mathrm{N}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) & 0
\end{array}\right) .
$$

The off-diagonal structure represents the total electron-hole interconversion that occurs for large perpendicular momenta. However, strictly speaking, electron-hole interconversion is not perfect, i.e., the amplitudes $R_{++}$and $R_{--}$do not vanish. Moreover, these charge-preserving amplitudes increase, as the angle of incidence approaches $\pi / 2$. Note the difference between the regimes of validity for the approximate expressions Eq. (5.53) and Eq. (5.58): the former becomes valid for $k_{\mathrm{F}} L \gg 1$ whereas the latter becomes valid for $k_{\mathrm{F}} L \gg 1$ and $\Delta / k_{\mathrm{F}}^{2} \ll 1$. However, in either case the MSE can be cast into the following effective Multiple Reflection Expansion:

$$
\begin{gather*}
\mathbf{G}^{\mathrm{ii}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\mathbf{G}^{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+2 \int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} \mathbf{G}^{\mathrm{N}} \mathbf{G}^{\mathrm{R}}+4 \int_{\partial \mathcal{V}} d \sigma_{\alpha} d \sigma_{\beta} \partial_{\alpha} \mathbf{G}^{\mathrm{N}} \partial_{\beta} \mathbf{G}^{\mathrm{R}} \mathbf{G}^{\mathrm{R}} \\
+8 \int_{\partial \mathcal{V}} d \sigma_{\alpha} d \sigma_{\beta} d \sigma_{\gamma} \partial_{\alpha} \mathbf{G}^{\mathrm{N}} \partial_{\beta} \mathbf{G}^{\mathrm{R}} \partial_{\gamma} \mathbf{G}^{\mathrm{R}} \mathbf{G}^{\mathrm{R}}+\cdots \tag{5.59}
\end{gather*}
$$

We stress two points about Eq. (5.59): (i) it is free of leading-order short-range critical points [i.e. the contributions of short-range critical points to $\mathbf{G}^{\mathrm{ii}}$ are smaller, by at least a factor of $\mathcal{O}\left(1 / k_{\mathrm{F}} R\right)$, than the leading-order contribution, and this is what we were aiming for]; and (ii) all propagations inside the superconductor have been integrated out, leading to the effective reflection expansion for $\mathbf{G}^{\mathrm{ii}}$.

## Chapter 6

## Density of states oscillations

In Sec. 5.4 we integrated out superconducting propagation by evaluating the shortrange critical points in the MSE and, hence, we obtained an effective expansion for the Green function, which we have termed an MRE. In the present Section we shall focus on the density of states $\rho(E)$, expressing this quantity in terms of the Green function which, in turn, we express via the MRE. In this section we shall ignore the effects of normal reflection, returning to them only in Sec. 6.2.4. Thus, by using the results of the previous section, we have

$$
\begin{align*}
& \rho(E) \approx \frac{1}{\pi} \int d^{d} x \operatorname{Im}\left\{g_{+}^{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+g_{-}^{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+4 \int_{\partial \mathcal{V}} \partial g_{+}^{\mathrm{N}} \partial g_{-}^{\mathrm{R}} g_{+}^{\mathrm{R}}\right. \\
&+4 \int_{\partial \mathcal{V}} \partial g_{-}^{\mathrm{N}} \partial g_{+}^{\mathrm{R}} g_{-}^{\mathrm{R}} \cdots+2^{2 n} \int_{\partial \mathcal{V}}\left(\partial g_{+}^{\mathrm{N}} \partial g_{-}^{\mathrm{R}} \cdots \partial g_{-}^{\mathrm{R}} g_{+}^{\mathrm{R}}\right) \\
&\left.\quad+2^{2 n} \int_{\partial \mathcal{V}}\left(\partial g_{-}^{\mathrm{N}} \partial g_{+}^{\mathrm{R}} \cdots \partial g_{+}^{\mathrm{R}} g_{-}^{\mathrm{R}}\right)+\cdots\right\}_{\mathbf{x}^{\prime}=\mathbf{x}}, \tag{6.1}
\end{align*}
$$

where terms with odd numbers of reflections vanish, owing to the off-diagonal structure of $\mathbf{G}^{\mathrm{R}}$. First, let us note that the first two terms on the right hand side of Eq. (6.1), which have zero-length propagation and hence vanishing action, do not introduce any oscillations into the DOS. As these terms do not involve any surface integrals (and, hence, do not involve any surface effects), they produce the bulk DOS
of a homogeneous N region [38]:

$$
\begin{equation*}
\left.\frac{1}{\pi} \int_{\mathcal{V}} d^{d} x \operatorname{Im} g_{ \pm}^{\mathrm{N}}\left(\boldsymbol{x}-\boldsymbol{x}^{\prime}\right)\right|_{x^{\prime}=x}=\frac{\mathcal{S}^{d-1}}{2(2 \pi)^{d}} V k_{ \pm}^{d-2} \tag{6.2}
\end{equation*}
$$

where $\mathcal{S}^{d-1}$ is the $(d-1)$-dimensional surface area of a $d$-dimensional unit sphere and $V$ is the volume of $\mathcal{V}$. We now deal with the (remaining) critical points that contribute at leading order. These critical points are closed classical trajectories consisting of the propagation of quasiparticles through the bulk of the billiard (i.e. the N region), connected by reflections from the billiard boundary (i.e. the N-S interface).

We shall distinguish between two asymptotic approximation schemes for $\rho$, both of which are obtained by evaluating the integrals in Eq. (6.1) within the stationaryphase approximation, valid for for large $k_{\mathrm{F}} L$ and small $\Delta / k_{\mathrm{F}}^{2}$. From a technical point of view, the difference between the two schemes concerns the stationary-phase points they use, which must be in accordance with the particular limits assumed for the parameters $k_{\mathrm{F}} L$ and $\Delta / k_{\mathrm{F}}^{2}$ (which the approximation becomes exact).

Scheme A: The first scheme is, in essence, equivalent to the (by now conventional) adiabatic approximation to the wave function, first introduced by Andreev [4]; it becomes exact when energy-level spacing goes to zero, which occurs for the following limit:

$$
\begin{equation*}
k_{\mathrm{F}} L \rightarrow \infty, \quad \Delta / k_{\mathrm{F}}^{2} \rightarrow 0, \quad \text { and } \quad L \Delta / k_{\mathrm{F}} \rightarrow \text { constant. } \tag{6.3}
\end{equation*}
$$

In this scheme, an excitation undergoes perfect retro-reflection (i.e. perfect velocityreversal) because in this limit the difference between $k_{+}$and $k_{-}$is ignored in the calculation of the critical trajectories. The resulting classical dynamics is confined to the chords of the billiard and, thus, is integrable, regardless of the shape of the billiard. However, for finite values of the parameters (i.e. large but finite $k_{\mathrm{F}} L$ and small but finite $\Delta / k_{\mathrm{F}}^{2}$ ), Scheme A produces the locally-energy-averaged DOS, which becomes numerically accurate only around the DOS singularities that it correctly captures. However, it fails to capture the DOS oscillations arising from the confinement of
quasiparticles to the billiard. To capture these oscillations is the main motivation of the following scheme.

Scheme B: In this scheme we shall take into account the previously neglected difference in electron and hole wavevectors. This leads to imperfectness in retro-reflection because, upon reflection, the wavelengths of the incoming and outgoing waves are no longer identical (as happens with refraction except, of course, that the waves are now on the same side of the boundary). Technically speaking, approximation scheme B becomes exact when

$$
\begin{equation*}
k_{\mathrm{F}} L \rightarrow \infty, \quad \text { and } \quad \Delta / k_{\mathrm{F}}^{2} \quad \text { is a small parameter. } \tag{6.4}
\end{equation*}
$$

Now the classical trajectories are determined by the reflection rule given by Eq. (5.6) and, consequently, the dynamics is no longer a priori integrable; on the contrary, it is weakly chaotic for most billiard shapes [39].

In order to understand the distinction between asymptotic schemes $A$ and $B$, consider the phase function $S$ for a process having $2 n$ reflections:

$$
\begin{equation*}
S\left(\boldsymbol{x}, \boldsymbol{\alpha}_{1}, \cdots, \boldsymbol{\alpha}_{2 n}\right)=k_{+} \ell_{x, \alpha_{1}}-k_{-} \ell_{\alpha_{1}, \alpha_{2}}+\cdots-k_{-} \ell_{\alpha_{2 n-1}, \alpha_{2 n}}+k_{+} \ell_{\alpha_{2 n}, x}, \tag{6.5}
\end{equation*}
$$

where $\ell_{\alpha_{i}, \alpha_{i}+1} \equiv\left|\boldsymbol{\alpha}_{i}-\boldsymbol{\alpha}_{i+1}\right|$. Notice that $S$ can be separated into two parts, a large one $S_{\mathrm{And}}$ and a small one $S_{\mathrm{imp}}$ so that $S=S_{\mathrm{And}}+S_{\mathrm{imp}}$, where

$$
\begin{align*}
S_{\mathrm{And}}\left(\boldsymbol{x}, \boldsymbol{\alpha}_{1}, \cdots, \boldsymbol{\alpha}_{2 n}\right) & \equiv \frac{k_{+}+k_{-}}{2}\left(\ell_{x, \alpha_{1}}-\ell_{\alpha_{1}, \alpha_{2}}+\cdots-\ell_{\alpha_{2 n-1}, \alpha_{2 n}}+\ell_{\alpha_{2 n}, x}\right)  \tag{6.6}\\
S_{\mathrm{imp}}\left(\boldsymbol{x}, \boldsymbol{\alpha}_{1}, \cdots, \boldsymbol{\alpha}_{2 n}\right) & \equiv \frac{k_{+}-k_{-}}{2}\left(\ell_{x, \alpha_{1}}+\ell_{\alpha_{1}, \alpha_{2}}+\cdots+\ell_{\alpha_{2 n-1}, \alpha_{2 n}}+\ell_{\alpha_{2 n}, x}\right) \tag{6.7}
\end{align*}
$$

For $n$ not too large (i.e. $\left.n \ll k_{F} L\right), S_{\text {And }}=\mathcal{O}\left(k_{F} L\right)$ and $S_{\text {imp }}=\mathcal{O}\left(L \Delta / k_{F}\right)$. Although it is clear that $S_{\text {And }}$ is large and must be included in any stationary phase calculation, whether or not $S_{\text {imp }}$ should be included in a stationary phase calculation depends precisely on the nature of the limiting scheme. If one solely uses $S_{\text {And }}$ for the calculation of stationary-phase points then the critical trajectories feature retro-reflection; this is
the content of Scheme A. On the other hand, with the inclusion of $S_{\mathrm{imp}}$ in the calculation of stationary-phase points, the critical trajectories feature small deviations from retro-reflection, due to the difference in electron and hole wavevectors. (Note that $S_{\text {imp }}$ is proportional to $k_{+}-k_{-.}$) This, in turn, is Scheme B. For larger $n$ (and thus higher resolution contributions to the DOS), $S_{\text {imp }}$ becomes larger, so that Scheme B should be used [40]. These higher-resolution contributions show up in the DOS as mesoscale oscillations due to the superconducting confinement of quasiparticles.

### 6.1 Scheme A: Andreev approximation

In the present section we shall evaluate Eq. (6.1) for the DOS within the stationaryphase approximation via asymptotic Scheme A. In this scheme, the stationary-phase points (i.e. the closed classical trajectories) are obtained by making stationary the phase $S_{\text {And }}$ alone. The factor $\exp i S_{\mathrm{imp}}$ is considered to be slowly-varying, and thus is evaluated at the critical points determined from $S_{\text {And }}$ alone. In all other factors, the difference between $k_{+}$and $k_{-}$can be neglected. The reflection rule can be obtained from Eq. (5.6) by letting $k_{+}, k_{-} \rightarrow k_{\mathrm{F}}$. In this limit, velocity vectors are, upon reflection, exactly reversed. The classical trajectories obtained by this reflection rule are tracings of the chords of the billiard. This allows us to label every classical trajectory by two boundary points, along with the number of reflections (or tracings). Therefore, the closed classical trajectories of Scheme A with $n$ tracings are degenerate (in the sense that their phases $S_{\text {And }}$ are identical) and they belong to a $(2 d-2)$ parameter family.

Let us first explore the underlying physics of Scheme A. In conventional billiards, degeneracy of closed trajectories is usually related to the symmetry of the billiard (e.g. rotational symmetry of a circular billiard). However, in Andreev billiards this degeneracy is due to the underlying electron-hole symmetry at the Fermi surface, and
is broken explicitly at nonzero energies by the term $S_{\mathrm{imp}}$ [41]. This (approximate) non-geometric symmetry is the reason for the (approximate) integrability of Andreev billiards, whatever the shape of the billiard.

The method for evaluating the DOS is as follows. We fix two reflection points (so that the degeneracy is lifted), evaluate the rest of the (surface and volume) integrals in the stationary-phase approximation, and then evaluate the remaining two surface integrals. Without loss of generality, we choose the first and last reflection points as those to be fixed, and label them $\boldsymbol{\alpha}$ and $\boldsymbol{\beta}$, respectively. We focus on the contribution to the DOS from the term having $2 n+2$ reflections:

$$
\begin{gather*}
\pi \rho_{2 n}^{ \pm}(E) \equiv \operatorname{Im} \int d \sigma_{\alpha} d \sigma_{\beta} d \sigma_{\alpha_{1}} \cdots d \sigma_{\alpha_{2 n}} \partial g_{\mp}\left(\boldsymbol{\alpha}, \boldsymbol{\alpha}_{1}\right) \partial g_{ \pm}\left(\boldsymbol{\alpha}_{1}, \boldsymbol{\alpha}_{2}\right) \cdots \partial g_{\mp}\left(\boldsymbol{\alpha}_{2 n}, \boldsymbol{\beta}\right) \\
\times 2^{2 n+2} \mathrm{e}^{2 n i \varphi} \int_{\mathcal{V}} d^{d} x g_{ \pm}(\boldsymbol{\beta}, \mathbf{x}) \partial g_{ \pm}(\mathbf{x}, \boldsymbol{\alpha}) \tag{6.8}
\end{gather*}
$$

where the $+/-$ sign represents the contribution from the electron/hole sector.


Figure 6.1: Geometry for $\boldsymbol{x}$ integration in Eq. (6.10)

We first evaluate the $\mathbf{x}$ integral. The stationary-phase points of the $\mathbf{x}$ integration lie on the line joining $\boldsymbol{\alpha}$ to $\boldsymbol{\beta}$. Because the phase does not vary as one moves the
point $\mathbf{x}$ along this line, it is natural to separate the $\mathbf{x}$ integration into longitudinal (i.e. parallel to $\boldsymbol{\alpha}-\boldsymbol{\beta}$ ) and transverse (i.e. perpendicular to $\boldsymbol{\alpha}-\boldsymbol{\beta}$ ) components; see Fig. 6.1 for the nomenclature and geometry). We use the asymptotic expression (5.9) for the homogeneous Green functions at large argument, together with the asymptotic expression for $\partial g$, i.e.,

$$
\begin{equation*}
2 \partial g_{ \pm}(\boldsymbol{\beta}, \boldsymbol{\alpha}) \approx\left(\frac{k_{ \pm}}{2 \pi \ell_{\alpha \beta}}\right)^{\frac{d-1}{2}} \cos \theta_{\alpha \beta} \exp \left( \pm i k_{ \pm} \ell_{\alpha \beta} \mp i \pi(d-1) / 4\right) \tag{6.9}
\end{equation*}
$$

where $\theta_{\alpha \beta}$ is the angle between the ray joining $\boldsymbol{\alpha}$ to $\boldsymbol{\beta}$ and the normal direction at the point $\boldsymbol{\alpha}$. Then, we fix $l$, evaluate the $\mathbf{t}$ integral in the stationary-phase approximation, and evaluate the $l$ integral:

$$
\begin{align*}
\int_{\mathcal{V}} d^{d} x g_{ \pm}(\boldsymbol{\beta}, \boldsymbol{x}) \partial g_{ \pm}(\boldsymbol{x}, \boldsymbol{\alpha}) \approx & \pm i \frac{k_{ \pm}^{d-2} \mathrm{e}^{\mp i \pi(d-1) / 2}}{2^{d+1} \pi^{d-1}} \cos \theta_{\alpha \beta} \int d l\left(l\left(\ell_{\alpha \beta}-l\right)\right)^{\frac{1-d}{2}} \\
& \times \int d^{d-1} t \mathrm{e}^{ \pm i k_{ \pm}\left(\sqrt{l^{2}+|t|^{2}}+\sqrt{\left(\ell_{\alpha \beta}-l\right)^{2}+|t|^{2}}\right)} \\
\approx & \pm i \frac{k_{ \pm}^{d-2} \mathrm{e}^{\mp i \pi(d-1) / 2}}{2^{d+1} \pi^{d-1}} \cos \theta_{\alpha \beta} \int d l\left(l\left(\ell_{\alpha \beta}-l\right)\right)^{\frac{1-d}{2}} \\
& \times \int d^{d-1} t \mathrm{e}^{ \pm i k_{ \pm} \ell_{\alpha \beta}\left(1+\frac{|t|^{2}}{2 l\left(\ell_{\alpha \beta}-l\right)}\right)} \\
= & \frac{ \pm i \ell_{\alpha \beta} \cos \theta_{\alpha \beta}}{4 k_{ \pm}}\left(\frac{k_{ \pm}}{2 \pi \ell_{\alpha \beta}}\right)^{\frac{(d-1)}{2}} \mathrm{e}^{ \pm i k_{ \pm} \ell_{\alpha \beta} \mp i \pi(d-1) / 4} . \tag{6.10}
\end{align*}
$$

Next, we use Eqs. (6.9) and (6.10), together with Eq. (6.8), to obtain the following asymptotic expression for $\rho_{2 n}^{ \pm}(E)$ :

$$
\begin{align*}
\rho_{2 n}^{ \pm}(E) \approx & \operatorname{Re} \int d \sigma_{\alpha} d \sigma_{\beta}\left(\cos \theta_{\alpha \beta} \cos \theta_{\beta \alpha}\right)^{n} I_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\beta} ; k_{\mathrm{F}}\right)\left(\frac{k_{\mathrm{F}}}{2 \pi \ell_{\alpha \beta}}\right)^{n d-n-1} \\
& \times \exp \left(-i 2 n \varphi+i n\left(k_{+}-k_{-}\right) \ell_{\alpha \beta}\right)  \tag{6.11}\\
I_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\beta} ; k_{\mathrm{F}}\right) \equiv & \int_{\mathrm{sp}} d \sigma_{\alpha_{1}} \cdots d \sigma_{\alpha_{2 n-2}} \mathrm{e}^{ \pm i k_{\mathrm{F}} S_{\mathrm{And}}} \\
S_{\mathrm{And}} & \equiv \ell_{\beta \alpha}-\ell_{\alpha \alpha_{1}}+\ell_{\alpha_{1} \alpha_{2}} \cdots-\ell_{\alpha_{2 n-2} \beta} \tag{6.12}
\end{align*}
$$

and $\int_{\text {sp }}$ indicates that the surface integrals over $\boldsymbol{\alpha}_{1} \cdots \boldsymbol{\alpha}_{2 n-2}$ should be evaluated in the stationary phase approximation. In order to do this evaluation, we expand the functions $\left\{\ell_{\alpha_{i} \alpha_{j}}\right\}$ around $\left(\boldsymbol{\alpha}_{2 i}, \boldsymbol{\alpha}_{2 i+1}\right)=(\boldsymbol{\alpha}, \boldsymbol{\beta})$.


Figure 6.2: Geometry for fluctuations of the reflection points
We now focus on the evaluation of the stationary phase integrals in the definition of $I_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\beta} ; k_{\mathrm{F}}\right)$. In the remainder of this section we shall work in three dimensions; after going through the calculation for this $d=3$ case, the extension to higher dimensions is straightforward. The coordinate system for evaluating these stationary phase integrals is as follows (see Fig. 6.2). We use a pair of coordinate systems: one for the set $\left\{\boldsymbol{\alpha}_{2 i}\right\}$ (which are near $\boldsymbol{\alpha}$ ); the other for the set $\left\{\boldsymbol{\alpha}_{2 i+1}\right\}$ (which are near $\boldsymbol{\beta}$.) Each of these coordinate systems is constructed from the normal vector at the point in question, along with two unit vectors in the tangent plane at this point. Thus, the coordinate system doublet comprises two sets: one formed by the unit vectors $\left\{\hat{\mathbf{x}}_{\alpha}, \hat{\mathbf{y}}_{\alpha}, \hat{\mathbf{n}}_{\alpha}\right\}$; the other formed by the unit vectors $\left\{\hat{\mathbf{x}}_{\beta}, \hat{\mathbf{y}}_{\beta}, \hat{\mathbf{n}}_{\beta}\right\}$, where $\hat{\mathbf{x}}_{\alpha / \beta}$ and $\hat{\mathbf{y}}_{\alpha / \beta}$ respectively lie on the tangent planes at the points $\boldsymbol{\alpha} / \boldsymbol{\beta}$. The choices of orientation of the axes within the tangent planes are somewhat arbitrary; to ease the calculation we shall choose $\hat{\mathbf{x}}_{\alpha}$ to be parallel to $\hat{\mathbf{x}}_{\beta}$ (and thus parallel to the line of intersection of the two tangent planes), which fixes $\hat{\mathbf{y}}_{\alpha}$ and $\hat{\mathbf{y}}_{\beta}$. In this coordinate doublet, we shall label the local coordinates of point $\boldsymbol{\alpha}_{i}$ by $\left(x_{i}, y_{i}\right)$. The three-dimensional position vectors are then determined as follows: (i) If $i$ is even, use the coordinate system at point $\boldsymbol{\alpha}$; otherwise use the coordinate system at point $\boldsymbol{\beta}$. (ii) Then the
three-dimensional position vector in the chosen coordinate system can be expressed as $\boldsymbol{\delta}_{i} \equiv\left(x_{i}, y_{i}, z_{\eta_{i}}\left(x_{i}, y_{i}\right)\right)$, where

$$
\boldsymbol{\eta}_{i} \equiv \begin{cases}\boldsymbol{\alpha}, & \text { for } i \text { even }  \tag{6.13}\\ \boldsymbol{\beta}, & \text { for } i \text { odd }\end{cases}
$$

and $z_{\eta_{i}}(x, y)$ is the local equation of $\partial \mathcal{V}$ near the point $\boldsymbol{\eta}_{i}$. We first expand $\ell_{\alpha_{2 i} \alpha_{2 i+1}}$ to second order:

$$
\begin{aligned}
& \ell_{\alpha_{2 i-1} \alpha_{2 i}} \cong \ell_{\alpha \beta}+\mathbf{n}_{\alpha \beta} \cdot\left(\boldsymbol{\delta}_{2 i-1}-\boldsymbol{\delta}_{2 i}\right)+\frac{\left|\boldsymbol{\delta}_{2 i-1}-\boldsymbol{\delta}_{2 i}\right|^{2}}{2 \ell_{\alpha \beta}}-\frac{\left(\mathbf{n}_{\alpha \beta} \cdot \boldsymbol{\delta}_{2 i-1}-\mathbf{n}_{\alpha \beta} \cdot \boldsymbol{\delta}_{2 i}\right)^{2}}{2 \ell_{\alpha \beta}} \\
& \ell_{\alpha_{2 i} \alpha_{2 i+1}} \cong \ell_{\alpha \beta}+\mathbf{n}_{\alpha \beta} \cdot\left(\boldsymbol{\delta}_{2 i+1}-\boldsymbol{\delta}_{2 i}\right)+\frac{\left|\boldsymbol{\delta}_{2 i+1}-\boldsymbol{\delta}_{2 i}\right|^{2}}{2 \ell_{\alpha \beta}}-\frac{\left(\mathbf{n}_{\alpha \beta} \cdot \boldsymbol{\delta}_{2 i+1}-\mathbf{n}_{\alpha \beta} \cdot \boldsymbol{\delta}_{2 i}\right)^{2}}{2 \ell_{\alpha \beta}}
\end{aligned}
$$

where $\mathbf{n}_{\alpha \beta} \equiv(\boldsymbol{\beta}-\boldsymbol{\alpha}) / \ell_{\alpha \beta}$. We then use this approximation to re-write $S_{\text {And }}$ :

$$
\begin{equation*}
S_{\mathrm{Andreev}} \approx-\frac{1}{\ell_{\alpha \beta}} \sum_{i=1}^{n-1}\left(\boldsymbol{\delta}_{2 i-1} \cdot \boldsymbol{\delta}_{2 i}+\left(\boldsymbol{\delta}_{2 i-1} \cdot \mathbf{n}_{\beta \alpha}\right)\left(\boldsymbol{\delta}_{2 i} \cdot \mathbf{n}_{\alpha \beta}\right)\right) . \tag{6.14}
\end{equation*}
$$

Note that $z_{\eta_{i}}(x, y)$ is a quadratic function of $x$ and $y$ and therefore, for the purposes of expanding $S_{\text {And }}$ to second order in $\left\{x_{i}, y_{i}\right\}$, we can neglect the $\mathbf{n}_{\eta_{i}}$ component of $\boldsymbol{\delta}_{i}$. Thus, the curvature of $\partial \mathcal{V}$ does not feature in $I_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\beta} ; k_{\mathrm{F}}\right)$, which implies that the stability of these trajectories does not depend on the curvature [43]. Then, $S_{\text {And }}$ can be written as follows:

$$
\begin{align*}
S_{\text {Andreev }} & \approx \frac{1}{\ell_{\alpha \beta}} \sum_{i=1}^{n-1}\left(\begin{array}{ll}
x_{2 i-1} & y_{2 i-1}
\end{array}\right) \cdot \mathbf{D} \cdot\binom{x_{2 i}}{y_{2 i}}  \tag{6.15}\\
\mathbf{D} & \equiv\left(\begin{array}{cc}
1-n_{x}^{2} & -n_{x}\left(n_{y} \cos \phi+n_{z} \sin \phi\right) \\
-n_{x} n_{y} & \cos \phi-n_{y}\left(n_{y} \cos \phi+n_{z} \sin \phi\right)
\end{array}\right) \tag{6.16}
\end{align*}
$$

where $n_{x}, n_{y}$ and $n_{z}$ are, respectively, the $x, y$ and $z$ coordinates of $\left.\mathbf{n}_{\alpha \beta}\right)$ in the coordinate system at point $\boldsymbol{\alpha}$, and $\cos \phi$ is the angle between the $y$ axes of the coordinate systems at the points $\boldsymbol{\alpha}$ and $\boldsymbol{\beta}$. Then

$$
\begin{equation*}
I_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\beta} ; k_{\mathrm{F}}\right)=\int \prod_{i=1}^{2 n-2} d x_{i} d y_{i} \exp \left(\left(i k_{\mathrm{F}} / \ell_{\alpha \beta}\right) \mathbf{X}^{\dagger} \cdot \boldsymbol{\mathcal { D }}(n) \cdot \mathbf{X}\right) \tag{6.17}
\end{equation*}
$$

$$
\mathbf{X} \equiv\left(\begin{array}{c}
x_{1}  \tag{6.18}\\
y_{1} \\
\vdots \\
x_{2 n-2} \\
y_{2 n-2}
\end{array}\right), \quad \boldsymbol{D}(n) \equiv\left(\begin{array}{ccccc}
\mathbf{0} & \mathbf{D} & \mathbf{0} & \cdots & \mathbf{0} \\
\mathbf{D}^{\dagger} & \mathbf{0} & -\mathbf{D} & & \\
\mathbf{0} & -\mathbf{D}^{\dagger} & \ddots & & \\
\vdots & & & & -\mathbf{D} \\
\mathbf{0} & & & -\mathbf{D}^{\dagger} & \mathbf{0}
\end{array}\right)_{(4 n-4) \times(4 n-4)}
$$

Thus we arrive at the following expression for $I_{n}$ :

$$
\begin{align*}
I_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\beta} ; k_{\mathrm{F}}\right) & =\left(2 \pi \ell_{\alpha \beta} / k_{\mathrm{F}}\right)^{2 n-2}(\operatorname{det} \boldsymbol{\mathcal { D } ( n )})^{-1 / 2} \exp (i \pi \operatorname{sgn} \mathcal{D} / 4) \\
& =\left(2 \pi \ell_{\alpha \beta} / k_{\mathrm{F}}\right)^{2 n-2}\left(\cos \theta_{\alpha \beta} \cos \theta_{\beta \alpha}\right)^{1-n} \tag{6.19}
\end{align*}
$$

where $\operatorname{sgn} \mathcal{D}$ denotes the signature of the matrix $\mathcal{D}$ (i.e. the number of positive eigenvalues minus the number of negative ones). By inserting this expression into Eq. (6.11) we obtain the contribution to the oscillatory part of the DOS associated with the $2 n$ reflection term:

$$
\begin{equation*}
\rho_{2 n}^{ \pm}(E) \approx \operatorname{Re} \int d \sigma_{\alpha} d \sigma_{\beta}\left(\frac{k_{\mathrm{F}} \cos \theta_{\alpha \beta} \cos \theta_{\beta \alpha}}{2 \ell_{\alpha \beta}}\right) \exp \left(-2 i n \varphi+i n\left(k_{+}-k_{-}\right) \ell_{\alpha \beta}\right) \tag{6.20}
\end{equation*}
$$

By collecting together the contributions from all numbers of reflections ( $n=1,2, \ldots$ ) we arrive at the following formula for the oscillatory part of the DOS in Scheme A:

$$
\begin{equation*}
\delta \rho_{\gamma}(E) \approx \int d \sigma_{\alpha} d \sigma_{\beta} \frac{k_{\mathrm{F}} \cos \theta_{\alpha \beta} \cos \theta_{\beta \alpha}}{4 \pi^{2} \ell_{\alpha \beta}} \operatorname{Re} \frac{\exp \left(-i 2 \varphi+i \frac{E}{k_{\mathrm{F}}} \ell_{\alpha \beta}-\frac{\gamma}{k_{\mathrm{F}}} \ell_{\alpha \beta}\right)}{1-\exp \left(-i 2 \varphi+i \frac{E}{k_{\mathrm{F}}} \ell_{\alpha \beta}-\frac{\gamma}{k_{\mathrm{F}}} \ell_{\alpha \beta}\right)} . \tag{6.21}
\end{equation*}
$$

A similar calculation for the $d$-dimensional case leads to the result

$$
\begin{equation*}
\delta \rho_{\gamma}(E) \approx \int d \sigma_{\alpha} d \sigma_{\beta} \frac{\cos \theta_{\alpha \beta} \cos \theta_{\beta \alpha}}{(2 \pi)^{d}}\left(\frac{k_{\mathrm{F}}}{\ell_{\alpha \beta}}\right)^{d-2} \operatorname{Re} \frac{\exp \left(-i 2 \varphi+i \frac{E}{k_{\mathrm{F}}} \ell_{\alpha \beta}-\frac{\gamma}{k_{\mathrm{F}}} \ell_{\alpha \beta}\right)}{1-\exp \left(-i 2 \varphi+i \frac{E}{k_{\mathrm{F}}} \ell_{\alpha \beta}-\frac{\gamma}{k_{\mathrm{F}}} \ell_{\alpha \beta}\right)} . \tag{6.22}
\end{equation*}
$$

Formula (6.22) for the oscillatory part of the DOS simplifies further in the $\gamma \rightarrow 0$ limit. To see this, consider the following familiar identity:

$$
\begin{align*}
\operatorname{Re} \frac{\mathrm{e}^{-i 2 \varphi+i \frac{E}{k_{\mathrm{F}}} \ell_{\alpha \beta}-\frac{\gamma}{k_{\mathrm{F}}} \ell_{\alpha \beta}}}{1-\mathrm{e}^{-i 2 \varphi+i \frac{E}{k_{\mathrm{F}}} \ell_{\alpha \beta}-\frac{\gamma}{k_{\mathrm{F}}} \ell_{\alpha \beta}}} & =\frac{1}{2} \sum_{n=-\infty}^{\infty} \mathrm{e}^{-i 2 n \varphi+i n \frac{E}{k_{\mathrm{F}}} \ell_{\alpha \beta}}-\frac{1}{2} \\
& =\pi \sum_{m=-\infty}^{\infty} \delta\left(\frac{E}{k_{\mathrm{F}}} \ell_{\alpha \beta}-2 \varphi-2 m \pi\right)-\frac{1}{2} \tag{6.23}
\end{align*}
$$

Applying this identity to Eq. (6.22), we obtain

$$
\begin{array}{r}
\delta \rho_{\gamma}(E) \approx \sum_{m=-\infty}^{\infty} \int d \sigma_{\alpha} d \sigma_{\beta} \frac{\cos \theta_{\alpha \beta} \cos \theta_{\beta \alpha}}{2(2 \pi)^{d-1}}\left(\frac{k_{\mathrm{F}}}{\ell_{\alpha \beta}}\right)^{d-2} \delta\left(\frac{E}{k_{\mathrm{F}}} \ell_{\alpha \beta}-2 \varphi-2 m \pi\right) \\
-\int d \sigma_{\alpha} d \sigma_{\beta} \frac{\cos \theta_{\alpha \beta} \cos \theta_{\beta \alpha}}{2(2 \pi)^{d}}\left(\frac{k_{\mathrm{F}}}{\ell_{\alpha \beta}}\right)^{d-2} \tag{6.24}
\end{array}
$$



Figure 6.3: Jacobian of the transformation from boundary points to scattering parameters

The last term in Eq. (6.24) is a constant term, and it exactly cancels the leadingorder Weyl (i.e. bulk) term. In order to see this, consider the coordinate transformation from $(\boldsymbol{\alpha}, \boldsymbol{\beta})$ to $(\mathbf{b}, \mathbf{n})$, where $\mathbf{n}$ is the direction of the chord and $\mathbf{b}$ is the position-vector specifying the intersection of the chord with the plane perpendicular to $\mathbf{n}$ (i.e. the impact parameter). The transformation of the surface elements are as follows (see Fig. 6.3):

$$
\begin{equation*}
d \sigma_{\alpha} d \sigma_{\beta} \cos \theta_{\alpha \beta} \cos \theta_{\beta \alpha} \ell_{\alpha \beta}^{1-d}=d \mathbf{n} d \mathbf{b} \tag{6.25}
\end{equation*}
$$

By using the (b, n) coordinate system, the last term in Eq. (6.24) can be cast into the following form:

$$
\begin{equation*}
-\frac{1}{2(2 \pi)^{d}} k_{\mathrm{F}}^{d-2} \int d \mathbf{n} d \mathbf{b} \ell_{\alpha \beta}=-\frac{1}{2(2 \pi)^{d}} k_{\mathrm{F}}^{d-2} V \int d \mathbf{n}=-\frac{1}{2(2 \pi)^{d}} k_{\mathrm{F}}^{d-2} V \mathcal{S}^{d-1} . \tag{6.26}
\end{equation*}
$$

A direct comparison with the leading-order Weyl term, given in Eq. (6.2), shows that the term in Eq. (6.26) identically cancels with the leading-order Weyl term. Thus, the DOS in Scheme A can be written as

$$
\begin{equation*}
\rho(E) \approx \sum_{m=-\infty}^{\infty} \frac{k_{\mathrm{F}}^{d-2} \ell_{\alpha \beta}}{2(2 \pi)^{d-1}} \int d \mathbf{n} d \mathbf{b} \delta\left(\frac{E}{k_{\mathrm{F}}} \ell_{\alpha \beta}-2 \varphi-2 \pi m\right), \tag{6.27}
\end{equation*}
$$

which is precisely the result found via the Andreev approximation, as stated in Sec. 3.3.

### 6.2 Scheme B: Mesoscale oscillations beyond the resolution of the Andreev approximation

The main motivation of Scheme B is to capture the mesoscale oscillations in the DOS that are caused by confinement of quasiparticles by the superconducting surround. The reason that Scheme A (and thus the Andreev approximation) is not capable of capturing such oscillations is that in Scheme A the transverse degrees of freedom are not quantized (i.e. quasiparticle motion on the chords is quantal, but there are chords arbitrarily close to one another, indicating that the transverse degrees of freedom are treated classically). On the other hand, in Scheme B we take into account: (i) the imperfectness in retro-reflection (arising from the previously-neglected difference between the wave vectors of incident and reflected electrons and holes); and (ii) the imperfectness in charge-interconversion, the amplitude for which is $\mathcal{O}\left(\Delta / k_{\mathrm{F}}^{2}\right)$. A priori, we know that mesoscale oscillations in the DOS must originate from one or more of the processes not yet taken into account, these being (i) and (ii), above, as well as the subleading quantal corrections (that are ignored in both schemes). In fact, as we shall see, it is the imperfectness in retro-reflection that is primarily responsible for the mesoscale oscillations in the DOS. (Note that the imperfectness in retro-reflection occurs transverse to the incoming direction.) The imperfectness in
charge-interconversion does modify these DOS oscillations. However, by itself (such as in a model with perfect retro-reflection but imperfect charge-interconversion) it is not capable of producing them. In order to clarify this issue, and thus to assess the significance of these processes, we shall define a useful intermediate model: the Perfectly Charge-Interconverting Model (PCIM). This PCIM has the feature of being fully quantum-mechanical; however, in it, any single reflection from the boundary is certain to have the effect of converting electrons to holes (and vice versa). As it contains all quantal effects, the comparison of its predictions with those from the semiclassical approach enable us to assess the importance of quantal effects beyond the semiclassical limit. Moreover, from this comparison it is possible to draw conclusions regarding whether it is imperfectness in retro-reflection that is capable of capturing the mesoscale oscillations in the DOS.

### 6.2.1 Perfectly Charge-Interconverting Model

We are now at a position to define a Perfectly Charge-Interconverting Model (PCIM). We start with the expansion for $\mathbf{G}$ in terms of $\mathbf{G}^{\mathbf{R}}$ in Eq. (5.59). However, we shall replace $\mathbf{G}^{\mathrm{R}}$ by its leading-order form, i.e., $-i \mathrm{e}^{-i \varphi} \boldsymbol{\sigma}_{1} \mathbf{G}^{\mathrm{N}}$. Then the resulting model is defined as an integral equation for $\mathbf{G}$, residing inside the billiard:

$$
\begin{equation*}
\mathbf{G}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\mathbf{G}^{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)-2 i \mathrm{e}^{-i \varphi} \int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial \mathbf{G}^{\mathrm{N}}(\mathbf{x}, \boldsymbol{\alpha}) \boldsymbol{\sigma}_{1} \mathbf{G}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) \tag{6.28}
\end{equation*}
$$

The off-diagonal matrix $\boldsymbol{\sigma}_{1}$ ensures that, upon each reflection from the boundary, electrons are totally converted into holes (and vice versa). Moreover, this model is fully quantum-mechanical, in the sense that it retains wave-propagation effects (as implied by the presence of surface integrals.) Let us now focus on the DOS of the PCIM, treated at the semiclassical level.

Due to the imperfectness in retro-reflection in Scheme B, the corresponding classical dynamics is no longer a priori integrable; on the contrary, it is weakly chaotic
for generic shapes [5]. However, the closed periodic orbits do fall into two quite distinct classes: one consists of multiple tracings of each stationary chord (i.e. chord of stationary length, which we refer to as a $\Sigma$ ); the other of much longer trajectories that "creep" around the billiard boundary (see Fig. 5.3). Correspondingly, the DOS is the sum of: (i) an average term, which depends on the volume of the billiard (i.e. the leading Weyl term); (ii) a finer-resolution term, having a universal lineshape that depends solely on the length and endpoint-curvatures of the $\Sigma s$; and (iii) highest resolution terms, which depend on the classical dynamics of the billiard in question.

### 6.2.2 Stationary chords

Here, we focus on the contribution to the DOS coming from stationary chords. As we shall derive below, it is possible to give a closed-form expression for the contribution of a $\Sigma$ to the DOS for billiards of generic shape. It is, however, necessary to distinguish between isolated $\Sigma$ s and degenerate $\Sigma$ s. First, let us consider an isolated $\Sigma$ in a 2D billiard, with endpoint curvatures $R_{1}$ and $R_{2}$. Then the $\Sigma$ contribution to the DOS from $2 n$ reflections is given by

$$
\begin{equation*}
\rho_{\Sigma, \pm}^{n}=\frac{\ell_{\Sigma}}{2 \pi k_{ \pm}} \operatorname{Re}\left(\frac{k_{+}}{2 \pi \ell_{\Sigma}}\right)^{n / 2}\left(\frac{k_{-}}{2 \pi \ell_{\Sigma}}\right)^{n / 2} \exp \left(i n\left(k_{+}-k_{-}\right) \ell_{\Sigma}-i 2 n \varphi\right) I_{n} \tag{6.29}
\end{equation*}
$$

Here, $I_{n}$ is the Gaussian integral resulting from the expansion of the action to second order:

$$
\begin{equation*}
I_{n} \equiv \int \prod_{i=1}^{2 n} d x_{i} \mathrm{e}^{i \sum_{i=1}^{n}\left(\frac{k_{+}}{\ell_{\Sigma}} x_{2 i-1} x_{2 i}-\frac{k_{-}}{\ell_{\Sigma}} x_{2 i} x_{2 i+1}+\frac{k_{+}-k_{-}}{2}\left(\frac{1}{\left.\left.\ell_{\Sigma}-\frac{1}{R_{1}}\right) x_{2 i-1}^{2}+\left(\frac{1}{\ell_{\Sigma}}-\frac{1}{R_{2}}\right) x_{2 i}^{2}\right)}, ~, ~\right.\right.} \tag{6.30}
\end{equation*}
$$

where $x_{2 n+1} \equiv x_{1}$. In order to evaluate this integral, we define a matrix $M$ such that:

$$
\begin{array}{r}
\sum_{i, j} x_{i} M_{i j} x_{j} \equiv \sum_{i=1}^{n} i\left\{\frac{k_{+}}{\ell_{\Sigma}} x_{2 i-1} x_{2 i}-\frac{k_{-}}{\ell_{\Sigma}} x_{2 i} x_{2 i+1}+\left(\frac{1}{\ell_{\Sigma}}-\frac{1}{R_{2}}\right) x_{2 i}^{2}\right. \\
\left.+\frac{k_{+}-k_{-}}{2}\left(\frac{1}{\ell_{\Sigma}}-\frac{1}{R_{1}}\right) x_{2 i-1}^{2}\right\} \tag{6.31}
\end{array}
$$

(Note that, by definition, an isolated stationary chord is one for which none of the eigenvalues of $M$ is zero; whenever a zero eigenvalue occurs, the stationary chord is
said to be degenerate.) Then, for an isolated $\Sigma, I_{n}$ can be expressed in terms of the determinant and signature of $M$ :

$$
\begin{equation*}
I_{n}=\frac{\pi^{n}}{\sqrt{|\operatorname{det} M|}} \exp \left(i \frac{\pi}{4} \operatorname{sgn} M\right) \tag{6.32}
\end{equation*}
$$

The eigenvalues of $M$ can be obtained from the eigenvalue equation $\sum_{j} M_{i j} x_{j}=\lambda x_{i}$. The symmetry of $M$ under the transformation $x_{i} \rightarrow x_{i+2}$ restricts the form of the eigenvectors to be

$$
x_{2 j}=A_{m} \mathrm{e}^{i 2 m \pi\left(\frac{2 j}{2 n}\right)}, \quad x_{2 j+1}=B_{m} \mathrm{e}^{i 2 m \pi\left(\frac{2 j+1}{2 n}\right)},
$$

where $m=1,2, \cdots, n$. Then the eigenvalue equation reduces to

$$
\left(\begin{array}{cc}
\frac{k_{+}-k_{-}}{2}\left(\frac{1}{\ell_{\Sigma}}-\frac{1}{R_{2}}\right) & \frac{k_{+}}{\ell_{\Sigma}} \mathrm{e}^{-i \pi \frac{m}{n}}-\frac{k_{-}}{\ell_{\Sigma}} \mathrm{e}^{i \pi \frac{m}{n}}  \tag{6.33}\\
-\frac{k_{-}}{\ell_{\Sigma}} \mathrm{e}^{-i \pi \frac{m}{n}}+\frac{k_{+}}{\ell_{\Sigma}} \mathrm{e}^{i \pi \frac{m}{n}} & \frac{k_{+}-k_{-}}{2}\left(\frac{1}{\ell_{\Sigma}}-\frac{1}{R_{1}}\right)
\end{array}\right)\binom{A_{m}}{B_{m}}=\lambda_{m}\binom{A_{m}}{B_{m}}
$$

The determinant of $M$ is readily obtained as

$$
\begin{align*}
\operatorname{det} M & =\prod_{m=1}^{n} \operatorname{det}\left(\begin{array}{cc}
\frac{k_{+}-k_{-}}{2}\left(\frac{1}{\ell_{\Sigma}}-\frac{1}{R_{2}}\right) & \frac{k_{+}}{\ell_{\Sigma}} \mathrm{e}^{-i \pi \frac{m}{n}}-\frac{k_{-}}{\ell_{\Sigma}} \mathrm{e}^{i \pi \frac{m}{n}} \\
-\frac{k_{-}}{\ell_{\Sigma}} \mathrm{e}^{-i \pi \frac{m}{n}}+\frac{k_{+}}{\ell_{\Sigma}} \mathrm{e}^{i \pi \frac{m}{n}} & \frac{k_{+}-k_{-}}{2}\left(\frac{1}{\ell_{\Sigma}}-\frac{1}{R_{1}}\right)
\end{array}\right) \\
& =\prod_{m=1}^{n}\left[\frac{\ell_{\Sigma}^{2}-\ell_{\Sigma} R_{1}-\ell_{\Sigma} R_{2}}{R_{1} R_{2}}\left(\frac{k_{+}-k_{-}}{\ell_{\Sigma}}\right)^{2}-4 \frac{k_{+} k_{-}}{\ell_{\Sigma}^{2}} \sin ^{2}\left(\pi \frac{m}{n}\right)\right] . \tag{6.34}
\end{align*}
$$

For $n \ll \sqrt{k_{+} k_{-}} /\left(k_{+}-k_{-}\right)$, this expression can be simplified to read

$$
\begin{align*}
\operatorname{det} M & \approx\left(\frac{\ell_{\Sigma}^{2}-\ell_{\Sigma} R_{1}-\ell_{\Sigma} R_{2}}{R_{1} R_{2}}\right)\left(\frac{k_{+}-k_{-}}{\ell_{\Sigma}}\right)^{2}\left(\frac{4 k_{+} k_{-}}{\ell_{\Sigma}^{2}}\right)^{n-1}\left(\prod_{m=1}^{n-1} \sin ^{2}\left(\pi \frac{m}{n}\right)\right)^{2} \\
& =(-1)^{n-1}\left(\frac{\ell_{\Sigma}^{2}-\ell_{\Sigma} R_{1}-\ell_{\Sigma} R_{2}}{R_{1} R_{2}}\right)\left(\frac{k_{+}-k_{-}}{\ell_{\Sigma}}\right)^{2}\left(\frac{k_{+} k_{-}}{\ell_{\Sigma}^{2}}\right)^{n-1} n^{2} . \tag{6.35}
\end{align*}
$$

The signature of $M$ is obtained from the eigenvalue problem (6.33) via the investigation of the signs of the eigenvalues $\lambda_{m}^{ \pm}$[or, equivalently, from the trace and the determinant of the matrix in Eq. (6.33)]. If the determinant is negative then $\lambda_{m}^{+}$ and $\lambda_{m}^{-}$are of differing signs and, hence, the associated pair cancel in the signature of $M$. If the determinant is positive then there is a pair of negative or of positive eigenvalues, the sign of which is determined by the trace. These considerations allow
us to express $\operatorname{sgn} M$ as follows:

$$
\begin{gather*}
\operatorname{sgn} M=\sum_{m=1}^{n}\left(1+\operatorname{sgn}\left[\left(\frac{\ell_{\Sigma}^{2}-\ell_{\Sigma} R_{1}-\ell_{\Sigma} R_{2}}{R_{1} R_{2}}\right)\left(\frac{k_{+}-k_{-}}{\ell_{\Sigma}}\right)^{2}-4 \frac{k_{+} k_{-}}{\ell_{\Sigma}^{2}} \sin ^{2}\left(\pi \frac{m}{n}\right)\right]\right) \\
\times \operatorname{sgn}\left(\frac{2}{\ell_{\Sigma}}-\frac{1}{R_{1}}-\frac{1}{R_{2}}\right) . \tag{6.36}
\end{gather*}
$$

Again there is simplification for $n \ll \sqrt{k_{+} k_{-}} /\left(k_{+}-k_{-}\right)$, because in this regime only $m=n$ term contributes to the sum above and, hence, the expression for $\operatorname{sgn} M$ becomes

$$
\begin{equation*}
\operatorname{sgn} M \approx\left(1+\operatorname{sgn}\left(\frac{\ell_{\Sigma}-R_{1}-R_{2}}{R_{1} R_{2}}\right)\right) \operatorname{sgn}\left(\frac{2}{\ell_{\Sigma}}-\frac{1}{R_{1}}-\frac{1}{R_{2}}\right) . \tag{6.37}
\end{equation*}
$$

In particular, when $R_{1}, R_{2}>0$ we have

$$
\begin{equation*}
\operatorname{sgn} M \approx \operatorname{sgn}\left(R_{1}+R_{2}-\ell_{\Sigma}\right)-1 \tag{6.38}
\end{equation*}
$$

Putting all the pieces together, the expression for the contribution to the DOS originating from isolated stationary chords is as follows:

$$
\begin{align*}
& \rho_{\Sigma, \pm}^{n} \approx \frac{\ell_{\Sigma}}{2 \pi k_{ \pm}} \operatorname{Re}\left(\frac{k_{+}}{\ell_{\Sigma}}\right)^{n / 2}\left(\frac{k_{-}}{\ell_{\Sigma}}\right)^{n / 2} \frac{1}{\sqrt{|\operatorname{det} M|}} \mathrm{e}^{i n\left(k_{+}-k_{-}\right) \ell_{\Sigma}-i 2 n \varphi+i \frac{\pi}{4} \operatorname{sig} M}  \tag{6.39}\\
& \approx \operatorname{Re} \frac{1}{n} \frac{1}{k_{ \pm}} \sqrt{\frac{k_{+} k_{-} \ell_{\Sigma} R_{1} R_{2}}{4 \pi^{2}\left(k_{+}-k_{-}\right)^{2}\left|\ell_{\Sigma}-R_{1}-R_{2}\right|}}  \tag{6.40}\\
& \times \exp \left(i n\left(k_{+}-k_{-}\right) \ell_{\Sigma}-i 2 n \varphi+i \frac{\pi}{4}\left(\operatorname{sgn}\left(R_{1}+R_{2}-\ell_{\Sigma}\right)-1\right)\right) \text {, }
\end{align*}
$$

where the second line is valid for $n \ll \sqrt{k_{+} k_{-}} /\left(k_{+}-k_{-}\right)$. We can now sum this expression to all orders in $n$ to obtain the general form for the contribution to the DOS originating from an isolated stationary chord:

$$
\begin{gather*}
\rho_{\Sigma} \approx \operatorname{Re} \sqrt{\frac{\left(k_{+}+k_{-}\right)^{2} \ell_{\Sigma} R_{1} R_{2}}{4 \pi^{2} k_{+} k_{-}\left(k_{+}-k_{-}\right)^{2}\left|\ell_{\Sigma}-R_{1}-R_{2}\right|}} \exp \left(i \frac{\pi}{4}\left(\operatorname{sgn}\left(R_{1}+R_{2}-\ell_{\Sigma}\right)-1\right)\right) \\
\times \ln \left\{1-\exp \left(i n\left(k_{+}-k_{-}\right) \ell_{\Sigma}-i 2 n \varphi\right)\right\} . \tag{6.41}
\end{gather*}
$$

The DOS expression in Eq. (6.41) is valid when its prefactor is free of singularities. Such singularities would indicate that the fluctuation determinant has a vanishing


Figure 6.4: Example of a billiard that is not circular but has degenerate stationary chords. Opposite segments of the boundary are partially coincident with a pair of concentric circles, as shown; when this happens, $\ell_{\Sigma}=R_{1}+R_{2}$.
eigenvalue. This would mean that the stationary chord is no longer isolated; instead there is a direction, corresponding to the vanishing eigenvalue, along which fluctuations do not change the phase.

Let us now focus on these prefactor singularities. The first kind occurs when $k_{+}=k_{-}$or, equivalently, when $E=0$. The reason that this singularity arises is that when $k_{+}=k_{-}$retro-reflection is perfect and, thus, all chords become stationary. As, in Scheme B, we are implicitly assuming that $k_{+}$and $k_{-}$are distinct, this kind of singularity is an artifact of the approximation scheme, and thus is unphysical. The second kind of singularity occurs when $\ell_{\Sigma}=R_{1}+R_{2}$. This is a geometrical singularity, in the sense that whenever the shape of the billiard is such that chords of stationary lengths are not isolated, such a singularity arises. For example, if the billiard is circular, the chord of stationary length is the diameter of the circle, and its length does not change when it is rotated. In this case $R_{1}=R_{2}=\ell_{\Sigma} / 2$. Another example is shown in Fig. 6.4; see the caption for details. Such singularities can be handled in Scheme B as follows: we fix one (or more, if necessary) reflection points, so that the $\Sigma$ is no longer degenerate; we then integrate the remaining surface integrals in the stationary phase approximations; and, after that, we evaluate the remaining surface integral. In this way, we obtain the general form for the stationary-chord
contribution to the DOS:

$$
\begin{equation*}
\rho_{\Sigma}(E) \approx \operatorname{Re} \sum_{\ell_{\Sigma}} Z_{\Sigma} \mathrm{e}^{i \lambda \pi / 4} \operatorname{Li}_{d-1-\frac{w}{2}}\left(1-\mathrm{e}^{i\left(k_{+}-k_{-}\right) \ell_{\Sigma}-2 i \varphi}\right) . \tag{6.42}
\end{equation*}
$$

Here, $\operatorname{Li}_{n}(z) \equiv \sum_{j=1}^{\infty} z^{j} / j^{n}$ is the polylogarithm function, $w$ is the dimensionality of the degeneracy of the $\Sigma$ (e.g. $w=1$ for a circle), $Z_{\Sigma}$ is a slowly varying real function of energy, which determines the size of the DOS oscillations, and $\lambda$ is a measure of the stability of the $\Sigma$ s, which determines whether the "tail" goes to higher or lower energies. For example, for a degenerate $\Sigma$ and $d=2$ we have

$$
\begin{aligned}
Z_{\Sigma} & =V_{\Sigma} \sqrt{\frac{\left(k_{+}+k_{-}\right)^{2} \ell_{\Sigma} R}{8 \pi^{3} k_{+} k_{-}\left(k_{+}-k_{-}\right)\left|\ell_{\Sigma}-R\right|}}, \\
\lambda & =\operatorname{sgn}\left(R-\ell_{\Sigma}\right)
\end{aligned}
$$

where $V_{\Sigma}$ is the volume of the degenerate $\Sigma$ and $R \equiv\left(R_{1}+R_{2}\right) / 2$.

### 6.2.3 Creeping orbits

In this section, we shall obtain the finer oscillatory structure that lies beyond the stationary-chord contribution. The method for obtaining this structure consists of: (i) finding the classical periodic orbits; (ii) evaluating the surface integrals in the stationary phase approximation (i.e. expanding the action to second order around each classical periodic orbit, thus reducing the surface integrals to Gaussian integrals, and then evaluating the resulting Gaussian integrals); and (iii) summing over all periodic-orbit contributions. For the purposes of illustrating this method, we now focus on the case of a circular Andreev billiard and obtain an expression for this finer oscillatory structure in DOS. In addition to its illustrative purposes, having obtained this expression for the finer DOS oscillations for circular Andreev billiard will become useful when we discuss an additional approximation, valid for shapes with slowly varying curvatures, as we shall do later in this thesis.
(i) Classical Dynamics: The classical dynamics of Andreev billiards is defined through the Scheme B reflection rule: $k_{+} \sin \theta_{+}=k_{-} \sin \theta_{-}$. Here, $\theta_{+/-}$is the angle of inci-
dence (or reflection) for the electron/hole. In the case of circular Andreev billiards, the classical periodic orbits can be specified (up to rigid rotations of the full trajectory) by the number of reflections ( $2 n$ ) and by the number of times the orbit "winds" around the billiard $(j)$. The fact that classical periodic orbits of circular Andreev billiards can be specified by $(n, j)$ is a consequence of the integrability of the classical dynamics.


Figure 6.5: A closed periodic orbit with imperfect retro-reflection; the degree of imperfectness is exaggerated for the purpose of illustration.

Our goal is to express dynamical quantities, such as the action $S_{\mathrm{cl}}\left(n, m, k_{\mathrm{F}}^{2}, E, R\right)$, angles of incidence and reflection, and the length of propagation between two successive reflections (which we shall denote by $\ell_{ \pm}$) in terms of the parameters ( $n, m, k_{\mathrm{F}}^{2}, E, R$ ). Here, $R$ is the radius of the circular billiard. First note that, for a given orbit, the angle at which an electron is incident or reflected is the same at all reflections; the same fact holds for holes (see Fig. 6.5). Next, note that each reflection rotates the direction of the momentum by $\delta \theta \equiv \theta_{-}-\theta_{+}$. Thus, in order for the orbit to close after $2 n$ reflections the momentum direction must be rotated by $2 \pi j$. In other words, $\delta \theta=\pi j / n$. However, $\theta_{ \pm}$, must also satisfy the classical reflection rule $k_{+} \sin \theta_{+}=k_{-} \sin \theta_{-}$. By
using these conditions we find (after a little algebra) that for a closed orbit specified by $n$ and $j$ we have

$$
\begin{equation*}
\cos \theta_{ \pm}= \pm \frac{k_{ \pm}-k_{\mp} \cos (\pi j / n)}{\Delta k}, \quad \ell_{ \pm}=2 R \cos \theta_{ \pm} \tag{6.43}
\end{equation*}
$$

where $\Delta k \equiv\left(k_{\mathrm{F}}^{2}-2 \sqrt{k_{\mathrm{F}}^{2}-E^{2}} \cos (\pi j / n)\right)^{1 / 2}$. Observe that, owing to the fact that $\cos \theta_{+}>\cos \theta_{-}>0$, the possible values of $n$ and $j$ are restricted:

$$
\begin{equation*}
\cos \left(\frac{\pi j}{n}\right)>\frac{k_{-}}{k_{+}} \tag{6.44}
\end{equation*}
$$

Next we evaluate the action:

$$
\begin{align*}
S_{\mathrm{cl}}\left(n, m, k_{\mathrm{F}}^{2}, E, R\right) & \equiv n k_{+} \ell_{+}-n k_{-} \ell_{-} \\
& =2 n\left(k_{+} R \cos \theta_{+}-k_{-} R \cos \theta_{-}\right) \\
& =2 n R\left(k_{\mathrm{F}}^{2}-2 \sqrt{\left.{k_{\mathrm{F}}^{2}-E^{2}}^{2} \cos (\pi j / n)\right)^{1 / 2}}\right. \\
& =2 n \Delta k R . \tag{6.45}
\end{align*}
$$

(ii) Evaluating surface integrals: Having obtained the critical points of the surface integrals in Eq. (6.1), we now proceed to expand the action around these points. In order to do this, we expand each $g_{ \pm}(\boldsymbol{\alpha}, \boldsymbol{\beta})$ in Eq. (6.1) around the critical values of its arguments (i.e. $\left.\left(\boldsymbol{\alpha}_{\mathrm{c}}, \boldsymbol{\beta}_{\mathrm{c}}\right)\right)$. This can be accomplished by writing

$$
\begin{equation*}
\boldsymbol{\alpha}=\boldsymbol{\alpha}_{\mathrm{c}}+\boldsymbol{\delta}_{\boldsymbol{\alpha}}, \quad \boldsymbol{\beta}=\boldsymbol{\beta}_{\mathrm{c}}+\boldsymbol{\delta}_{\boldsymbol{\beta}} \tag{6.46}
\end{equation*}
$$

and expressing $\boldsymbol{\delta}_{\boldsymbol{\alpha} / \boldsymbol{\beta}}$ in the coordinate system at $\boldsymbol{\alpha} / \boldsymbol{\beta}$ with coordinate axes specified by the normal (i.e. $\mathbf{n}_{\boldsymbol{\alpha} / \boldsymbol{\beta}}$ ) and tangent (i.e. $\mathbf{t}_{\boldsymbol{\alpha} / \boldsymbol{\beta}}$ ) directions of $\partial \mathcal{V}$ at $\boldsymbol{\alpha} / \boldsymbol{\beta}$. In this coordinate system, the coordinates of the boundary point $\boldsymbol{\delta}_{\boldsymbol{\alpha} / \boldsymbol{\beta}}$ can, up to quadratic order, be parametrized as

$$
\begin{equation*}
\boldsymbol{\delta}_{\boldsymbol{\alpha} / \boldsymbol{\beta}} \approx s_{\boldsymbol{\alpha} / \boldsymbol{\beta}} \mathrm{t}_{\boldsymbol{\alpha} / \boldsymbol{\beta}}+\frac{s_{\boldsymbol{\alpha} / \boldsymbol{\beta}}^{2}}{2 R_{\boldsymbol{\alpha} / \boldsymbol{\beta}}} \mathbf{n}_{\boldsymbol{\alpha} / \boldsymbol{\beta}} \tag{6.47}
\end{equation*}
$$



Figure 6.6: A closed periodic orbit with imperfect retro-reflection; the degree of imperfectness is exaggerated for the purpose of illustration, but less so than in Fig. 6.5.
where $R_{\boldsymbol{\alpha} / \boldsymbol{\beta}}$ are the radii of curvature at $\boldsymbol{\alpha} / \boldsymbol{\beta}$. The surface elements can also be expressed using the parameter $s_{\boldsymbol{\alpha} / \boldsymbol{\beta}}$ :

$$
\begin{equation*}
d \sigma_{\boldsymbol{\alpha} / \boldsymbol{\beta}}=d s_{\boldsymbol{\alpha} / \boldsymbol{\beta}}+\mathcal{O}(1 / R) \tag{6.48}
\end{equation*}
$$

Then $\partial g_{ \pm}$can be approximated as

$$
\begin{equation*}
2 \partial g_{ \pm}\left(s_{\boldsymbol{\alpha}}, s_{\boldsymbol{\beta}}\right) \approx \mp \sqrt{\frac{k_{\mathrm{F}} \cos ^{2} \theta_{\boldsymbol{\beta}}}{2 \pi \ell_{\boldsymbol{\alpha}_{\mathrm{c}} \boldsymbol{\beta}_{\mathrm{c}}}}} \exp \left( \pm i k_{ \pm} \ell_{\boldsymbol{\alpha}_{\mathrm{c}} \boldsymbol{\beta}_{\mathrm{c}}} \mp i \pi / 4 \pm i k_{ \pm} \Phi\left(s_{\boldsymbol{\alpha}}, s_{\boldsymbol{\beta}}\right)\right) \tag{6.49}
\end{equation*}
$$

where the phase function $\Phi$ is given by

$$
\begin{array}{r}
\Phi\left(s_{\boldsymbol{\alpha}}, s_{\boldsymbol{\beta}}\right) \equiv s_{\boldsymbol{\beta}} \sin \theta_{\boldsymbol{\beta}}-s_{\boldsymbol{\alpha}} \sin \theta_{\boldsymbol{\alpha}}-s_{\boldsymbol{\beta}}^{2} \cos \theta_{\boldsymbol{\beta}} / 2 R_{\boldsymbol{\beta}}-s_{\boldsymbol{\alpha}}^{2} \cos \theta_{\boldsymbol{\alpha}} / 2 R_{\boldsymbol{\alpha}} \\
+\left(s_{\boldsymbol{\beta}} \cos \theta_{\boldsymbol{\beta}}+s_{\boldsymbol{\alpha}} \cos \theta_{\boldsymbol{\alpha}}\right)^{2} / 2 \ell_{\boldsymbol{\alpha}_{\mathrm{c}} \boldsymbol{\beta}_{\mathrm{c}}} \tag{6.50}
\end{array}
$$

The terms linear in $s_{\boldsymbol{\alpha} / \boldsymbol{\beta}}$ cancel with the linear terms of the next and previous Green function, due to the stationarity feature of the critical point. For the circle we have $\theta_{\boldsymbol{\alpha}}=\theta_{\boldsymbol{\beta}}=\theta_{ \pm}, R_{\boldsymbol{\alpha}}=R_{\boldsymbol{\beta}}=R$, and $\ell_{\boldsymbol{\alpha}_{\boldsymbol{c}} \boldsymbol{\beta}_{\mathrm{c}}}=2 R \cos \theta_{ \pm}$. This allows us to write

$$
\begin{equation*}
\Phi_{ \pm}\left(s_{\boldsymbol{\alpha}}, s_{\boldsymbol{\beta}}\right)=-\frac{\cos \theta_{ \pm}}{4 R}\left(s_{\boldsymbol{\alpha}}-s_{\boldsymbol{\beta}}\right)^{2} \tag{6.51}
\end{equation*}
$$

and thus $2 \partial g_{ \pm}\left(s_{\boldsymbol{\alpha}}, s_{\boldsymbol{\beta}}\right)=2 \partial g_{ \pm}\left(s_{\boldsymbol{\alpha}}-s_{\boldsymbol{\beta}}\right)$. We are now in a position to evaluate the DOS oscillations due to a periodic orbit with $2 n$ reflections that winds $j$ times around the circular billiard:

$$
\begin{align*}
\rho_{ \pm}^{n, j} & =\frac{\ell_{ \pm}}{2 \pi k_{ \pm}} \operatorname{Re} \exp \left(i S_{\mathrm{cl}}\left(n, j, k_{\mathrm{F}}^{2}, E, R\right)-i 2 n \varphi\right) I_{n}  \tag{6.52}\\
I_{n} & \equiv \int \prod_{i=1}^{2 n} d s_{i} 2 \partial g_{+}\left(s_{1}\right) 2 \partial g_{-}\left(s_{1}-s_{2}\right) \cdots 2 \partial g_{+}\left(s_{2 n-1}\right) \tag{6.53}
\end{align*}
$$

Although it is possible to evaluate this Gaussian integral by the usual formula that relates it to the determinant and signature of the quadratic form in the phase, the (rotationally invariant) form of $\Phi$ in Eq. (6.51) allows the evaluation of the integral above in an easier way, viz., via Fourier decomposition. By defining the Fourier transform

$$
\begin{equation*}
2 \partial g_{ \pm}(p) \equiv \int_{-\infty}^{\infty} d s \mathrm{e}^{i p s} 2 \partial g_{ \pm}(s)=i \exp \left( \pm i p^{2} R / k_{ \pm} \cos \theta_{ \pm}\right) \tag{6.54}
\end{equation*}
$$

we are able to write $I_{n, j}$ as

$$
\begin{align*}
I_{n, j} & =\int_{-\infty}^{\infty} \frac{d p}{2 \pi}\left(-2 \partial g_{+}(p) 2 \partial g_{-}(p)\right)^{n} \\
& =\int_{-\infty}^{\infty} \frac{d p}{2 \pi}(-i)^{n} \exp \left(i n p^{2} R\left(\left(k_{+} \cos \theta_{+}\right)^{-1}-\left(k_{-} \cos \theta_{-}\right)^{-1}\right)\right) \\
& =(-i)^{n} \sqrt{\frac{k_{+} \cos \theta_{+} k_{-} \cos \theta_{-}}{4 \pi n R\left(k_{+} \cos \theta_{+}-k_{-} \cos \theta_{-}\right)}} \exp -i \pi / 4 \tag{6.55}
\end{align*}
$$

Then the contribution to the DOS oscillations originating from a creeping orbit with $2 n$ reflections and winding number $j$ can be written as

$$
\begin{align*}
\rho^{n, j}= & \operatorname{Re}\left(\frac{\cos \theta_{+}}{k_{+}}+\frac{\cos \theta_{-}}{k_{-}}\right) \sqrt{\frac{R^{3} k_{+} \cos \theta_{+} k_{-} \cos \theta_{-}}{\pi n\left(k_{+} \cos \theta_{+}-k_{-} \cos \theta_{-}\right)}} \\
& \times(-i)^{n} \exp \left(i S_{\mathrm{cl}}\left(n, j, k_{\mathrm{F}}^{2}, E, R\right)-2 i n \varphi-i \pi / 4\right) \\
= & \operatorname{Re} \sqrt{\frac{R^{3}\left(k_{+}-k_{-} \cos (\pi j / n)\right)\left(k_{+} \cos (\pi j / n)-k_{-}\right)}{\pi n(\Delta k)^{5} k_{+} k_{-}}} E \cos (\pi j / n) \\
& \times \exp (2 i n \Delta k R-2 i n \varphi-i \pi / 4) \tag{6.56}
\end{align*}
$$

(iii) Summing over all creeping orbits: In order to obtain the creeping-orbit contribution to the DOS oscillations, we must sum the contributions specified by Eq. (6.56) for all $n$ and $j$, obeying the restriction given in Eq. (6.44). Although this summation may appear to be problematic, we remind the reader that $E$ has a nonzero imaginary part $\Gamma$, viz., the smoothing width, which suppresses exponentially (in $n$ ) the contributions coming from high $n$ values. Thus, for a given $\Gamma$ it is possible to perform the sum up to a value of $n$ such that any contribution from higher $n$ would be invisible on the scale set by the truncated sum.


Figure 6.7: Density of states oscillations for a circular Andreev billiard: $k_{\mathrm{F}} R=150$; $\Delta / k_{\mathrm{F}}^{2}=0.08$; smoothing width $\Gamma / k_{\mathrm{F}}^{2}=1.1 \times 10^{-4}$.

In Fig. 6.7 we plot three versions of the oscillatory part of the DOS: the Scheme A result (dashed line), which includes all chords but is dominated by the stationary ones; the Scheme B result (dotted line), which includes creeping orbits and stationary chords; and the exact PCIM result (full line), obtained by the numerical solution of the PCIM. Observe that, although the local average behavior of the exact DOS is essentially that captured by Scheme A, in order to capture the mesoscale oscillations beyond this average behavior one must use Scheme B.

Poisson summation and the semiclassical quantization condition: In the case of a circular Andreev billiard the summation over $n$ and $j$ can be performed approximately to all orders. The procedure for doing this involves using the Poisson summation
formula and evaluating the integrals in the stationary phase approximation. By using this procedure it is possible to obtain the energies at which the DOS has simple poles, viz., it is possible to obtain a semiclassical quantization condition. Our starting point is the expression Eq. (6.56) for $\rho^{n, j}$. In terms of this, the DOS oscillations can be written in the form

$$
\begin{equation*}
\rho \approx \operatorname{Re} \sum_{n, j} \frac{a(\pi j / n)}{\sqrt{n}} \exp (i 2 n \Delta k(\pi j / n) R-2 i n \varphi-i \pi / 4) \tag{6.57}
\end{equation*}
$$

where the amplitude $a$ is defined via

$$
a(x) \equiv \sqrt{\frac{R^{3}\left(k_{+}-k_{-} \cos (x)\right)\left(k_{+} \cos (x)-k_{-}\right)}{\pi \Delta k(x)^{5} k_{+} k_{-}}} E \cos (x)
$$

By using the Poisson summation formula we first cast the expression for the DOS in the form

$$
\rho \approx \operatorname{Re} \sum_{n, m} \int d j \frac{a(\pi j / n)}{\sqrt{n}} \exp (2 i n \Delta k(\pi j / n) R-2 i n \varphi-2 i \pi m j-i \pi / 4) .
$$

Here, $m$ has an interpretation as the angular momentum quantum number. We then evaluate the $j$ integral in the stationary phase approximation. The stationary phase points satisfy

$$
\begin{align*}
\cos \left(\frac{\pi j_{c}}{n}\right) & =\left(\lambda \pm \sqrt{(1-\lambda)^{2}-\frac{E^{2}}{k_{\mathrm{F}}^{2}}}\right)\left(1-\frac{E^{2}}{k_{\mathrm{F}}^{2}}\right)^{-1 / 2} \\
& \equiv f\left(m, R, k_{\mathrm{F}}^{2}, E\right) \tag{6.58}
\end{align*}
$$

where $\lambda \equiv m^{2} /\left(k_{\mathrm{F}} R\right)^{2}$, and only real and positive values of $\cos \left(\pi j_{c} / n\right)$ are allowed. The important point to observe here is that $\cos \left(\pi j_{c} / n\right)$ is independent of $n$, which implies that $\Delta k$ is independent of $n$, too. This allows us to write the expression for $\rho$ as

$$
\begin{align*}
\rho \approx \sum_{m}\left(\pi k_{\mathrm{F}}\right. & R(f(m)-\lambda))^{-1 / 2} a\left(\cos ^{-1} f(m)\right) \\
& \times \operatorname{Re} \sum_{n} \exp \text { in }\left(2 \Delta k(m) R-2 \varphi-2 m \cos ^{-1} f(m)\right) . \tag{6.59}
\end{align*}
$$

We are now in a position to perform the sum over $n$ :

$$
\rho \approx \sum_{m}\left(\pi k_{\mathrm{F}} R(f(m)-\lambda)\right)^{-1 / 2} a\left(\cos ^{-1} f(m)\right) \operatorname{Re} \frac{\mathrm{e}^{i\left(2 \Delta k(m) R-2 \varphi-2 m \cos ^{-1} f(m)\right)}}{1-\mathrm{e}^{i\left(2 \Delta k(m) R-2 \varphi-2 m \cos ^{-1} f(m)\right)}}
$$

Thus, the semiclassical approximation to the eigenenergies for a given angular momentum quantum number $m$ (i.e. the DOS peaks) are given implicitly as the roots of

$$
\begin{equation*}
\exp i\left(2 \Delta k(m) R-2 \varphi-2 m \cos ^{-1} f(m)\right)=1 \tag{6.60}
\end{equation*}
$$

In Table 6.1 we compare the eigenenergies of the PCIM calculated via this semiclassical scheme and exactly. From this Table we see that, as expected, the semiclassical results agree with the exact results upto contributions of relative order $1 / k_{\mathrm{F}} R$.

### 6.2.4 Incorporating ordinary reflection

Thus far in our semiclassical treatment we have ignored all amplitudes involving ordinary reflection. For non-grazing incidence [i.e. for $\theta-(\pi / 2) \sim 1$ ] the amplitude for ordinary reflection is very small (in fact, of order $\Delta / k_{\mathrm{F}}^{2} \cos ^{2} \theta$ ). However, orbits that contribute dominantly to the oscillatory structure of the DOS obey $|\theta-(\pi / 2)| \ll 1$, and therefore ordinary reflection amplitudes are not negligible and must be incorporated. This can be done by returning to Eq. (5.59) and re-evaluating the trace formula using the full expression for $\mathbf{G}^{\mathrm{R}}$ (i.e. not just the leading, off-diagonal, term). Then, as a result of treating both charge-interconverting and charge-preserving reflections, the classical limit changes drastically: the initial conditions of specifying position and momentum (and charge, in the case of Andreev billiards) no longer determines the full orbit; on the contrary, each reflection splits the incoming ray into two rays: one (albeit imperfectly) retro-reflecting; the other, ordinarily reflecting. Thus, the classical dynamics is no longer deterministic (i.e. specifying the position and momentum no longer determines the full orbit) and, instead, the initial conditions specify a superposition of orbits. (A similar situation emerges in the context of Schrödinger billiards

Table 6.1: Comparison of the eigenenergies of the PCIM for a circular Andreev billiard computed semiclassically and computed exactly (numerically) for $k_{\mathrm{F}} R=150$; $\Delta / k_{\mathrm{F}}^{2}=0.08$.

| $m$ | $n$ | $E_{m, n} / k_{\mathrm{F}}^{2}$ | $E_{m, n} / k_{\mathrm{F}}^{2}$ | $\Delta E / k_{\mathrm{F}}^{2}$ |
| :--- | :--- | :--- | :--- | :--- |
| 0 | 4 | 0.066694 | 0.066695 | 0.000001 |
| 1 | 4 | 0.066698 | 0.066696 | 0.000002 |
| 2 | 4 | 0.066699 | 0.066700 | -0.000001 |
| 3 | 4 | 0.066708 | 0.066707 | 0.000001 |
| 4 | 4 | 0.066716 | 0.066716 | 0.000000 |
| 70 | 4 | 0.073978 | 0.073967 | 0.000011 |
| 71 | 4 | 0.074204 | 0.074206 | -0.000002 |
| 72 | 4 | 0.074444 | 0.074450 | -0.000006 |
| 73 | 4 | 0.074711 | 0.074700 | 0.000011 |
| 74 | 4 | 0.074951 | 0.074954 | -0.000003 |
| 110 | 3 | 0.067310 | 0.067285 | 0.000025 |
| 111 | 3 | 0.067842 | 0.067880 | -0.000038 |
| 112 | 3 | 0.068466 | 0.068495 | -0.000029 |
| 113 | 3 | 0.069169 | 0.069132 | 0.000037 |
| 114 | 3 | 0.069840 | 0.069791 | 0.000049 |
| $*$ PCIM |  |  |  |  |
| $\dagger$ Semiclassical |  |  |  |  |

when the billiard has sharp jumps in the single-particle potential [20].) Generically, the number of closed orbits increases exponentially as a function of number of reflections (as opposed to linearly, as is the case when ordinary reflection is neglected), but the amplitudes for these orbits are suppressed exponentially (owing to the fact that the amplitudes for ordinary and Andreev reflections are smaller than unity), allowing the number of reflections to take care of any divergences arising from this ray-splitting feature.

The exponential increase in the number of closed orbits, as a function of the number of reflections, makes evaluation of the periodic-orbit sum difficult. For sufficiently smooth shapes there is a way to circumvent this difficulty, which involves resorting to a different approximation scheme, as we shall shortly show. The main motivation for this approximation scheme is as follows.

1. The closer a primitive creeping orbit is to the boundary (i.e. the closer $|\theta-(\pi / 2)|$ is to zero), the shorter it its total length. As the periodic-orbit contributions are suppressed exponentially with their lengths, the creeping orbits that are closer to the boundary contribute more strongly to the DOS oscillations.
2. The closer a creeping orbit is to the boundary, the bigger the ordinary reflection amplitude $\left(\sim \Delta / k_{\mathrm{F}}^{2} \cos ^{2} \theta\right.$.) Thus, the orbits that involve ordinary reflection contribute more strongly if they are close to the boundary.
3. For the orbits close to the boundary (which, by the previous considerations, dominate), consecutive reflections take place very near to each other, and thus "see" only the local curvature of the boundary.

These considerations motivate us to perform an "adiabatic" approximation to the expansion in Eq. (5.59), in which we assume that the curvature of the boundary varies slowly, relative to the rate at which creeping orbits sample the boundary.

Our starting point is the integral equation (5.40) that generates the MRE:

$$
\mathbf{G}^{\mathrm{ii}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \equiv \mathbf{G}^{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\mathbf{x}, \boldsymbol{\alpha}) \cdot \boldsymbol{\mu}^{\mathrm{ii}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right)
$$

$$
\begin{equation*}
\boldsymbol{\mu}^{\mathrm{ii}} \approx 2 \mathbf{G}^{\mathrm{R}}+2 \partial \mathbf{G}^{\mathrm{R}} \boldsymbol{\mu}^{\mathrm{ii}} \tag{6.61}
\end{equation*}
$$

where $\mathbf{G}^{\mathrm{R}}$ is given in Eq. (5.46). These equations can be cast to the following form:

$$
\begin{equation*}
\mathbf{G}^{\mathrm{ii}} \equiv \mathbf{G}^{\mathrm{N}}+2 \partial \mathbf{G}^{\mathrm{N}}\left(1-2 \partial \mathbf{G}^{\mathrm{R}}\right)^{-1} \mathbf{G}^{\mathrm{R}} \tag{6.62}
\end{equation*}
$$

which shows that the approximate poles of $\mathbf{G}^{\mathrm{ii}}$ are given by the poles of $\mathbf{K} \equiv$ $\left(1-2 \partial \mathbf{G}^{R}\right)^{-1}$. Thus, in order to obtain the energy eigenvalues it is sufficient to consider the following integral equation defined on the surface $\partial \mathcal{V}$ :

$$
\mathbf{K}(\boldsymbol{\alpha}, \boldsymbol{\beta})=\mathbf{I} \delta(\boldsymbol{\alpha}, \boldsymbol{\beta})+2 \int_{\partial \mathcal{V}} d \sigma_{\gamma} \partial \mathbf{G}^{\mathrm{R}}(\boldsymbol{\alpha}, \gamma) \mathbf{K}(\boldsymbol{\gamma}, \boldsymbol{\beta})
$$

This equation will have a regular solution (i.e. one without poles) if and only if none of the eigenvalues of the operator $2 \partial \mathbf{G}^{R}$ is equal to unity; conversely, poles of $\mathbf{K}$ occurs at energies for which at least one of the eigenvalues of $2 \partial \mathbf{G}^{\mathrm{R}}$ is equal to unity. Consequentially, we now focus on the following eigenvalue problem defined on the boundary of the billiard:

$$
\begin{equation*}
2 \int_{\partial \mathcal{V}} d \sigma_{\beta} \partial \mathbf{G}^{\mathrm{R}}(\boldsymbol{\alpha}, \boldsymbol{\beta}) \mathbf{u}(\boldsymbol{\beta})=\lambda \mathbf{u}(\boldsymbol{\alpha}) \tag{6.63}
\end{equation*}
$$

We shall work in the coordinate system in which the boundary is parametrized by its arc length $s$. (Recall that we are considering 2D billiards.) Thus, the equation for the boundary is given by the vector function $\boldsymbol{\alpha}(s)$,

$$
\begin{equation*}
\mathbf{t}(s) \equiv d \boldsymbol{\alpha} / d s \tag{6.64}
\end{equation*}
$$

is the tangent vector,

$$
\begin{equation*}
\mathbf{n}(s) \equiv-R(s) d^{2} \boldsymbol{\alpha} / d s^{2} \tag{6.65}
\end{equation*}
$$

is the unit normal vector, and

$$
\begin{equation*}
R(s) \equiv\left|d^{2} \boldsymbol{\alpha} / d s^{2}\right|^{-1} \tag{6.66}
\end{equation*}
$$

is the curvature at the point $\boldsymbol{\alpha}(s)$. Furthermore, we shall transform to the coordinates

$$
\begin{equation*}
t \equiv s-s^{\prime} \quad \text { and } \quad S \equiv \frac{s+s^{\prime}}{2} \tag{6.67}
\end{equation*}
$$

and define

$$
\begin{equation*}
\overline{\mathbf{G}}^{\mathrm{R}}(t, S) \equiv \mathbf{G}^{\mathrm{R}}\left(s, s^{\prime}\right) \tag{6.68}
\end{equation*}
$$

The virtue of this coordinate system is that for a constant-curvature boundary (viz. a circle) $\mathbf{G}^{\mathrm{R}}\left(s, s^{\prime}\right)=\overline{\mathbf{G}}^{\mathrm{R}}\left(s-s^{\prime}\right)$. Thus, if the curvature is slowly varying then $\overline{\mathbf{G}}^{\mathrm{R}}(t, S)$ is a slowly varying function of $S$. In this coordinate system, then, the eigenvalue equation Eq. (6.63) becomes

$$
\begin{align*}
2 \int_{0}^{L} d t \partial \overline{\mathbf{G}}^{\mathrm{R}}(t, s-t / 2) \mathbf{u}(s-t) & =\lambda \mathbf{u}(s)  \tag{6.69}\\
\mathbf{u}(0) & =\mathbf{u}(L) \tag{6.70}
\end{align*}
$$

where $L$ is the length of the boundary. From periodic-orbit theory we already know that the dominant mesoscale oscillations in the DOS arise from the part of phase space in which the component of the excitation momentum lying tangent to the boundary is $\mathcal{O}\left(k_{\mathrm{F}}\right)$. Thus, in order to capture the dominant mesoscale oscillations, it is sufficient to solve the eigenvalue equation (6.69) for the sector of eigenfunctions varying on the length scale $\mathcal{O}\left(k_{\mathrm{F}}\right)$. In this sector of rapidly-varying eigenfunctions, only the small- $t$ behavior of the kernel $\partial \overline{\mathbf{G}}^{\mathrm{R}}(t, S)$ is relevant. In the following, we shall obtain an expression for $\partial \overline{\mathbf{G}}^{\mathrm{R}}(t, S)$ valid for small $t$.

Our starting point is the general expression for $\mathbf{G}$ in polar coordinates,

$$
\mathbf{G}\left(s, s^{\prime}\right)=\sum_{m} \mathbf{R} \cdot\left(\begin{array}{cc}
\frac{i}{4} J_{m}\left(k_{+} r_{<}\right) H_{m}\left(k_{+} r_{>}\right) & 0  \tag{6.71}\\
0 & -\frac{i}{4} J_{m}\left(k_{-} r_{<}\right) H_{m}\left(k_{-} r_{>}\right)
\end{array}\right) \mathrm{e}^{i m \Theta}
$$

where $\mathbf{R}(m / R(s))$ is defined in Eq. (5.47). For $s$ near $s^{\prime}$ (i.e. for small $t$ ), one can choose the "polar" coordinate system, in which a circle of radius $R\left(\left(s+s^{\prime}\right) / 2\right)$ coincides locally with the surface at point $\left(s+s^{\prime}\right) / 2$. Then the expansion (6.71) can be written as

$$
\overline{\mathbf{G}}(t, S) \approx \sum_{m} \mathbf{R} \cdot\left(\begin{array}{cc}
\frac{i}{4} J_{m}\left(k_{+} R(S)\right) H_{m}\left(k_{+} R(S)\right) & 0 \\
0 & -\frac{i}{4} J_{m}\left(k_{-} R(S)\right) H_{m}\left(k_{-} R(S)\right) \tag{6.72}
\end{array}\right)
$$

The normal derivative $\partial \mathbf{G}$ is given by

$$
\left.\partial \overline{\mathbf{G}} \approx \sum_{m} \mathbf{R} \cdot \frac{i}{4}\left(\begin{array}{cc}
\frac{d}{d R}\left(J_{m}\left(k_{+} R\right) H_{m}\left(k_{+} R\right)\right) & 0 \\
0 & -\frac{d}{d R}\left(J_{m}\left(k_{-} R\right) H_{m}\left(k_{-} R\right)\right) \tag{6.73}
\end{array}\right)\right|_{R=R(S)}
$$

Having obtained the approximate form of $\partial \overline{\mathbf{G}}$, valid for small $t$, we shall now seek the eigenfunctions. We shall assume that the rapidly-varying eigenfunctions have the following form:

$$
\begin{equation*}
\mathbf{u}(s)=\overline{\mathbf{u}}(s) \exp i m s / R(s), \tag{6.74}
\end{equation*}
$$

viz., a slowly-varying envelope and a rapidly varying part. Then the eigenvalue equation becomes

$$
\begin{align*}
\lambda \overline{\mathbf{u}}(s) \mathrm{e}^{i m(s) / R(s)} & =2 \int_{0}^{L} d t \partial \overline{\mathbf{G}}^{\mathrm{R}}(t, s-t / 2) \overline{\mathbf{u}}(s-t) \mathrm{e}^{i m(s-t) / R(s-t)},  \tag{6.75}\\
\lambda \overline{\mathbf{u}}(s) & \approx 2 \int_{0}^{L} d t \partial \overline{\mathbf{G}}^{\mathrm{R}}(t, s) \mathrm{e}^{-i m t / R(s)} \overline{\mathbf{u}}(s),  \tag{6.76}\\
\lambda(m, s) \overline{\mathbf{u}}(s) & =2 \int_{0}^{L} d t \partial \overline{\mathbf{G}}^{\mathrm{R}}(t, s) \mathrm{e}^{-i m t / R(s)} \overline{\mathbf{u}}(s),  \tag{6.77}\\
\lambda(m, s) \overline{\mathbf{u}}(s) & =\sum_{m^{\prime}} 2 \partial \overline{\mathbf{G}}^{\mathrm{R}}\left(m^{\prime}, s\right) \int_{0}^{2 \pi R(s)} d t \mathrm{e}^{-i\left(m-m^{\prime}\right) t / R(s)} \\
\lambda(m, s) \overline{\mathbf{u}}(s) & =\sum_{m^{\prime}} 2 \partial \overline{\mathbf{G}}^{\mathrm{R}}\left(m^{\prime}, s\right) 2 \pi R(s) \delta_{m, m^{\prime}}^{\mathrm{R}}\left(m^{\prime}, s\right) \int_{0}^{2 \pi R(s)-L} d t \mathrm{e}^{-i\left(m-m^{\prime}\right) t / R(s)} \overline{\mathbf{u}}(s),  \tag{6.78}\\
& \approx \sum_{m^{\prime}} 2 \partial \overline{\mathbf{G}}^{\mathrm{R}}\left(m^{\prime}, s\right)\left(2 \pi R(s) \delta_{m, m^{\prime}}-(2 \pi R(s)-L)\right) \overline{\mathbf{u}}(s), \\
& \approx 2 \partial \overline{\mathbf{G}}^{\mathrm{R}}\left(m^{\prime}, s\right) \int_{0}^{\mathrm{R}}(m, s) 2 \pi R(s) \overline{\mathbf{u}}(s)+\mathcal{O}(|d R / d s| L) . \tag{6.79}
\end{align*}
$$

By ignoring the term $\mathcal{O}(|d R / d s| L)$ we have reduced the eigenvalue equation to

$$
\begin{align*}
& \frac{\lambda(m, s)}{i \pi R(s)} \overline{\mathbf{u}}(s)=  \tag{6.82}\\
& \left.\mathbf{R} \cdot\left(\begin{array}{cc}
\frac{d}{d R}\left(J_{m}\left(k_{+} R\right) H_{m}\left(k_{+} R\right)\right) & 0 \\
0 & -\frac{d}{d R}\left(J_{m}\left(k_{-} R\right) H_{m}\left(k_{-} R\right)\right)
\end{array}\right)\right|_{R=R(s)} \overline{\mathbf{u}}(s),
\end{align*}
$$

The important point to note about Eq. (6.82) is that it is simply a $2 \times 2$ matrix eigenvalue equation and, thus, its solution is straightforward. After obtaining the eigenvalues $\lambda(m, s)$, the DOS is given by the following formula:

$$
\begin{equation*}
\rho(E)=\int_{0}^{L} \frac{d s}{2 \pi R(s)} \sum_{m} \delta(1-\lambda(m, s ; E))\left|\frac{\partial \lambda(m, s ; E)}{\partial E}\right| \tag{6.83}
\end{equation*}
$$



Figure 6.8: Density of states oscillations for a circular Andreev Billiard: $k_{\mathrm{F}} R=150$; $\Delta / k_{\mathrm{F}}^{2}=0.08 ;$ smoothing width $\Gamma / k_{\mathrm{F}}^{2}=1.1 \times 10^{-4}$.

In Fig. 6.8 the DOS arising from the solution of the eigenvalue equation Eq. (6.82) is compared to the the one arising from the exact solution of the full BDG eigenproblem for the case of a circular Andreev billiard, which we have computed numerically.

In this section we have shown that mesoscale oscillations in DOS essentially arise from imperfectness in retro-reflection. However, in order to correctly account for these oscillations it is necessary to account for the effects of imperfect charge-interconversion as well. The latter effects can be incorporated via an extension of the trace formula, in which a generic closed periodic orbit has both charge-interconverting and chargepreserving reflections. However, for billiard shapes having slowly-varying curvatures, it is possible to obtain an adiabatic approximation to the DOS that bypasses the periodic-orbit summation (which has a number of terms that increases exponentially with the number of reflections) and, hence, reduces the task to the solution of $2 \times 2$ matrix eigenvalue equation. The strength of this method is that it relates the DOS to
the closed form of $\mathbf{R}(m / R(s))$ and, thus, it is readily extendible to cases in which the N-S boundary is not clean (i.e. reflection amplitudes are modified). As seen above, in order to calculate the DOS one simply needs the reflection amplitudes at the billiard boundary.

## Chapter 7

## Concluding remarks; Perspectives

In the first part of this thesis, we have explored semiclassical approaches to the oscillatory part of the density of states of electron-hole excitations confined to Andreev billiards. We have done this by deriving two semiclassical trace formulas, each corresponding to one of two limiting schemes of the physical parameters specifying the billiard. The first of these trace formulas (viz. the Scheme A trace formula, discussed in Sec. 6.1) is essentially equivalent to the conventional quasiclassical approximation scheme first introduced by Andreev [4]. The physical ingredients of this scheme are perfect charge-interconversion and perfect retro-reflection. It captures the coarsest oscillations in the DOS. The second trace formula (viz. the Scheme B trace formula, discussed in Sec. 6.2) not only captures the coarsest oscillations, but also goes beyond this resolution by capturing mesoscale oscillations. At the semiclassical level, mesoscale oscillations arise from orbits featuring imperfect retro-reflection. Although such oscillations are sensitive to charge-preserving reflection amplitudes (in addition to charge-interconverting reflection amplitudes) from the N-S interface, they are present even if there is no charge-preserving reflection.

The methods developed in the present thesis readily apply to settings such as the superconducting proximity effect or the Josephson effects. Cases in which the phase of the pair-potential is relevant can be addressed by the appropriate modification of
the renormalized Green function for complex $\Delta$. In particular, by using the methods described in Sec. 5.4 , it is possible to show that the leading-order behavior of $\mathbf{G}^{R}$ is modified [cf. Eq. (5.58)], becoming

$$
\mathbf{G}^{\mathrm{R}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) \approx \exp (-i \varphi+i \phi)\left(\begin{array}{cc}
0 & -g_{-}^{\mathrm{N}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right)  \tag{7.1}\\
g_{+}^{\mathrm{N}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) & 0
\end{array}\right)
$$

where the phase $\varphi \equiv \cos ^{-1}(E /|\Delta|)$ and the phase $\phi$ is the phase of the pair-potential. Thus, as one might have anticipated, the phase of the pair-potential simply adds to the phase acquired by the reflection. By using this modified form for $G^{R}$ one can, e.g., account for the zero-energy states observed in $\pi$-junctions (i.e. Josephson junctions through which the pair-potential undergoes a single sign-change).

The Multiple Scattering Expansion that we have developed applies to systems consisting of piecewise homogeneous N or S regions. It is possible to generalize this expansion to handle smoothly-varying $\Delta(\mathbf{x})$. This can be achieved via an "energyslicing" construction (rather than the time-slicing kind familiar from, say, the derivation of the Feynman path integral [50]). In this way, one arrives at a functional version of the Multiple Scattering Expansion. To get a feeling why, let us divide the pair-potential range $\left(0, \Delta_{0}\right)$ [where $\left.\Delta_{0} \equiv \max \Delta(\mathbf{x})\right]$ into a large number $N$ of equallyspaced subintervals. Let us also break x-space into sub-regions, in each of which the pair-potential has values lying in solely one of the $N$ energy subintervals. For $N$ large, $\Delta(\mathbf{x})$ can be taken to be constant in each subregion. Then we can apply the Multiple Scattering Expansion formalism for this intermediate system to obtain its Bogoliubov-de Gennes Green function. Finally, by taking the limit $N \rightarrow \infty$ one can recover the full Bogoliubov-de Gennes Green function associated with $\Delta(\mathbf{x})$. Other spatial inhomogeneities, such as in the single-particle potential, can also be handled in this way. Semiclassical approximations, as discussed here, can be applied to this expansion. For systems that have distinct regions of N and S , but in which these regions are modified from the ideal piecewise-homogeneous state because they feature a
smooth variation of pair-potential, can still be regarded as Andreev billiards. In fact, as the amplitude for charge-preserving reflection is expected to diminish for more slowly varying pair-potentials, we expect the Perfectly Charge-Interconverting Model (PCIM) to be more accurate for such systems. On the other hand, billiards constructed by fabricating a normal region inside a superconductor should feature more charge-preserving reflection, owing to interface effects. Such cases can be modelled by suitably modifying the reflection matrix $\mathbf{R}$ in the definition of the renormalized Green function $\mathbf{G}^{\mathrm{R}}$.

Apart from problems related to superconductivity, the methods presented here are also applicable to more conventional topics in quantum chaos. One such application is to the so-called Ray-splitting billiards. These billiards consist of regions of piecewise homogeneous (single-particle) potential, and the (sharp) boundaries between these regions serve as ray-splitting boundaries. By changing the homogeneous N and S Green functions to the homogeneous Helmholtz Green functions appropriate for a given constant potential, the formulation in the present thesis is readily extended [20]. Another potential application is to multidimensional tunneling, studied in path-integral language in Ref. [44]. The Green-function language adopted in the first part of this thesis is especially well suited to the study of tunneling, owing to the energy (rather than time) dependence of the Green function.

One experimentally relevant application of the work presented so far is to antidot Andreev billiards. Such billiards (in particular their magnetotransport properties) were recently studied experimentally by Eroms et al. [26]. In this realization of Andreev billiards, it is the S region that is embedded in the N region, rather than the converse. (An experimental virtue of this geometry is that it is well suited to the study of the effects of weak magnetic fields, as the magnetic flux need not be quantized.) The methods presented in the present thesis are readily applicable to this antidot geometry. In particular, in App. C we describe the new features that
emerge when the N region is nonconvex, as it is in antidot billiards. Moreover, the incorporation of magnetic fields - at least in the weak field case, so that it is simply excluded from the S region-is handled by modifying the Green functions $\left(\mathbf{G}^{\mathrm{N}}\right.$ and $\mathbf{G}^{\mathrm{S}}$ ) to include the magnetic field. As the electromagnetic vector potential further increases the difference between the action (or accumulated phase) of electrons and holes, it will increase the degree of imperfectness in the retro-reflection.

The present work, and in particular approximation Scheme B, provides insight into the general question of when electron dynamics should be handled separately from the hole dynamics in inhomogeneous superconductors. In doing this, it also provides a semiclassical framework for studying the effects of electron/hole symmetry breaking beyond Andreev approximation. From the point of view of quantum chaos, such electron/hole differences lead to the novel dynamics featuring in the present work. Phenomena associated with this should be accessible via experiment and, indeed, the extremely recent experiments by Eroms et al. [26] are paving the way to a thorough experimental exploration of the novel quantal dynamics of electrons and holes in Andreev billiards.

## Chapter 8

## Intermezzo: d-wave superconductivity

High-temperature superconductivity in the cuprate materials was discovered by Bednorz and Müller in 1986 [73]. After this discovery, several families of cuprate materials have been discovered to be high-temperature superconductors. These cuprate families have rather complicated physical/chemical structures; currently there is no widespread consensus regarding the detailed mechanism causing superconductivity in these materials.

The many common features related to superconductivity shared by these materials do, however, suggest an underlying theoretical picture. In particular, it is agreed that what is responsible for superconductivity is the formation and long-range coherence of Cooper pairs. The degree of this pairing can be characterized by the pair-potential $\Delta(\mathbf{x}, \boldsymbol{\rho})$, where $\mathbf{x}$ is the center of mass coordinate and $\boldsymbol{\rho}$ is the relative coordinate associated with the pair. In conventional superconductors (such as those we have dealt with in the first part of this thesis) the geometrical symmetry of the pairing is such that $\Delta$ has no dependence on $\rho$, i.e. pairing is isotropic and we have what is called s-wave symmetry. In contrast, it is now widely accepted that pairing in cuprates is anisotropic and has predominantly d-wave symmetry [51], i.e. $\Delta$ has the
$\boldsymbol{\rho}$-dependence $\Delta(\mathbf{x}, \boldsymbol{\rho}) \propto \cos 2 \theta_{\rho}$, where $\theta_{\rho}$ is the angular variable of the vector $\boldsymbol{\rho}$.
Perhaps the most important feature of the d-wave symmetry of the pair-potential in cuprate superconductors is that $\Delta$ has differing signs on different parts of the Fermi surface. This has important consequences for the quasiparticle spectrum, especially is cases in which it is possible for quasiparticles to scatter between momenta for which $\Delta$ has distinct signs. These consequences have been explored both theoretically [57-60] and experimentally [61] when the quasiparticle scattering is caused by (flat) boundaries. In the next two chapters we shall exploit the consequences of the sign-changing nature of the pair-potential on the quasiparticle states in cuprates in the presence of a strong, extended (but ultimately localized) variation in the single-particle potential as caused, e.g., by an impurity. As we shall see, our semiclassical approach yields an enlightening picture of the origin and properties of low-energy quasiparticle excitations in disordered d-wave superconductors.

## Chapter 9

## Low-energy quasiparticle states near extended scatterers in d-wave

## superconductors

### 9.1 Introduction

In the present work we shall explore the low-energy quasiparticle states available in d-wave superconductors due to the presence of an extended scatterer such as a boundary or an impurity more than a few Fermi wavelengths across. In the context of boundary scattering, such states represent an important signature of sign-variations of the superconducting order parameter, as they have been shown to originate in the possibility of scattering between momentum orientations that are subject to superconducting pair-potentials of differing sign. The main aims of our work are to extend the idea that sign-variations in the superconducting pair-potential lead to low-energy quasiparticle states to the context of scattering by general single-particle potentials, such as those due to impurities (i.e., beyond scattering by boundaries), and to explore the robustness of this effect.

The theoretical framework that we shall adopt is the semiclassical approach to
the quantum-mechanical problem of scattering from the single-particle potential, via which the eigenvalue problem at hand reduces to a family of effectively onedimensional problems for the particle-hole dynamics in the presence of the superconducting pair-potential. Through this approach, we shall be able to see that the density of low-energy quasiparticle states (DOS) is determined solely by the classical scattering properties of the single-particle potential and, furthermore, that this DOS is insensitive to any suppression of the pair-potential that the impurity might cause. This approach also provides us with a framework for classifying and calculating corrections to the DOS at low energies, such as those due to diffraction during scattering from the single-particle potential itself, or due to any pair-potential modifications beyond mere suppression (such as the induction of any out-of-phase components of the pair-potential).

Along the way, we shall discuss the fact that the emerging one-dimensional eigenproblem is a realization of Witten's supersymmetric quantum-mechanical model [52, $53]$ which, via the Witten index $[52,53]$, provides a natural setting in which to explore zero-energy states $[54,55]$. Through this identification with Witten's model we shall see that the conditions under which zero-energy states exist are indeed those mentioned above, viz., propagation between pair-potentials of differing signs. In addition, we shall examine the role played by the semiclassical approximation to the scattering problem vis-à-vis the existence of zero-energy states, and thus see how it is that going beyond this semiclassical approximation introduces transition amplitudes between classical scattering trajectories, causing the dispersion of the formerly zero-energy states, e.g., into one or more low-energy peaks in the DOS.

We would like to stress at the outset that the issue of the origin of the low-energy states, viz., sign changes in the pair-potential, has already been soundly understood and extensively developed theoretically in several contexts: notable examples include the works of Buchholtz and Zwicknagl [57] on p-wave superconductors near surfaces;
and of Hu [59], Buchholtz et al. [58], and Fogelström et al. [60] on d-wave superconductors near flat surfaces. Low-energy states have also received extensive experimental attention in the context of boundary-scattering in high-temperature superconductors. In particular, measurements of the (macroscopic) tunneling conductance [61] reveal a zero-bias anomaly indicating the existence of low-energy states near boundaries.

Apart from the effects of flat boundaries, theoretical research on low-energy quasiparticle resonances in d-wave materials has mostly been concerned with the effects of point-like impurities (i.e., impurities for which the size of the impurity is not much larger than the Fermi wavelength $\lambda_{F}$ ). Of particular interest has been the effect of the impurity strength on the energies and wave functions of the resonances [62, 63]. More recently, attention has been paid to the effects on these resonances of impurityinduced suppression of the superconducting order parameter [64, 65]. Emerging from this body of work is a picture in which each strong, point-like impurity gives rise to a low-energy resonance. This resonance, which would show up in the tunneling DOS as a pair of peaks symmetrically located around zero energy, transforms (in the particle-hole symmetric case) into a single, marginal, bound state at zero energy in the unitary scattering limit. As the impurity strength is reduced, the energy of this resonance moves towards the gap maximum. Moreover, the quantitative details of the band structure and/or order parameter can play important roles [66]. In particular, in particle-hole asymmetric systems the energies of the resonances no longer tend to zero in the unitary limit.

In contrast, the present work suggests that an extended (rather than point-like) impurity induces a zero-energy peak in the DOS with a weight of order the linear size of the impurity (measured in units of the Fermi wavelength). Moreover, the resulting low-energy DOS is much less sensitive to details such as the precise form of the band structure and any in-phase order parameter variations, i.e., the peak at zero energy is inert. In this respect, extended impurities behave more like flat boundaries than
like point-like impurities.
The theoretical distinctions between point-like and extended impurities raised here have, to some extent, been addressed experimentally via scanning tunneling spectroscopy on $\mathrm{Bi}_{2} \mathrm{Sr}_{2} \mathrm{CaCu}_{2} \mathrm{O}_{8}$ surfaces [67, 68]. Work on native defects [67, 68], which often appear to be essentially point-like in STM imaging, yield weak signatures in the (smeared, local) DOS near each defect. Such signatures can each be interpreted as being induced by a point-like impurity that yields a resonance of unit weight. In contrast, the artificially-induced defects described in Ref. [68], which appear to be more extended in STM imaging, show much stronger signatures in the DOS. This is consistent with the idea that extended impurities produce many states, as the present work indicates.

### 9.2 Bogoliubov-de Gennes eigenproblem

We regard the single-quasiparticle excitations as being described by the Bogoliubovde Gennes (BdG) eigenproblem [4, 69]

$$
\left(\begin{array}{cc}
\hat{h} & \hat{\Delta}  \tag{9.1}\\
\hat{\Delta}^{\dagger} & -\hat{h}
\end{array}\right)\binom{u}{v}=E\binom{u}{v}
$$

where the components $u(\mathbf{x})$ and $v(\mathbf{x})$ of the energy eigenstate respectively give the amplitudes for finding an electron and a hole at the position $\mathbf{x}, E$ is the energy eigenvalue, and $\hat{h}=-\nabla^{2}-k_{\mathrm{F}}^{2}+V(\mathbf{x})$ is the one-particle hamiltonian, in which $k_{\mathrm{F}}^{2}$ is the chemical potential [i.e., $k_{\mathrm{F}}\left(\equiv 2 \pi / \lambda_{\mathrm{F}}\right)$ is the Fermi wave vector] and $V$ is the single-particle potential. We have adopted units in which $\hbar^{2} / 2 m=1$, where $m$ is the (effective) mass of the electrons and holes. The operator $\hat{\Delta}$ (which should ultimately be determined self-consistently) is the pair-potential (integral) operator, whose action on the wave functions is specified by the (nonlocal) kernel $\Delta\left(\mathbf{x}, \mathrm{x}^{\prime}\right)$ via: $[\hat{\Delta} v](\mathbf{x})=\int d \mathbf{x}^{\prime} \Delta\left(\mathbf{x}, \mathbf{x}^{\prime}\right) v\left(\mathbf{x}^{\prime}\right)$. We assume that sufficiently far from the scatterer $\Delta$ returns to the value that characterizes the bulk superconductor (e.g., s-wave, d-
wave, mixed, etc.). As we shall see below, our computation of the low-energy DOS is insensitive to the precise form of any suppression of the superconducting order induced by the single-particle potential, and therefore continues to hold when $\Delta$ is replaced by its self-consistent value. However, as we shall also see below, induced modifications of the superconducting order parameter that go beyond simple suppression in a manner that causes local supercurrents [i.e., via the addition of any intrinsically out-of-phase component to $\Delta]$ spoil this robustness.

### 9.3 Andreev's approximation for a strong singleparticle potential

To analyze the BdG eigenproblem we first apply a semiclassical approximation, which reduces the full problem to a family of first-order differential eigenproblems labeled by the classical trajectories of a particle at the Fermi energy in the presence of the full single-particle potential. This amounts to extending the Andreev approximation to situations in which there is a single-particle potential whose energy scale $V_{0}$ is not negligible compared with the Fermi energy. In technical terms, we are making an asymptotic approximation valid when $k_{\mathrm{F}}^{2} \gg\left(\Delta_{0}, E\right), V_{0} \sim k_{\mathrm{F}}^{2}$, and $V(\mathbf{x})$ is slowly varying relative to $\lambda_{\mathrm{F}}$. To implement this approximation we consider the semiclassical solution of

$$
\begin{equation*}
\left(-\nabla^{2}-k_{\mathrm{F}}^{2}+V(\mathbf{x})\right)\left(\mathcal{M}(\mathbf{x}) \mathrm{e}^{i k_{\mathrm{F}} S(\mathbf{x})}\right)=0 \tag{9.2}
\end{equation*}
$$

i.e., the "large" part of the BdG eigenproblem, where both $\mathcal{M}(\mathbf{x})$ and $S(\mathbf{x})$ are taken to be slowly varying (with respect to $\lambda_{F}$ ) [70]. By retaining the first and second powers in $k_{\mathrm{F}}$ we obtain, from Eq. (9.2), the Hamilton-Jacobi equation $|\nabla S(\mathbf{x})|^{2}=$ $1-k_{\mathrm{F}}^{-2} V(\mathbf{x})$ and the conservation condition $\boldsymbol{\nabla} \cdot\left(\mathcal{M}(\mathbf{x})^{2} \boldsymbol{\nabla} S(\mathbf{x})\right)=0$. We then use the resulting semiclassical solution, which is specified in terms of the incoming momentum orientation $\mathbf{n}$ via the asymptotic behavior $S(\mathbf{x} ; \mathbf{n}) \sim \mathbf{n} \cdot \mathbf{x}[71]$ (for $\mathbf{x}$ far from the
scattering center) and includes all of the fast (i.e., order of $\lambda_{F}$ ) variations of the exact BdG eigenfunctions, to perform a generalized separation of rapidly and slowly varying components by writing $(u(\mathbf{x}), v(\mathbf{x}))$ as $\mathcal{M}(\mathbf{x}) \mathrm{e}^{i k_{\mathrm{F}} S(\mathbf{x} ; \mathbf{n})}(\bar{u}(\mathbf{x}), \bar{v}(\mathbf{x}))$, where $\bar{u}$ and $\bar{v}$ are assumed to be slowly varying relative to $\lambda_{F}$. Then, by inserting this form into Eq. (9.1) we obtain $\left[\hat{h}\left(\mathcal{M} \mathrm{e}^{i k_{\mathrm{F}} S} \bar{u}\right)\right](\mathbf{x}) \sim-2 i k_{\mathrm{F}} \mathcal{M}(\mathbf{x}) \mathrm{e}^{i k_{\mathrm{F}} S(\mathbf{x} ; \mathbf{n})}(\boldsymbol{\nabla} S) \cdot(\boldsymbol{\nabla} \bar{u})$.

We now turn to the "small" part of the BdG eigenproblem, which involves the off-diagonal integral operator $\hat{\Delta}$. It is convenient to transform to relative and center-of-mass coordinates, $\mathbf{r}$ and $\mathbf{R}: \bar{\Delta}(\mathbf{r}, \mathbf{R}) \equiv \Delta\left(\mathbf{x}, \mathbf{x}^{\prime}\right), \mathbf{r} \equiv \mathbf{x}-\mathbf{x}^{\prime}, 2 \mathbf{R} \equiv \mathbf{x}+\mathbf{x}^{\prime}$. Then the action of $\hat{\Delta}$ can be asymptotically approximated (for $k_{\mathrm{F}}^{2} \gg \Delta_{0}$ ) as

$$
\begin{align*}
{\left[\hat{\Delta}\left(\mathcal{M} \mathrm{e}^{i k_{\mathrm{F}} S} \bar{u}\right)\right](\mathbf{x})=} & \int d \mathbf{r} \bar{\Delta}(\mathbf{r}, \mathbf{x}-\mathbf{r} / 2) \bar{u}(\mathbf{x}-\mathbf{r} / 2) \mathcal{M}(\mathbf{x}-\mathbf{r} / 2) \mathrm{e}^{i k_{\mathrm{F}} S(\mathbf{x}-\mathbf{r} / 2 ; \mathbf{n})} \\
\approx & \left(\mathcal{M}(\mathbf{x}) \mathrm{e}^{i k_{\mathrm{F}} S(\mathbf{x} ; \mathbf{n})}\right) \bar{u}(\mathbf{x}) \Delta_{\mathrm{eff}}(\mathbf{x} ; \mathbf{n})  \tag{9.3}\\
\Delta_{\text {eff }}(\mathbf{x} ; \mathbf{n}) \equiv & \int d \mathbf{r} \bar{\Delta}(\mathbf{r}, \mathbf{x}-\mathbf{r} / 2) \mathcal{M}(\mathbf{x})^{-1} \mathcal{M}(\mathbf{x}-\mathbf{r} / 2) \\
& \quad \times \exp \left(i k_{\mathrm{F}} S(\mathbf{x}-\mathbf{r} ; \mathbf{n})-i k_{\mathrm{F}} S(\mathbf{x} ; \mathbf{n})\right) \tag{9.4}
\end{align*}
$$

provided we assume that $(\bar{u}(\mathbf{x}), \bar{v}(\mathbf{x}))$ varies much more slowly than $\lambda_{F}$. Thus the task of solving the full BdG eigenproblem (9.1) is reduced to the task of solving the (classical) Hamilton-Jacobi equation, along with the $(2 \times 2)$ first-order partial differential eigenproblem

$$
\left(\begin{array}{cc}
-2 i k_{\mathrm{F}} \boldsymbol{\nabla} S \cdot \boldsymbol{\nabla} & \Delta_{\mathrm{eff}}(\mathbf{x} ; \mathbf{n})  \tag{9.5}\\
\Delta_{\mathrm{eff}}^{*}(\mathbf{x} ; \mathbf{n}) & 2 i k_{\mathrm{F}} \boldsymbol{\nabla} S \cdot \boldsymbol{\nabla}
\end{array}\right)\binom{\bar{u}}{\bar{v}}=E\binom{\bar{u}}{\bar{v}}
$$

In fact, the eigenproblem is an ordinary rather than partial one. To see this, recall the element of Hamilton-Jacobi theory in which one establishes that the solution $S$ of the Hamilton-Jacobi equation (at least for classically allowed regions) is indeed the action computed along the classical trajectory $\mathbf{x}_{c}(\cdot)$ that solves Newton's equation $k_{\mathrm{F}}^{2} \partial_{s}^{2} \mathbf{x}_{\mathrm{c}}(s)=-\nabla V\left(\mathbf{x}_{\mathrm{c}}\right)$ subject to the condition $\left|\partial_{s} \mathbf{x}_{c}(s)\right| \rightarrow 1$ as $s \rightarrow \pm \infty$ (so that the classical motion is at the Fermi energy). Owing to this connection between $\nabla S$
and $\dot{\mathbf{x}}_{\mathrm{c}}$, Eq. (10.7) can be rewritten as [70]

$$
\hat{H}\binom{\bar{u}}{\bar{v}}=E\binom{\bar{u}}{\bar{v}}, \quad \hat{H} \equiv\left(\begin{array}{cc}
-2 i k_{\mathrm{F}} \partial_{s} & \Delta_{\mathrm{eff}}(s) \\
\Delta_{\mathrm{eff}}^{*}(s) & 2 i k_{\mathrm{F}} \partial_{s}
\end{array}\right),
$$

where $\Delta_{\text {eff }}(s)$ is defined to be $\Delta_{\text {eff }}\left(\mathbf{x}_{\mathrm{c}}(s) ; \mathbf{n}\right)$. This family of first-order ordinary differential eigenproblems is parametrized by $\mathbf{n}$ and the impact parameter $b$, which uniquely specify the classical trajectory $\mathbf{x}_{\mathrm{c}}(\cdot)$ from amongst those having energy $k_{\mathrm{F}}^{2}$.

### 9.4 Zero-energy states

To search for zero-energy states it is useful to reduce the eigenproblem via the following sequence of steps. We apply the unitary transformation (in electron-hole space) $\hat{U} \equiv \frac{1}{\sqrt{2}}\left(\begin{array}{cc}1 & 1 \\ i & -i\end{array}\right)$, under which

$$
\begin{align*}
\hat{H} & \rightarrow \hat{H}^{\prime} \equiv \hat{U}^{\dagger} \hat{H} \hat{U}=\left(\begin{array}{cc}
0 & \hat{A} \\
\hat{A}^{\dagger} & 0
\end{array}\right)  \tag{9.6}\\
\hat{A} & \equiv-2 i k_{\mathrm{F}} \partial_{s}-i \Delta_{\mathrm{eff}}(s), \hat{A}^{\dagger} \equiv-2 i k_{\mathrm{F}} \partial_{s}+i \Delta_{\mathrm{eff}}(s) \tag{9.7}
\end{align*}
$$

We emphasize that it is not possible to arrive at this structure for values of $\Delta_{\text {eff }}$ that are intrinsically complex (i.e., cannot be made real by an elementary gauge transformation), as is the case, e.g., for supercurrent-carrying states. The virtue of the structure of Eqs. (9.6) and (9.7) is that it allows us to recognize that zeroenergy eigenfunctions of $\hat{H}^{\prime}$ have the form $\binom{\varphi_{+}}{0}$ or $\binom{0}{\varphi_{-}}$, where the functions $\varphi_{ \pm}$ obey $\left(2 k_{\mathrm{F}} \partial_{s} \mp \Delta_{\text {eff }}\right) \varphi_{ \pm}=0$, provided they exist (i.e., are normalizable). Owing to their first-order nature, these (zero-energy) eigenproblems may readily be integrated to give $\varphi_{ \pm}(s) \propto \exp \pm\left(2 k_{\mathrm{F}}\right)^{-1} \int^{s} d s^{\prime} \Delta_{\text {eff }}\left(s^{\prime}\right)$. However, the ability to normalize $\varphi_{ \pm}$, and therefore the existence of zero-energy eigenvalues, depends on the form of $\Delta_{\text {eff }}$ via the limiting values $\Delta_{ \pm} \equiv \lim _{s \rightarrow \infty} \Delta_{\text {eff }}( \pm s)$ for a given semiclassical path $\mathbf{x}_{\mathrm{c}}(\cdot)$. Specifically, for semiclassical paths for which $\Delta_{+} \Delta_{-}$is negative, one or other (but not both) of $\varphi_{ \pm}$is normalizable and, therefore, for such paths provide precisely one zero-energy eigenvalue. On the other hand, for semiclassical paths for which $\Delta_{+} \Delta_{-}$
is positive, neither of $\varphi_{ \pm}$is normalizable, and therefore such paths provide no zeroenergy eigenvalues.

This diagnostic for when semiclassical paths lead to zero-energy states allows us to assemble the zero-energy contributions to the DOS. If, for the sake of concreteness, we restrict our attention to two-dimensional systems then our approximation to the low-energy DOS has the form

$$
\begin{equation*}
\rho_{\mathrm{SC}}(E)=\delta(E) \frac{k_{\mathrm{F}}}{2 \pi} \int d \mathbf{n} d b\left(1-\operatorname{sgn} \Delta_{+} \operatorname{sgn} \Delta_{-}\right) . \tag{9.8}
\end{equation*}
$$

This formula should have corrections, which vanish as $E$ tends to zero, coming from the nodes in the gap of the homogeneous d-wave state, as well as suppression of the superconducting state near the impurity.

Let us now highlight some features related to Eq. (9.8): (i) The evaluation of Eq. (9.8) requires only knowledge of the classical scattering trajectories for $V$. (ii) The DOS peak is located at zero energy. Corrections to this result, owing inter alia to particle-hole asymmetry, are of relative order $\max \left(1 / k_{\mathrm{F}} R, \Delta_{0} / k_{\mathrm{F}}^{2}\right)$ (where $R$ is the characteristic linear extent of the impurity potential). These corrections lead to the splitting and dispersion of the zero-energy peak. (iii) Only the asymptotic signs of $\Delta$ at the ends of the classical trajectories feature; the DOS is unchanged by deformations of the pair-potential, provided the asymptotic signs are preserved and no out-of-phase components are induced. (iv) The degeneracy of the zero-energy level, and hence the weight of the zero-bias STM peak, is proportional to the typical impurity scattering cross-section (more precisely, of order $R / \lambda_{\mathrm{F}}$ ), the exact weight depending on the form of $V$. (v) For a sparse collection of (extended) scatterers (i.e., scatterers resolvable at the scale $k_{\mathrm{F}} / \Delta_{0}$ ) their contributions to the DOS are roughly additive; for a dense collection (i.e., for nearly overlapping impurity potentials) the contribution scales with the linear size of the collection.

### 9.5 Connection with Witten's model of supersymmetric quantum mechanics and index theory

Having seen, within the context of an explicit computation, the emergence (or otherwise) or zero-energy states, we now discuss the structure that underlies this issue, namely index theory [56]. The relevant aspect of index theory is Witten's index from Witten's model of supersymmetric quantum mechanics (SUSY QM). The specific connection is as follows: $\hat{H}^{\prime 2}$ (c.f. our 9.6) is Witten's SUSY Hamiltonian; $\Delta_{\text {eff }}$ (our 9.4) is Witten's SUSY potential; $A$ and $A^{\dagger}$ (our 9.7) are proportional to Witten's annihilation and creation operators. Indeed, the analysis leading from Eq. (9.6) to the conditions for the existence of a zero-energy state, mirrors the standard SUSY QM analysis.

In SUSY QM, an important tool is the Witten index, i.e., the number of zeroenergy states of the form $\binom{0}{\varphi_{-}}$minus the number of the form $\binom{\varphi_{+}}{0}$. If the Witten index is nonzero then there certainly are zero-energy states (i.e., SUSY is good; see, e.g., Ref. [53], Sec. 2.1). If the Witten index is zero then there may or may not be zeroenergy states, as contributions may cancel. In the present context, we are not prima facie concerned with the Witten index and its properties, but rather with ascertaining the number of zero-energy states. However, owing to the fact that there is at most one zero-energy state for any semiclassical trajectory (because the normalizability condition cannot be simultaneously satisfied by both $\varphi_{+}$and $\varphi_{-}$) the (modulus of the) Witten index does indeed permit the counting of the zero-energy states.

### 9.6 Discussion and outlook

The condition on the existence of zero-energy states, together with Eq. (9.4), provide us with a way of calculating the DOS at low energies by a simple counting of the
number of classical trajectories that start and end with different signs of the superconducting pair-potential [see Eq. (9.8)]. Thus, the DOS at low energies depends only on the classical scattering properties of the single-particle potential.

As we have stressed earlier, this result is valid in the regime in which the singleparticle potential is both spatially extended and strong and the pair-potential is much smaller than the Fermi energy. Before turning to a discussion (and classification) of the generic corrections to this result for the DOS, which arise upon the relaxation of these conditions, we remark that the foregoing approximation scheme and results also hold for spatially extended single-particle potentials that are weaker than the Fermi energy. Moreover, in the regime $V_{0} \lesssim \Delta_{0}$ our results can be extended to the case of rapidly-varying single-particle potentials (such as are due to point-like impurities). However, as the strength of the single-particle potential is diminished, the classical trajectories tend towards straight lines and, hence, the number of trajectories that "see" different signs of the pair-potential is reduced. This results in a corresponding decrease in the degeneracy of the zero-energy level, in accordance with formula (9.8). Indeed, for $V_{0} \lesssim \Delta_{0}$ the trajectories are essentially straight lines. Thus, there would be no zero-energy states, but additional resonances (due to the impurity) may arise if the pair-potential is suppressed.

By contrast, in the regime $V_{0} \sim k_{\mathrm{F}}^{2}$ but $V(\mathrm{x})$ rapidly varying (e.g., for strong, point-like impurities), the approximation scheme that enabled us to reduce the problem to a family of one-dimensional eigenproblems breaks down, due to the fact that the previously-neglected $\boldsymbol{\nabla} \mathcal{M}$ term becomes comparable to previously-retained $\boldsymbol{\nabla} S$ term. The former term introduces diffraction effects in the (quantum-mechanical) scattering from the single-particle potential, as well as tunneling through the classicallyforbidden region. These effects can be viewed as consequences of nonzero transition amplitudes between states associated with the classical trajectories, and would result in the dispersion of the previously-degenerate zero energy states.

Let us conclude by remarking that the presence of an extended-impurity induced subdominant component to the pair-potential, provided it is in-phase with the dominant component, would not change the picture presented here: specifically, formula (9.8) would continue to hold. On the other hand, if an out-of-phase component is induced (i.e. time-reversal symmetry is locally broken, as would be the case, e.g., if the state were locally to become $\mathrm{d}+$ is), this would cause the zero-energy peak in the DOS to split into two peaks of nonzero width [60,61], symmetrically located about zero energy, their lineshapes depending on the full (rather than solely the asymptotic) details of the pair-potential and computable via perturbation theory. If observed (and provided other sources of splitting such as diffraction effects can be ruled out), such a splitting would provide strong evidence for extended-impurity induced local time-reversal symmetry breaking.

## Chapter 10

## Probing d-wave pairing correlations in the pseudogap regime of the cuprate

 superconductors via low-energy states near impurities
### 10.1 Introduction

Among the challenges presented by the high-temperature superconductors, one of the most persistent concerns the anomalous properties of the normal state of these materials. In particular, the suppression of single-particle spectral weight around the Fermi energy [74-76] for temperatures above the superconducting transition temperature $T_{\mathrm{c}}$ of the underdoped cuprates indicates that the electronic behavior of these materials deviates substantially from that of conventional superconductors. There have been several theoretical scenarios proposed to explain this so-called pseudogap behavior [77]. In this Chapter, we shall be concerned with a particular one of these,
viz., the phase-fluctuation scenario [83-86]. According to this scenario, superconducting correlations in the form of Cooper pairing are presumed to exist (and to be responsible for the loss of single-particle spectral weight) for temperatures $T$ below the pseudogap onset temperature $T^{*}$. However, in the intermediate temperature range (i.e. $T_{\mathrm{c}}<T \lesssim T^{*}$ ) the long-range spatial and temporal coherence in the phase of Cooper-pair wave functions (occurring for $T<T_{\mathrm{c}}$ ) is presumed to present only up to intermediate length scales, having been disrupted on longer length scales by the presence and motion of vortex excitations. In other words, although long-range phase coherence is absent in it, the pseudogap regime is quantitatively distinguished from the conventional non-superconducting state by the presence of substantial, residual, local d-wave pairing correlations.

Several recent experimental investigations support the phase-fluctuation scenario as the origin of pseudogap phenomenology, including Refs. [87, 88]. Further steps towards an understanding the nature of the pseudogap regime would be furnished by experimental probes that are targeted towards the question of the existence of the putative local superconducting correlations [89-92]. Such probes would have the potential to discriminate between scenarios based on pairing correlations and those in which the pseudogap is due to some other mode of electronic ordering [77]. The purpose of the present Chapter is to identify one such probe: scanning tunneling spectroscopy (STS) measurements of the single-particle spectral function near extended impurities in the pseudogap regime [93].

Before explaining the nature of this probe, let us pause to recall one of its essential ingredients. It is has long been recognized [58-60] that the scattering of quasiparticles between states corresponding to differing signs of the d-wave pair-potential leads to the existence of low-energy states [94]. Such scattering, and hence low-energy state-formation, occurs, e.g., at suitably oriented surfaces in the cuprates, leading to the observed zero-bias anomaly in the tunneling conductance [61]. In the setting of
impurity (rather than boundary) scattering in d-wave superconductors, low-energy states, which in this case are localized near the impurities, have been observed in STS experiments [67, 68], and have been discussed theoretically for the case of pointlike [95] and extended [96, 97] impurities. We remind the reader that these states coexist with the continuum of low-energy quasiparticle states associated with the nodes at which the d-wave pair-potential vanishes. Distinguishing between impurity states and nodal states is straightforward because the former give a peaked contribution to the spectral function whereas the contribution from the latter vanishes linearly at zero energy. (A clear example of this is furnished by the data reported in Ref. [68].)

Returning to our main task, viz., probing the pseudogap regime for pairing correlations, we now state the central idea on which the present Chapter is based. Let us suppose that the pseudogap state is indeed distinguished by the presence of local (but not long range) d-wave pairing correlations. The lack of long-range phase coherence in such a state complicates the experimental observation of the more conventional signatures of superconducting correlations (such as the Meissner effect). However, the low-energy quasiparticle states occurring near extended impurities are localized in space, i.e., their existence only relies on the presence of local d-wave pairing correlations. In consequence, these states should be only weakly affected by the destruction of long-range superconducting order that occurs at $T_{\mathrm{c}}$, and hence should "survive" the transition into the pseudogap state. Thus, if STS experiments were to reveal a sharp feature in the single-particle spectral function at low voltage bias, this would be evidence for the presence of local d-wave pairing correlations. And if such experiments were able to characterize the temperature- and doping-dependence of the width of this spectral feature, this would provide a characterization of the nature of these finite-range spatial and temporal correlations. Of course, the interpretation of any experiment conducted at nonzero temperature would have to contend with broadening (arising, e.g., from thermal fluctuations of the sample). In order to minimize the
consequent smearing of the spectral function, which has the potential to obscure the very feature being sought, it would be preferable to examine cuprate systems having low values of $T_{\mathrm{c}}$ (e.g. heavily underdoped systems).

Why are we focusing on the case of (non-magnetic) extended impurities as a probe of the pseudogap regime? The main reason underlying this choice is as follows. In the setting of extended impurities (as well as boundaries) in the superconducting state, it has been shown that the existence of low-energy quasiparticle states is a direct consequence of the d-wave nature (or, more precisely, the sign-changing nature) of the pair-potential [96]. Extending this reasoning to the case of extended impurities in the pseudogap regime, there the existence of low-energy states strongly depends on the presence of superconducting correlations of d-wave type.

We note that Kruis et al. [98] have studied the density of states near a pointlike impurity in the context of a simple phenomenological picture of the pseudogap regime. In this picture, the physics of the pseudogap regime is incorporated through the hypothesis that, in the absence of the impurity, the single-particle density of states vanishes linearly at the Fermi energy. At present, the extent to which an approach based on this picture can yield information about pairing correlations in the pseudogap regime is not clear.

The present Chapter is organized as follows. In Sec. 10.2 we provide a framework for discussing the influence of local d-wave pairing correlations on quasiparticle states near extended impurities [99], focusing on the single-particle spectral function. As we shall see, our expression for this spectral function will take the form of a density of states (determined at a fixed, locally phase-randomized d-wave pair-potential) averaged over the fluctuations of the phase field. In Sec. 10.3 we develop a semiclassical scheme for computing this density of states at fixed pair-potential in which we treat the long-wavelength pair-potential phase variations via perturbation theory. This scheme allows us to focus on the contribution to the density of states at low energies,
which involves states arising from changes in sign (as the momentum is varied) of the local pair-potential. At this point we will have obtained an expression for the spectral function which consists of a sum of terms each associated with one classical scattering trajectory that passes through the vicinity of the impurity potential. In Sec. 10.4.1 we perform the average over the configurations of the fluctuating pair-potential arrived at via locally randomizations of the phase, under the assumption that the distribution of phase-field configurations is Gaussian. We make an approximation to the resulting expression for the spectral function that is valid for the case of phase correlations persisting beyond the Cooper-pair size, finally arriving at an expression for the spectral function near an extended impurity that consists of a Gaussian peak the linewidth of which is proportional to the typical gradient of the phase of the pair-potential. In this section we also compute the linewidth of the spectral function in the pseudogap regime, doing so by assuming that the phase fluctuations are governed by the BKT theory of the two-dimensional $X Y$-model. finally, in Sec. 10.5 we make a numerical estimate of the linewidth by invoking the results of recent high-frequency conductivity data [87], and provide some concluding remarks.

### 10.2 Model of cuprate superconductors with an extended impurity

In the present section we formulate the task of obtaining the single-particle spectral function in settings of systems of fermions interacting via some fermion-fermion coupling and also interacting with an external single-fermion potential (which could, e.g., represent an impurity potential). As our aim is to address the phase-fluctuation picture of the pseudogap regime of the cuprates, we envision following the standard field-theoretic route (see, e.g., Ref. [100]) of exchanging the fermion-fermion coupling for a suitable collective quantum field $\Delta$. Thus we arrive at the following formula for
the one-fermion Green function $\mathcal{G}\left(x, x^{\prime}\right)$ :

$$
\begin{gather*}
\mathcal{G}\left(x, x^{\prime}\right)=\frac{\int D \Delta^{\dagger} D \Delta \mathrm{e}^{-S[\Delta]} G\left(x, x^{\prime} ; \Delta\right)}{\int D \Delta^{\dagger} D \Delta \mathrm{e}^{-S[\Delta]}}  \tag{10.1}\\
\left(\begin{array}{cc}
\partial_{\tau}+\hat{h} & \hat{\Delta} \\
\hat{\Delta}^{\dagger} & \partial_{\tau}-\hat{h}
\end{array}\right) G\left(x, x^{\prime} ; \Delta\right)=-\delta\left(x-x^{\prime}\right) \tag{10.2}
\end{gather*}
$$

i.e., a Bogoliubov-de Gennes Green function for pair-potential $\Delta$, suitably averaged over quantum field $\Delta$, the action for which is $S[\Delta]$. Here, $x \equiv(\mathbf{r}, \tau)$ where $\mathbf{r}$ and $\tau$ are respectively the spatial position (in the two-dimensional $\mathrm{CuO}_{2}$ plane) and the Matsubara time. Furthermore, $\hat{h} \equiv-\nabla^{2}-k_{\mathrm{F}}^{2}+V(\mathbf{r})$, in which $k_{\mathrm{F}}^{2}$ is the chemical potential [i.e. $k_{\mathrm{F}}\left(\equiv 2 \pi / \lambda_{\mathrm{F}}\right)$ is the Fermi wave vector], $V$ is the single-particle potential, and we have adopted units in which $\hbar^{2} / 2 m=1, m$ being the (effective) mass of the electrons and holes. The operator $\hat{\Delta}$ is the pair-potential (integral) operator; how it acts is specified by the nonlocal kernel $\Delta\left(x, x^{\prime}\right)$ via $[\hat{\Delta} v](x)=\int d x^{\prime} \Delta\left(x, x^{\prime}\right) v\left(x^{\prime}\right)$.

Our primary interest is in spectral function $\rho(E)$ in the pseudogap regime as well as the superconducting-to-pseudogap transition regime. This quantity can be obtained in the usual way from $\mathcal{G}\left(x, x^{\prime}\right)$ as follows:

$$
\begin{align*}
& \rho(E) \equiv-\left.\pi^{-1} \operatorname{Im} \operatorname{tr} \int d^{2} r \mathcal{G}\left(\mathbf{r}, \mathbf{r} ; \omega_{n}\right)\right|_{i \omega_{n} \rightarrow E+i \delta}  \tag{10.3}\\
& \mathcal{G}\left(\mathbf{r}, \mathbf{r}^{\prime} ; \omega_{n}\right) \equiv \int_{0}^{\beta} d \tau \mathrm{e}^{i \omega_{n} \tau} \mathcal{G}\left(\mathbf{r}, \mathbf{r}^{\prime} ; \tau, 0\right) \tag{10.4}
\end{align*}
$$

where $\operatorname{tr}$ denotes a trace only in the $2 \times 2$ particle-hole space, $\beta \equiv 1 / T$ (i.e., we have chosen units in which Boltzmann's constant $k_{\mathrm{B}}=1$ ) and the $\omega_{n} \equiv(2 n+1) \pi T$ (with $n$ integral)are fermionic Matsubara frequencies, and $\delta=0+$. We shall assume that the temperature is sufficiently high to validate the neglect of nonzero Matsubarafrequency modes of $\Delta$, which amounts to treating $\Delta$ as a classical (i.e. non-quantal) statistical field. Under this static condition, in which the fermion dynamics takes place in the presence of an unchanging $\Delta$ field, the spectral function $\rho(E)$ may be
expressed in the form

$$
\begin{align*}
\rho(E) & =\langle\rho(E ; \Delta)\rangle  \tag{10.5}\\
\rho(E ; \Delta) & =\sum_{n} \delta\left(E-E_{n}\right), \tag{10.6}
\end{align*}
$$

where $\left\{E_{n}\right\}$ is the collection of energy eigenvalues of the following Bogoliubovde Gennes eigenproblem in the presence of a generic classical configuration of $\Delta$ :

$$
\left(\begin{array}{cc}
\hat{h} & \hat{\Delta}  \tag{10.7}\\
\hat{\Delta}^{\dagger} & -\hat{h}
\end{array}\right)\binom{u}{v}=E\binom{u}{v} .
$$

The notation $\langle\cdots\rangle$ denotes the aforementioned static average over $\Delta$, i.e.,

$$
\begin{equation*}
\langle\mathcal{O}\rangle \equiv \frac{\int D \Delta^{\dagger} D \Delta \mathrm{e}^{-S[\Delta]} \mathcal{O}}{\int D \Delta^{\dagger} D \Delta \mathrm{e}^{-S[\Delta]}} . \tag{10.8}
\end{equation*}
$$

(We do not specifically indicate that this functional average is only over static configurations of $\Delta$.) Thus we have expressed the single-particle spectral function $\rho(E)$ in terms of a suitably averaged density of states $\rho(E ; \Delta)$ for a corresponding Bogoliubovde Gennes eigenproblem.

Before proceeding with the analysis of the eigenproblem given in Eq. (10.7), we address the issue of the form of the pairing fluctuations that contribute dominantly to the average in Eq. (10.5). This amounts to a statement about the physical physical picture of the pseudogap regime that we are concerned with. Now, as discussed shortly before Eq. (10.5), we are considering only static configurations of $\Delta$, and therefore we shall henceforth simplify the notation by writing the pair-potential kernel $\Delta\left(x, x^{\prime}\right)$ as $\Delta\left(x, x^{\prime}\right)=\Delta\left(\mathbf{r}, \mathbf{r}^{\prime}\right)$. In addition, it is convenient to transform $\Delta\left(\mathbf{r}, \mathbf{r}^{\prime}\right)$ to relative and center-of-mass coordinates, $\boldsymbol{\rho}$ and $\mathbf{R}$ :

$$
\begin{equation*}
\bar{\Delta}(\boldsymbol{\rho}, \mathbf{R}) \equiv \Delta\left(\mathbf{r}, \mathbf{r}^{\prime}\right), \quad \boldsymbol{\rho} \equiv \mathbf{r}-\mathbf{r}^{\prime}, \quad 2 \mathbf{R} \equiv \mathbf{r}+\mathbf{r}^{\prime} \tag{10.9}
\end{equation*}
$$

It is then convenient to introduce the Fourier transform of $\bar{\Delta}$ with respect to the relative coordinate $\rho$, viz., $\bar{\Delta}(\mathbf{k}, \mathbf{R})$ :

$$
\begin{equation*}
\bar{\Delta}(\mathbf{k}, \mathbf{R}) \equiv \int d^{2} \rho \mathrm{e}^{-i \mathbf{k} \cdot \boldsymbol{\rho}} \bar{\Delta}(\boldsymbol{\rho}, \mathbf{R}) \tag{10.10}
\end{equation*}
$$

thus obtaining a pair-potential at center-of-mass position $\mathbf{R}$ and relative momentum k.

As mentioned in Sec. 10.1, the scenario for the pseudogap regime with which we are concerned is based on the dominance of configurations of $\Delta(\mathbf{k}, \mathbf{R})$ of the form

$$
\begin{equation*}
\bar{\Delta}(\mathbf{k}, \mathbf{R})=\bar{\Delta}_{0}(\mathbf{k}, \mathbf{R}) \mathrm{e}^{i \vartheta(\mathbf{R})} . \tag{10.11}
\end{equation*}
$$

Here, the non-fluctuating factor $\bar{\Delta}_{0}(\mathbf{k}, \mathbf{R})$ is taken to have d-wave form (and can therefore be taken to be real) [101]. The fluctuating factor $\exp i \vartheta(\mathbf{R})$ varies slowly with $\mathbf{R}$. By assuming this form for $\bar{\Delta}(\mathbf{k}, \mathbf{R})$ we are adopting the physical picture of the state as being one in which there is no long-range superconducting order, but there are local d-wave superconducting correlations, embodied in $\bar{\Delta}_{0}(\mathbf{k}, \mathbf{R})$ [102].

In the following section, we obtain $\rho(E ; \Delta)$ by making use of an elaboration of Andreev's semiclassical approach [103] to the Bogoliubov-de Gennes eigenproblem. This elaboration is appropriate for the setting at hand, viz., one in which there is a strong single-particle potential. This scheme was used in Ref. [96] in order to address the low-energy density of states near an extended impurity in a d-wave superconducting state having a negligibly fluctuating, well-formed condensate. In contrast, our focus here is on situations in which the fluctuations in the amplitude of $\Delta$ are small, but the phase of $\Delta$ is strongly fluctuating. Thus, although there is local pairing the system does does not exhibit long-range order.

### 10.3 Semiclassical approach to the Bogoliubov-de Gennes eigenproblem

In Sec. 10.2 we showed how the single-particle spectral function $\rho(E)$ can be expressed as a density of states $\rho(E ; \Delta)$ for the Bogoliubov-de Gennes eigenproblem at arbitrary pair-potential $\Delta$, averaged over $\Delta$ with a suitable weight, provided $\Delta$ can be treated
in the static approximation. To make progress with this Bogoliubov-de Gennes eigenproblem we invoke, in the present section, a semiclassical approximation under which Eq. (10.7) reduces to a family of one-dimensional eigenproblems labeled by the classical scattering trajectories of a particle in the presence of the single-particle potential $V(\mathbf{r})$. This approximation scheme, which was developed in Ref. [96], is valid provided that $k_{\mathrm{F}}^{2} \gg\left\{\Delta_{0}, E\right\}$ for physically relevant configurations of $\Delta$ (where $\Delta_{0}$ is the magnitude of $\Delta$ ), and provided that $V(\mathbf{r})$ and relevant configurations of $\Delta$ vary slowly on the scale of the Fermi wavelength $\lambda_{\mathrm{F}}$.


Figure 10.1: Schematic illustration of a classical trajectory for a quasiparticle scattering from an extended impurity. Also shown is a depiction of the background d-wave pair-potential; dots indicate incoming and outgoing quasiparticle momenta.

We now turn to the family of one-dimensional eigenproblems arising in our semiclassical scheme. Following Ref. [96], it is straightforward to determine that the emerging trajectory-dependent eigenproblem has the form

$$
\begin{align*}
& \hat{H}\binom{\bar{u}}{\bar{v}}=E\binom{\bar{u}}{\bar{v}},  \tag{10.12}\\
& \hat{H} \equiv\left(\begin{array}{cc}
-2 i k_{\mathrm{F}} \partial_{s} & \Delta_{0}(s) \exp i \vartheta(s) \\
\Delta_{0}(s) \exp -i \vartheta(s) & 2 i k_{\mathrm{F}} \partial_{s}
\end{array}\right),  \tag{10.13}\\
& \Delta_{0}(s) \simeq \bar{\Delta}_{0}\left(k_{\mathrm{F}} \partial_{s} \mathbf{x}_{\mathrm{c}}(s), \mathbf{x}_{\mathrm{c}}(s)\right) . \tag{10.14}
\end{align*}
$$

Here, $\vartheta(s) \simeq \vartheta\left(\mathbf{x}_{\mathrm{c}}(s)\right)$, and the parameter $s$ measures the position along a particular classical trajectory $\mathbf{x}_{\mathrm{c}}(s)$, the latter obeying the Newton equation

$$
\begin{equation*}
k_{\mathrm{F}}^{2} \partial_{s}^{2} \mathbf{x}_{\mathrm{c}}(s)=-\boldsymbol{\nabla} V\left(\mathbf{x}_{\mathrm{c}}(s)\right) \tag{10.15}
\end{equation*}
$$

Each such classical trajectory (and associated eigenproblem) is labeled by an impact parameter $b$ and an incoming momentum direction $\mathbf{n}$.

Our next task is to obtain the low-energy eigenvalues associated with Eq. (10.12) for the case of a d-wave pair-potential subject to generic spatial phase variations, i.e., Eq. 10.11. To do this, we perform a local unitary transformation of $\hat{H}$, i.e., $\hat{H} \rightarrow U \hat{H} U^{\dagger}$, where

$$
U(s) \equiv\left(\begin{array}{cc}
\mathrm{e}^{-i \vartheta(s) / 2} & 0  \tag{10.16}\\
0 & \mathrm{e}^{i \vartheta(s) / 2}
\end{array}\right)
$$

and thus our eigenproblem acquires the form

$$
\begin{align*}
& \left(\hat{H}_{0}+\hat{H}_{1}\right)\binom{\bar{u}}{\bar{v}}=E\binom{\bar{u}}{\bar{v}},  \tag{10.17}\\
& \hat{H}_{0}(s) \equiv\left(\begin{array}{cc}
-2 i k_{\mathrm{F}} \partial_{s} & \Delta_{0}(s) \\
\Delta_{0}(s) & 2 i k_{\mathrm{F}} \partial_{s}
\end{array}\right)  \tag{10.18}\\
& \hat{H}_{1}(s) \equiv\left(\begin{array}{cc}
k_{\mathrm{F}} \partial_{s} \vartheta & 0 \\
0 & k_{\mathrm{F}} \partial_{s} \vartheta
\end{array}\right) \tag{10.19}
\end{align*}
$$

The Hamiltonian for this eigenproblem now consists of a term arising from the underlying d-wave pair-potential (i.e. $\hat{H}_{0}$ ) as well as a term that contains all the phasevariation information (i.e. $\hat{H}_{1}$ ). Our strategy is to treat $\hat{H}_{1}$ within perturbation theory, the starting point for which is the identification of the eigenstates of $\hat{H}_{0}$. As our purpose is to address low-energy states, it is adequate for us to focus on the zeroenergy eigenstate of $\hat{H}_{0}$, if any there be. As we shall discuss below, such states are guaranteed to arise for certain classical trajectories $\mathbf{x}_{\mathrm{c}}(s)$ [52, 96]. The condition for the existence of such a state is determined by the asymptotic properties of $\Delta_{0}(s)$ : If $\Delta_{0}(s)$ changes sign an odd number of times along an entire trajectory then the Hamiltonian $\hat{H}_{0}$ associated with that trajectory has precisely one zero-energy eigenstate.

Said more formally, if $\Delta_{ \pm} \equiv \lim _{s \rightarrow \pm \infty} \Delta_{0}(s)$ then the condition for the existence of the zero-energy state is that $\Delta_{+} \Delta_{-}<0$. The explicit form of this eigenstate is

$$
\begin{align*}
\psi_{ \pm}(s) & =\frac{1}{\sqrt{2}}\binom{\varphi_{ \pm}(s)}{ \pm i \varphi_{ \pm}(s)}  \tag{10.20}\\
\varphi_{ \pm}(s) & \equiv \alpha_{ \pm} \exp \pm\left(2 k_{\mathrm{F}}\right)^{-1} \int^{s} d s^{\prime} \Delta_{0}\left(s^{\prime}\right) \tag{10.21}
\end{align*}
$$

where $\psi_{+}$corresponds to the case $\Delta_{+}<0$ (so that $\Delta_{-}>0$ ) and $\psi_{-}$corresponds to the case $\Delta_{+}>0$ (so that $\Delta_{-}<0$ ). The pre-factors $\alpha_{ \pm}$are normalization factors, chosen so that $\int_{-\infty}^{\infty} \varphi_{ \pm}(s)^{2} d s=1$.

Let us emphasize some of the important features of these zero-energy states that hold provided the scatterer creates only trajectories with at most a single sign change in the pair-potential [104]. First, the presence or absence of these states depends only on the properties of the d-wave pair-potential far from any impurity, and is insensitive to any amplitude variations of the d-wave pair-potential that might occur in the vicinity of this impurity. Second, the wavefunctions $\psi_{ \pm}(s)$ exhibit exponential decay away from the impurity with a decay constant of order $k_{\mathrm{F}} /\left|\Delta_{ \pm}\right|$, which is proportional to the BCS correlation length. As this length-scale is known to be short in the cuprates, these states are indeed an extremely local probe of local superconducting correlations.

In the absence of phase fluctuations (i.e. in the pure d-wave superconductor), these zero-energy eigenstates lead to a sharp peak in the density of states at lowenergies [96]. To access the impact of order-parameter phase variations on this peak, we include the effect of $\hat{H}_{1}$ on the zero-energy state $\psi_{ \pm}^{\dagger}(s)$ within quantum-mechanical perturbation theory. Thus, the shift in energy of the previously zero-energy eigenstate associated with a particular trajectory $\mathbf{x}_{\mathrm{c}}(s)$ (parametrized by $\mathbf{n}$ and $b$ ) is given by

$$
\begin{align*}
\epsilon(\mathbf{n}, b) & \equiv \int d s \psi_{ \pm}^{\dagger}(s) \hat{H}_{1}(s) \psi_{ \pm}(s)  \tag{10.22}\\
& =k_{\mathrm{F}} \int d s \partial_{s} \vartheta(s) \varphi_{ \pm}(s)^{2} \tag{10.23}
\end{align*}
$$

where we have used Eq. (10.20) to get from Eq. (10.22) to Eq. (10.23). The spectral function at fixed pair-potential $\rho(E, \Delta)$ consists of contributions from all trajectories $\mathbf{x}_{\mathrm{c}}(s)$. As we are restricting our attention to low energies, it is sufficient to consider only trajectories for which $\Delta_{+} \Delta_{-}<0$ (i.e., those for which $\hat{H}_{0}$ has a zero-energy state). The contribution to the low-energy spectral function at constant pair-potential due to the perturbed zero-energy states is given by

$$
\begin{align*}
\rho(E, \Delta) \simeq k_{\mathrm{F}} \int & \frac{d \mathbf{n}}{2 \pi} \int d b \delta(E-\epsilon(\mathbf{n}, b)) \\
& \times\left(1-\operatorname{sgn} \Delta_{+} \operatorname{sgn} \Delta_{-}\right), \tag{10.24}
\end{align*}
$$

where the factor $\left(1-\operatorname{sgn} \Delta_{+} \operatorname{sgn} \Delta_{-}\right)$ensures that only trajectories that satisfy the condition $\Delta_{+} \Delta_{-}<0$ (i.e. exhibit an asymptotic sign change in $\left.\Delta_{0}(s)\right)$ contribute.

It is important to emphasize that Eq. (10.24) only includes the contribution from the perturbed zero-energy states, and that there will be additional, smaller contributions to the density of states at low energies. Sources of these include remnants of the near-nodal quasiparticles states that exist in pure, d-wave superconductors at arbitrarily low energies.

### 10.4 Averaging over phase fluctuations

In Sec. 10.3 we obtained an expression for the Bogoliubov-de Gennes density of states $\rho(E, \Delta)$ in the presence of an extended impurity for the case of a pair-potential with local d-wave character and specific realization of the spatially varying phase $\vartheta(\mathbf{r})$; see Eq. (10.24). In the present section, we calculate the spectral function $\rho(E)$ in the pseudogap regime by averaging $\rho(E, \Delta)$ over suitable configurations of $\vartheta(\mathbf{r})$, i.e., by inserting Eq. (10.24) into Eq. (10.5). We remind the reader that our scheme for computing $\rho(E)$ applies to settings in which the temperature is high enough that we may treat the phase fluctuations classically, as discussed after Eq. (10.4).

### 10.4.1 Gaussian model for phase fluctuations

In order to evaluate the average over configurations of the pair-potential in Eq. (10.5) explicitly, we need a model for the weight of the various configurations. For the sake of simplicity, we choose the phase-field configurations to have a (zero mean) Gaussian distribution characterized by the correlator $\left\langle\nabla_{\alpha} \vartheta(\mathbf{r}) \nabla_{\beta} \vartheta\left(\mathbf{r}^{\prime}\right)\right\rangle$. In terms of this correlator, it is straightforward to show that the spectral function $\rho(E)$ is a superposition of Gaussian distributions, one associated with each classical trajectory on which $\Delta_{0}$ changes sign:

$$
\begin{align*}
\rho(E) \simeq & k_{\mathrm{F}} \int \frac{d \mathbf{n}}{2 \pi} \int d b \frac{\exp \left(-E^{2} / 2\left\langle\epsilon(\mathbf{n}, b)^{2}\right\rangle\right)}{\sqrt{2 \pi\left\langle\epsilon^{2}\right\rangle}}  \tag{10.25}\\
& \times\left(1-\operatorname{sgn} \Delta_{+} \operatorname{sgn} \Delta_{-}\right) \\
\left\langle\epsilon(\mathbf{n}, b)^{2}\right\rangle \equiv & k_{\mathrm{F}}^{2} \int d s \int d s^{\prime} \varphi_{ \pm}^{2}(s) \varphi_{ \pm}^{2}\left(s^{\prime}\right)\left\langle\dot{\vartheta}(s) \dot{\vartheta}\left(s^{\prime}\right)\right\rangle . \tag{10.26}
\end{align*}
$$

Here, $\dot{\vartheta}(s) \equiv \partial_{s} \vartheta(s)$, and the width $\left\langle\epsilon(\mathbf{n}, b)^{2}\right\rangle$ of each Gaussian contribution is sensitive to the degree to which local superconducting correlations have been disrupted, as can be seen by the presence of the phase-phase correlator $\left\langle\dot{\vartheta}(s) \dot{\vartheta}\left(s^{\prime}\right)\right\rangle$ in Eq. (10.26). Our next task, then, is to evaluate the integrals over the parameters $s$ and $s^{\prime}$ in Eq. (10.26) in order to determine the width $\left\langle\epsilon(\mathbf{n}, b)^{2}\right\rangle$ associated with each trajectory $(\mathbf{n}, b)$. In the next section, we carry out this evaluation within an approximation that is valid for the case of long-wavelength phase fluctuations.

### 10.4.2 Approximate evaluation of trajectory integrals

In Sec. 10.4.1, we obtained the expression (10.25) for the spectral function $\rho(E)$ near an extended impurity in the pseudogap regime. In the present section, we make an approximation to our expression for $\rho(E)$ that makes use of the local nature of the low-energy states, as well as the long-wavelength nature of the pair-potential phase variations (appropriate for $T \gtrsim T_{\mathrm{c}}$ ). By local we mean that the wave functions $\varphi_{ \pm}(s)$ exhibit exponential decay over a length-scale $\xi \sim k_{\mathrm{F}} /\left|\Delta_{ \pm}\right|$(where $\Delta_{0}$ is the
bulk d-wave pair-potential), i.e., the BCS correlation length; this can be seen by examining Eq. (10.21); see [105]. In the cuprate superconductors, this length-scale is expected to be much shorter than the length-scale $\xi_{\vartheta}$ for typical pair-potential phase variations (i.e. the inter-vortex spacing). Thus one has a separation of lengthscales: $\xi_{\vartheta}>\xi$. (Such a separation is a natural ingredient of the phase-fluctuation picture of the pseudogap regime because for inter-vortex spacings on the order of $\xi$ the conventional meaning of local Cooper pairs to break down.)

To make the approximation to our expression for $\rho(E)$, consider the integrations over the trajectory parameters $s$ and $s^{\prime}$ in Eq. (10.26). Now, the correlation function $\left\langle\dot{\vartheta}(s) \dot{\vartheta}\left(s^{\prime}\right)\right\rangle$ varies appreciably only over length-scales on the order of $\xi_{\vartheta}$ or longer, whereas the wavefunctions $\varphi_{ \pm}$decay exponentially, as mentioned, on the length-scale $\xi$. Hence, one can make an asymptotic approximation to the $s$ and $s^{\prime}$ integrations which amounts to pulling the correlator out of the integrals. Thus, owing to the normalization of $\varphi_{ \pm}$, we have

$$
\begin{equation*}
\left\langle\epsilon^{2}\right\rangle \simeq k_{\mathrm{F}}^{2}\langle\dot{\vartheta}(0) \dot{\vartheta}(0)\rangle, \tag{10.27}
\end{equation*}
$$

independent of $\mathbf{n}$ and $b$.
Next, we turn to the interpretation of the quantity $\langle\dot{\vartheta}(0) \dot{\vartheta}(0)\rangle$. The derivative of the phase along a particular trajectory is given by

$$
\begin{equation*}
\dot{\vartheta}(s)=\left.\dot{\mathbf{x}}_{\mathrm{c}}(s) \cdot \boldsymbol{\nabla} \vartheta(\mathbf{x})\right|_{\mathbf{x}=\mathbf{x}_{\mathrm{c}}(s)}, \tag{10.28}
\end{equation*}
$$

By inserting Eq. (10.28) into Eq. (10.27), we see that the correlator of interest is $\left\langle\nabla_{\alpha} \vartheta(\mathbf{r}) \nabla_{\beta} \vartheta(\mathbf{r})\right\rangle$ which, by spatial isotropy, has the form $\left.\left.\delta_{\alpha \beta}\langle | \nabla \vartheta(\mathbf{r})\right|^{2}\right\rangle / 2$. According to the phase-fluctuation scenario, vortex excitations are the dominant mechanism for generating phase gradients .

How does this information about the phase correlations translate into information about the spectral function? By noting that trajectories $\mathbf{x}_{\mathrm{c}}(s)$ involving sign-changes
in the d-wave pair-potential obey $\left|\dot{\mathbf{x}}_{\mathrm{c}}\right| \sim 1$ and using the aforementioned phasegradient correlator, we arrive at the width

$$
\begin{equation*}
\left.\left.\left\langle\epsilon^{2}\right\rangle \sim k_{\mathrm{F}}^{2}\langle | \boldsymbol{\nabla} \vartheta(\mathbf{r})\right|^{2}\right\rangle / 2 \tag{10.29}
\end{equation*}
$$

and, hence, our final expression for the spectral function $\rho(E)$ :

$$
\begin{align*}
\rho(E) \simeq & \frac{\exp \left(-E^{2} / 2\left\langle\epsilon^{2}\right\rangle\right)}{\sqrt{2 \pi\left\langle\epsilon^{2}\right\rangle}}  \tag{10.30}\\
& \quad \times k_{\mathrm{F}} \int \frac{d \mathbf{n}}{2 \pi} \int d b\left(1-\operatorname{sgn} \Delta_{+} \operatorname{sgn} \Delta_{-}\right)
\end{align*}
$$

The first factor is a Gaussian in the energy $E$ with a linewidth $\left\langle\epsilon^{2}\right\rangle^{1 / 2}$ associated with the r.m.s. fluctuations in the phase gradient. The second factor (i.e. the integral over the impact parameter $b$ and the incoming momentum direction $\mathbf{n}$ ) determines the scale for $\rho(E)$, essentially by counting the number of sign-changing trajectories, and is expected to be only weakly temperature dependent. In the next section, we compute the linewidth within a Berezinskiil-Kosterlitz-Thouless-like model of phase fluctuations near $T_{\mathrm{c}}$.

### 10.4.3 Linewidth near the superconducting phase boundary

To make further progress, we now attempt to calculate the spectral function linewidth due to the phase fluctuations accompanying the destruction of superconducting order in the neighborhood of $T_{\mathrm{c}}$. We shall do this by choosing a particular weight for the phase fluctuations, viz., that associated with the two-dimensional $X Y$-model [106108]. Our analysis is reminiscent of that due to Franz and Millis [84] who addressed the bulk single-particle spectral function in the pseudogap regime. In the present context, this analysis is provided solely for illustrative purposes, and is only meant only to provide a rough estimate of the linewidth.

Let us consider the $X Y$-model action,

$$
\begin{equation*}
S[\Delta]=\frac{K}{2} \int d^{2} r|\nabla \vartheta|^{2} \tag{10.31}
\end{equation*}
$$

where $K \equiv \rho_{s}(T) / T$ in which $\rho_{s}(T)$ is the temperature-dependent superfluid density [109]. Note that we are not suggesting that the true critical fluctuations of the cuprate superconductors necessarily lie in the universality class of the two-dimensional $X Y$-model, but simply that the intermediate length-scale fluctuations proposed as leading to pseudogap phenomena may adequately be modeled by Eq. (10.31). The form for $\left.\left.\langle | \nabla \vartheta(\mathbf{r})\right|^{2}\right\rangle$ for the $X Y$-model may be calculated following the Debye-Hückeltype analysis of Halperin and Nelson [111], giving

$$
\begin{equation*}
\left.\left.\langle | \nabla \vartheta(\mathbf{r})\right|^{2}\right\rangle \simeq \frac{2}{\pi \xi_{\vartheta}^{2} K^{*}} \ln \Lambda \xi_{\vartheta}, \tag{10.32}
\end{equation*}
$$

in which $K^{*}$ is the short length-scale stiffness (obtained using the Kosterlitz renormalizationgroup equations), $\Lambda$ is a short-distance cutoff, and $\xi_{\vartheta}$ is a length-scale characterizing the typical inter-vortex spacing. The principal temperature dependence in Eq. (10.32) arises via $\xi_{\vartheta}$ : near $T_{\mathrm{BKT}}$ [i.e. the transition temperature of the model (10.31), which is expected to lie not far below $\left.T_{\mathrm{c}}\right] \xi_{\vartheta}$ is proportional to $\exp \sqrt{\Theta /\left(T-T_{\mathrm{BKT}}\right)}$, where $\Theta$ is a constant of order unity (which we do not try to calculate in detail).


Figure 10.2: Spectral linewidth versus $\left(T-T_{\mathrm{BKT}}\right) / \Theta$. Inset: Immediate vicinity of the transition.

The proliferation of unbound vortex excitations upon warming through the BKT transition is reflected by this strong diminution of $\xi_{\vartheta}$ and causes a concomitant dramatic increase in the linewidth of the spectral function:

$$
\begin{equation*}
\left\langle\epsilon^{2}\right\rangle^{1 / 2} \propto \exp \left(-\sqrt{\Theta /\left(T-T_{\mathrm{BKT}}\right)}\right) \tag{10.33}
\end{equation*}
$$

Observation of a fluctuation-broadened peak in the spectral function, e.g., via STS measurements, would provide striking evidence in support of the phase-fluctuation scenario. Moreover, the temperature dependence of the linewidth would, e.g., provide access to the details of the vortex-unbinding transition.

### 10.5 Numerical estimate of the linewidth; Concluding remarks

In the previous section we saw how one could estimate the temperature dependence of the linewidth of the spectral function near the superconducting transition temperature $T_{\mathrm{c}}$. In the present section we make a rough numerical estimate of this linewidth at one particular temperature in the pseudogap regime by appealing to the data obtained in the high-frequency conductivity experiments of Corson et al. [87] on $\mathrm{Bi}_{2} \mathrm{Sr}_{2} \mathrm{CaCu}_{2} \mathrm{O}_{8+\delta}$. We shall be specifically concerned with the interpretation of these data inasmuch as they provide access to the characteristic vortex density.

Until now we have been working with a system of units in which $\hbar^{2} / 2 m=1$. Restoring conventional units in Eq. (10.29) gives for the linewidth

$$
\begin{equation*}
\left.\left.\left\langle\epsilon^{2}\right\rangle^{1 / 2} \sim \frac{\hbar^{2}}{2 m} k_{\mathrm{F}}\langle | \boldsymbol{\nabla} \vartheta(\mathbf{r})\right|^{2} / 2\right\rangle^{1 / 2} \tag{10.34}
\end{equation*}
$$

To estimate this width, we turn to the Corson et al. experiments and its analysis by Corson et al., which is based on the notion that, at sufficiently high frequencies, the conductivity probes short-length-scale pairing correlations, and leads to an estimate for the characteristic density of free vortices $n_{\mathrm{f}}$. Assuming that is is vortex excitations that lead to phase fluctuations, one expects that, up to a constant of order unity, $\left.\left.\langle | \boldsymbol{\nabla} \vartheta(\mathbf{r})\right|^{2}\right\rangle \simeq n_{\mathrm{f}}$.

Now, Corson et al. obtain values of $n_{f}$ that are on the order of $0.003 a_{\mathrm{vc}}$ (for $T \sim 75 K$ ), where $a_{\mathrm{vc}}$ is the area of the core of a vortex. If we take the vortex core to
be a disk of radius $\xi \simeq 1 \mathrm{~nm}$, this leads to the value $n_{\mathrm{f}} \sim 10^{3} \mu \mathrm{~m}^{-2}$. Then, using the order-of-magnitude estimate $k_{\mathrm{F}} \simeq 1 \mathrm{~nm}^{-1}$, Eq. (10.34) gives $\left\langle\epsilon^{2}\right\rangle^{1 / 2} \simeq 9 \mathrm{meV}$.

The value of this estimate is that it shows that, for at least one cuprate material, there is a temperature at which the free-vortex density is small enough that phase fluctuations only weakly perturb the energies of the quasiparticle states. Hence, the linewidth arising from perturbed zero-energy states can be rather smaller than the scale of the superconducting energy gap (and hence small enough to justify the our picture of perturbed zero-modes) but be large enough to be resolvable in STS measurements, such as those of Yazdani et al. [68]. Larger densities of free vortices, and hence large line-widths, would result from working at higher temperatures. Moreover, near $T_{\mathrm{c}}$ the vortex density is expected to show a very strong dependence on temperature, which should confer a strong temperature dependence on the linewidth. Of course, if conducted at the temperature of the Corson et al. experiments, thermal broadening would complicate the task of accessing the intrinsic linewidth (i.e. the linewidth due to phase fluctuations); recall that 10 K is equivalent to 1 meV . Thus, as emphasized in Sec. 10.1, one should consider performing experiments on materials having a lower $T_{\mathrm{c}}$, so that the pseudogap regime can be explored at temperatures at which thermal broadening is less significant. These considerations indicate that it is at least conceivable that STS experiments near extended scatterers could provide a sharp test of the phase fluctuation scenario for the pseudogap regime.

## Appendix A

## Application of boundary integral technique to the Helmholtz and Bogoliubov-De Gennes wave equations

The underlying strategy employed in this Paper is the boundary integral technique [45], the origin of which was Fredholm's analysis of the existence of soluions of the interior Laplace problem subject to Dirichlet boundary conditions. In this scheme, Fredholm transformed the task of solving the Laplace partial differential equation (subject to Dirichlet boundary conditions) to one of solving a certain integral equation residing on the boundary. This prompted Fredholm to develop the theory of what are now known as Fredholm integral equations and, in particular, to prove the existence of a solution of the corresponding Laplace problem.

In the present context of spectral geometry, the virtue of this boundary integral technique is that it allows one to harness the piecewise homogeneity of the system (and the corresponding simplicity of the fundamental Green functions in the locally homogeneous regions) and, thereby, to study the physical implications of the bound-
ary in as direct and a natural manner as possible.
The aim of this appendix is to provide a guide to the boundary integral technique, beginning with the simplest setting and working towards the setting of the BDG eigenproblem. When discussing the simplest settings we shall borrow heavily from Refs. [46, 47]. The elaborations that we shall be needing for the BDG setting arise from (i) the multicomponent nature of the eigenproblem, and (ii) the presence of matching rather than boundary conditions. We mention that Jackson [48] gives a highly readable discussion of the physics of the potential discontinuities that are a pivotal feature of boundary integral techniques.

## A. 1 Review of elementary ingredients

In this section we shall discuss the origin of the discontinuities in the three-dimensional potential and the fields generated by surface charge (which we call single-layer) and dipole (which we call double-layer) densities, as well as present derivations of explicit formulas quantifying such discontinuities. Such formulas will become useful in the next section, where we discuss parametrizations of wave functions in terms of these single and double layers.

Before turning to the derivation, we consider a simple example which contains the essential features: a planar charge layer with constant charge density $\nu$. Without loss of generality, let us assume that the charge layer lies in the $x y$ plane, so that the normal direction nis $\hat{\mathbf{z}}$. Then the potential $\varphi$ and the field $\mathbf{E}$ are given by the following surface integrals:

$$
\begin{align*}
4 \pi \varphi(\mathbf{x}) & =\nu \int d \sigma_{\alpha} \frac{1}{|\mathbf{x}-\boldsymbol{\alpha}|}  \tag{A.1}\\
4 \pi \mathbf{E}(\mathbf{x}) & =-\nu \int d \sigma_{\alpha} \boldsymbol{\nabla}_{x}\left(\frac{1}{|\mathbf{x}-\boldsymbol{\alpha}|}\right)=\nu \int d \sigma_{\alpha} \frac{\mathbf{x}-\boldsymbol{\alpha}}{|\mathbf{x}-\boldsymbol{\alpha}|^{3}} \tag{A.2}
\end{align*}
$$

where $\mathbf{x} \equiv(x, y, z), \boldsymbol{\alpha} \equiv\left(x^{\prime}, y^{\prime}, 0\right)$, and $d \sigma_{\alpha} \equiv d x^{\prime} d y^{\prime}$. Owing to translational invariance in the $x y$ plane, $\varphi(\mathbf{x})=\varphi(z)$ and $\mathbf{E}(\mathbf{x})=\hat{\mathbf{z}} E(z)$. Without loss of generality,
let us choose $x=y=0$, and focus on $E(z)$. By introducing the polar coordinates $\left(x^{\prime}, y^{\prime}\right)=(r \cos \theta, r \sin \theta)$, the integral for $E(z)$ becomes

$$
\begin{equation*}
E(z)=\nu \int r d r d \theta \frac{z}{4 \pi\left(z^{2}+r^{2}\right)^{3 / 2}}=\frac{\nu}{2} \int_{0}^{\infty} r d r \frac{z}{\left(z^{2}+r^{2}\right)^{3 / 2}}=\frac{\nu}{2} \operatorname{sgn}(z) \tag{A.3}
\end{equation*}
$$

For $z>0, E(z)$ is independent of $z$ and equal to $\nu / 2$; for $z<0, E(z)$ is independent of $z$, however it is equal to $-\nu / 2$ and for $z=0, E(z)$ vanishes. The mathematical origin of this discontinuity is the noncommutativity of the limit $z \rightarrow 0$ and the surface integral in Eq. (A.3). Thus the value of $\varphi$ is continuous as it approaches the surface but its normal derivative (in this case partial derivative with respect to $z$ ) is not. This discontinuity can be summarized by the equation

$$
\begin{align*}
& \lim _{z \rightarrow 0^{+}} \varphi=\left.\varphi\right|_{z=0}, \quad \lim _{z \rightarrow 0^{-}} \varphi=\left.\varphi\right|_{z=0},  \tag{A.4}\\
& \lim _{z \rightarrow 0^{+}} \frac{\partial \varphi}{\partial z}=-\frac{\nu}{2}+\left.\frac{\partial \varphi}{\partial z}\right|_{z=0^{\prime}}, \quad \lim _{z \rightarrow 0^{-}} \frac{\partial \varphi}{\partial z}=\frac{\nu}{2}+\left.\frac{\partial \varphi}{\partial z}\right|_{z=0} . \tag{A.5}
\end{align*}
$$

Now consider a generic charge layer $\nu(\boldsymbol{\alpha})$ on the surface $\partial \mathcal{V}$, and focus on the potential generated by via the Helmholtz (instead of Coulomb) Green function

$$
\begin{equation*}
\varphi(\mathbf{x})=\int_{\partial \mathcal{V}} d \sigma_{\alpha} g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\alpha}) \nu(\boldsymbol{\alpha}) \tag{A.6}
\end{equation*}
$$

As in the case of the simple example of homogeneous planar charge layer, the potential generated by this generic charge layer is continuous,

$$
\begin{equation*}
\lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \varphi(\mathbf{x})=\int_{\partial \mathcal{V}} d \sigma_{\alpha} g^{\mathrm{H}}(\boldsymbol{\beta}, \boldsymbol{\alpha}) \nu(\boldsymbol{\alpha}) \tag{A.7}
\end{equation*}
$$

but its normal derivative is not. In order to see this, consider the normal derivative $\mathbf{n}_{\beta} \cdot \boldsymbol{\nabla}_{x} \varphi(\mathbf{x})$ as $\mathbf{x}$ in $\overline{\mathcal{V}}$ tends to a generic point on $\partial \mathcal{V}$, which we denote by $\boldsymbol{\beta}$, along the interior normal and divide the domain of the surface integration into two parts: (i) a small region $\mathcal{D}_{\delta} \equiv \mathcal{C}_{\delta} \bigcap \partial \mathcal{V}$, where $\mathcal{C}_{\delta}$ is a sphere of radius $\delta$ around $\boldsymbol{\beta}$ and (ii) remaining domain $\overline{\mathcal{D}}_{\delta} \equiv \partial \mathcal{V}-\mathcal{D}_{\delta}$, then

$$
\begin{align*}
& \lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \mathbf{n}_{\beta} \cdot \nabla_{x} \varphi(\mathbf{x})=\lim _{\epsilon \rightarrow 0} \int_{\mathcal{D}_{\delta}} d \sigma_{\alpha} \frac{\partial}{\partial \epsilon} g^{\mathrm{H}}\left(\boldsymbol{\beta}+\epsilon \mathbf{n}_{\beta}, \boldsymbol{\alpha}\right) \nu(\boldsymbol{\alpha}) \\
&+\lim _{\epsilon \rightarrow 0} \int_{\overline{\mathcal{D}}_{\delta}} d \sigma_{\alpha} \frac{\partial}{\partial \epsilon} g^{\mathrm{H}}\left(\boldsymbol{\beta}+\epsilon \mathbf{n}_{\beta}, \boldsymbol{\alpha}\right) \nu(\boldsymbol{\alpha}) \tag{A.8}
\end{align*}
$$

where $\epsilon$ is the perpendicular distance between $\mathbf{x}$ and $\boldsymbol{\beta}$. The virtue of this separation is that the singularity of $g^{\mathrm{H}}(\boldsymbol{\beta}, \boldsymbol{\alpha})$ at $\boldsymbol{\beta}=\boldsymbol{\alpha}$ is now contained in the first term on the right hand side of the Eq. (A.8). As the integrand of second term on the right hand side of this equation is free of singularities, the limit can be taken inside the integral sign. Moreover within $\mathcal{D}_{\delta}$ and for $\delta$ and $\epsilon$ very small,

$$
\begin{equation*}
\frac{\partial}{\partial \epsilon} g^{\mathrm{H}}\left(\boldsymbol{\beta}+\epsilon \mathbf{n}_{\beta}, \boldsymbol{\alpha}\right) \approx \frac{\epsilon}{4 \pi\left(\rho^{2}+\epsilon^{2}\right)^{3 / 2}}, \quad \rho \equiv|\boldsymbol{\alpha}-\boldsymbol{\beta}|^{2} \tag{A.9}
\end{equation*}
$$

Then the integral over the domain $\mathcal{D}_{\delta}$ can be evaluated for small $\delta$ as follows
$\lim _{\epsilon \rightarrow 0} \int_{\mathcal{D}_{\delta}} d \sigma_{\alpha} \frac{\partial}{\partial \epsilon} g^{\mathrm{H}}\left(\boldsymbol{\beta}+\epsilon \mathbf{n}_{\beta}, \boldsymbol{\alpha}\right) \nu(\boldsymbol{\alpha})=\lim _{\epsilon \rightarrow 0} \nu(\boldsymbol{\beta}) \int_{0}^{2 \pi} d \theta \int_{0}^{\delta} \rho d \rho \frac{\epsilon}{4 \pi\left(\rho^{2}+\epsilon^{2}\right)^{3 / 2}}=\frac{1}{2} \nu(\boldsymbol{\beta})$.

Now that we have evaluated the singular part of the surface integral we take the limit $\delta$ tends to zero:

$$
\begin{align*}
\lim _{\mathbf{x} \in \mathcal{V} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \mathbf{n}_{\beta} \cdot \boldsymbol{\nabla}_{x} \varphi(\mathbf{x}) & =\frac{1}{2} \nu(\boldsymbol{\beta})+\lim _{\delta \rightarrow 0} \int_{\overline{\mathcal{D}}_{\delta}} d \sigma_{\alpha} \mathbf{n}_{\beta} \cdot \boldsymbol{\nabla}_{\beta} g^{\mathrm{H}}(\boldsymbol{\beta}, \boldsymbol{\alpha}) \nu(\boldsymbol{\alpha})  \tag{A.11}\\
& =\frac{1}{2} \nu(\boldsymbol{\beta})+\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{n}_{\beta} \cdot \boldsymbol{\nabla}_{\beta} g^{\mathrm{H}}(\boldsymbol{\beta}, \boldsymbol{\alpha}) \nu(\boldsymbol{\alpha}) \tag{A.12}
\end{align*}
$$

We note that this discontinuity originates again from the noncommutativity of limit $\mathbf{x} \rightarrow \boldsymbol{\beta}$ and the surface integral. Moreover the calculation of the amount of this discontinuity involves only the form of $g^{\mathrm{H}}$ for small values of distance between its arguments and this form is simply the Coulomb Green function. Thus, in essence, this discontinuity is the same as the discontinuity in the simple case of a planar, electrostatic-charge layer discussed at the beginning of this section.

Let us now consider the potential generated by a generic dipole density $\mu(\boldsymbol{\alpha})$ on $\partial \mathcal{V}$, viz.

$$
\begin{equation*}
\varphi(\mathbf{x})=\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{n}_{\alpha} \cdot \nabla_{\alpha} g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\alpha}) \mu(\boldsymbol{\alpha}) \tag{A.13}
\end{equation*}
$$

Notice the similarity of this form to that of the normal derivative of the potential generated by the charge layer, only difference being the normal derivative acting on the second rather than the first argument. However, as the $\mathbf{G}^{H}$ is a function of the
difference between its arguments, one expects a similar discontinuity on the potential across the surface. Let us now consider the case in which $\mathbf{x}$ approaches to a surface point denoted by $\boldsymbol{\beta}$ from $\overline{\mathcal{V}}$. Indeed,

$$
\begin{align*}
\lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \varphi(\mathbf{x})= & \lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{n}_{\alpha} \cdot \boldsymbol{\nabla}_{\alpha} g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\alpha}) \mu(\boldsymbol{\alpha})  \tag{A.14}\\
= & \lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \int_{\mathcal{D}_{\delta}} d \sigma_{\alpha} \mathbf{n}_{\alpha} \cdot \boldsymbol{\nabla}_{\alpha} g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\alpha}) \mu(\boldsymbol{\alpha}) \\
& +\lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \int_{\overline{\mathcal{D}}_{\delta}} d \sigma_{\alpha} \mathbf{n}_{\alpha} \cdot \boldsymbol{\nabla}_{\alpha} g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\alpha}) \mu(\boldsymbol{\alpha})  \tag{A.15}\\
= & -\lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \int_{\mathcal{D}_{\delta}} d \sigma_{\alpha} \mathbf{n}_{\beta} \cdot \boldsymbol{\nabla}_{x} g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\alpha}) \mu(\boldsymbol{\alpha}) \\
& \quad \lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \int_{\overline{\mathcal{D}}_{\delta}} d \sigma_{\alpha} \mathbf{n}_{\alpha} \cdot \boldsymbol{\nabla}_{\alpha} g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\alpha}) \mu(\boldsymbol{\alpha}), \tag{A.16}
\end{align*}
$$

where in order to get to the third line we have used $\boldsymbol{\nabla}_{x} g^{\mathrm{H}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=-\boldsymbol{\nabla}_{x^{\prime}} g^{\mathrm{H}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$, and that in $\mathcal{D}_{\delta}$, as $\delta$ goes to zero, $\mathbf{n}_{\alpha} \rightarrow \mathbf{n}_{\beta}$. Notice that the first term in the right hand side of Eq. (A.16) is equal to the first term in Eq. (A.10). Thus in the limit $\delta$ goes to zero, we have:

$$
\begin{equation*}
\lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \varphi(\mathbf{x})=-\frac{1}{2} \mu(\boldsymbol{\beta})+\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{n}_{\alpha} \cdot \boldsymbol{\nabla}_{\alpha} g^{\mathrm{H}}(\boldsymbol{\beta}, \boldsymbol{\alpha}) \mu(\boldsymbol{\alpha}) . \tag{A.17}
\end{equation*}
$$

The discontinuity in $\varphi$ as $\mathbf{x}$ approaches to $\boldsymbol{\beta}$ from $\mathcal{V}$ (rather than $\overline{\mathcal{V}}$ ), is obtained similarly, the result is:

$$
\begin{equation*}
\lim _{\mathbf{x} \in \mathcal{V} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \varphi(\mathbf{x})=\frac{1}{2} \mu(\boldsymbol{\beta})+\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{n}_{\alpha} \cdot \boldsymbol{\nabla}_{\alpha} g^{\mathrm{H}}(\boldsymbol{\beta}, \boldsymbol{\alpha}) \mu(\boldsymbol{\alpha}) \tag{A.18}
\end{equation*}
$$

## A. 2 Single- and double-layer parametrizations of wave functions; Jump conditions

The first ingredient needed for the construction of a MSE is the parametrization of a wave function in terms of single or double layers. As this aspect of classical potential theory might be unfamiliar to some readers, we first illustrate it in the simpler setting of the Helmholtz wave equation, before turning to the BDG wave equation.

## A. 3 Parametrizations for scalar Helmholtz wave functions

Consider the Helmholtz wave equation,

$$
\begin{equation*}
\left(\nabla^{2}+E\right) \varphi(\mathbf{x})=0, \tag{A.19}
\end{equation*}
$$

for the wave function $\varphi(\mathbf{x})$ in the region $\overline{\mathcal{V}}$ (which we define to be the region outside some region $\mathcal{V}$ bounded by the closed surface $\partial \mathcal{V}$ ). Then, from potential theory [46, 47] it is known that the solutions $\varphi(\mathbf{x})$ can be parametrized in terms of a function $\boldsymbol{\nu}(\boldsymbol{\alpha})$ defined only on $\partial \mathcal{V}$, via the integral

$$
\begin{equation*}
\varphi(\mathrm{x})=\int_{\partial \nu} d \sigma_{\alpha} g^{\mathrm{H}}(\mathbf{x}-\boldsymbol{\alpha}) \nu(\boldsymbol{\alpha}) . \tag{A.20}
\end{equation*}
$$

Here, $\mathbf{x}$ and $\mathrm{x}^{\prime}$ are positions lying in $\overline{\mathcal{V}}$, Greek letters, such as $\boldsymbol{\alpha}$, represent vectors on the boundary $\partial \mathcal{V}$ (as they do throughout this Paper), and $g^{\mathrm{H}}\left(\mathrm{x}-\mathrm{x}^{\prime}\right)$ is the fundamental Green function for the Helmholtz wave equation, which satisfies

$$
\begin{equation*}
\left(\nabla^{2}+E\right) g^{\mathrm{H}}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)=\delta^{(3)}\left(\mathbf{x}-\mathbf{x}^{\prime}\right) . \tag{A.21}
\end{equation*}
$$

One can interpret the parametrization by saying that the wave function $\varphi(\mathbf{x})$ is the Helmholtz wave function due to a single layer of charge of surface density $\nu(\boldsymbol{\alpha})$.

Where does this parametrization come from? First, note that the wave function (A.20) does indeed satisfy Eq. (A.19). To see this, observe that $\mathbf{x}$ and $\boldsymbol{\alpha}$ are never coincident ( $\boldsymbol{\alpha}$ lying on $\partial \mathcal{V}$ but $\mathbf{x}$ lying in $\overline{\mathcal{V}}$ ) so that $g^{\mathrm{H}}(\mathbf{x}-\boldsymbol{\alpha})$ solves Eq. (A.19) for any $\boldsymbol{\alpha}$, and the parametrization A. 20 is simply a superposition of such solutions. Second, recall that by Green's theorem one has

$$
\begin{equation*}
\varphi(\mathbf{x})=\int_{\partial V} d \sigma_{\alpha} \partial_{\alpha} g^{A}(\mathbf{x}, \boldsymbol{\alpha}) \varphi(\boldsymbol{\alpha})-\int_{\partial V} d \sigma_{\alpha} g^{A}(\mathbf{x}, \boldsymbol{\alpha}) \partial_{\alpha} \varphi(\boldsymbol{\alpha}), \tag{A.22}
\end{equation*}
$$

where $g^{A}(\mathbf{x}, \boldsymbol{\alpha})$ is any Helmhotz Green function (i.e. not necessarily the fundamental one). In the most common setting, one then chooses the Green function that satisfies
the homogeneous version of the boundary condition on $\varphi$, thus eliminating all absent boundary information and arriving at an expression for $\varphi(\mathbf{x})$ in terms of information known about $\varphi$ on $\partial \mathcal{V}$. Here, instead, one takes a different tack. One selects for $g^{\text {A }}$ the fundamental Green function $g^{\mathrm{H}}$, jettisons the first contribution to the right hand side of Eq. (A.22), and accommodates for this by replacing the boundary information on $\partial \varphi$ by the (as-yet unknown) single-layer $\nu(\boldsymbol{\alpha})$.

Representations of wave functions by surface integrals are available in other settings, too. We have considered wave functions satisfying the Helmholtz wave equation outside the region $\mathcal{V}$ (i.e. the so-called exterior problem). One can also consider wave functions satisfying the Helmholtz wave equation inside the region $\mathcal{V}$ (i.e. the so-called interior problem).

Furthermore, one can parametrize wave functions satisfying the Helmholtz wave equation in other ways. For example, consider wave functions inside the region $\mathcal{V}$, which one can parametrize as

$$
\begin{equation*}
\varphi(\mathbf{x})=\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} g^{\mathrm{H}}(\mathbf{x}-\boldsymbol{\alpha}) \mu(\boldsymbol{\alpha}) \tag{A.23}
\end{equation*}
$$

In this case $\varphi(\mathbf{x})$ is the Helmholtz wave function due to a layer of dipoles on $\partial \mathcal{V}$ of local strength $\mu(\boldsymbol{\alpha})$ and local orientation normal to $\partial \mathcal{V}$ at each point $\boldsymbol{\alpha}$. (We denote such surface normal vectors as $\mathbf{n}_{\alpha}$, and adopt the convention that they point towards the interior of $\mathcal{V}$.) One can, of course, regard this dipole layer as consisting of two single layers, vanishingly close to one another and locally carrying opposite charges, in the limit that the charges become large and the layer separation becomes correspondingly small. Such layers are referred to as a double layers. The normal component of the gradient acting on the Green function accounts for the fact that this parametrization features opposing, vanishingly close, layers. As with the single-layer parametrization (A.20), that the parametrization (A.23) satisfies Eq. (A.19) follows because $\mathbf{x}$ and $\boldsymbol{\alpha}$ are never coincident, so that $g^{\mathrm{H}}(\mathbf{x}-\boldsymbol{\alpha})$ solves Eq. (A.19) for any $\boldsymbol{\alpha}$. Motivation for the parametrization (A.23) also follows from consideration of Green's
theorem, Eq. (A.22), but with the second contribution on the right hand side being jettisoned, rather than the first.

There, of course, remains the issue that whether all solutions can be expressed in terms of these parametrizations. It turns out that if one uses the double-layer parametrization for interior wave functions, and the single-layer parametrization for exterior wave functions then any solution can be thus parametrized. The converse problem (i.e. parametrizing interior wave functions using double layers and exterior wave functions using single layers) is also possible, provided certain supplementary conditions are satisfied; see, e.g., Refs. [46, 47]

We note that the strategy that we are adopting can be implemented in more general settings. For example, one might consider the case of disconnected superconducting regions connected by normal regions, and thus address the issue of Josephson tunneling between them. One might also consider disconnected normal regions connected by superconducting regions, and thus address the issue of single-particle tunneling between them.

The utility of these parametrizations is that they can be used to transform partial differential equations in $\mathcal{V}$ or $\overline{\mathcal{V}}$ into Fredholm integral equations that reside on $\partial \mathcal{V}$, and as we shall see in the following section, such integral equations prove useful in some cases, especially if the integral equation is of the second type, i.e., if an iterative solution is possible.

We now turn to the second ingredient needed for the construction of a MSE, viz., jump conditions. It is an important result of potential theory that the parametrization of wave functions in terms of single and double layers, such as those given in Eqs. (A.20,A.23), leads to representations of wave functions that behave in a singular fashion for field points $\mathbf{x}$ on the surface $\partial \mathcal{V}$. It is precisely this singular behavior, and the attendant jump conditions, that are responsible for the utility of these parametrizations and, as we shall see shortly, lead to the formulation of integral equa-
tions for the layer strengths, single or double, known as boundary integral equations. These equations incorporate what ever boundary conditions one wishes to impose on the wave function. By solving boundary integral equations one arrives at layer strengths that parametrize the wave functions.

There are three virtues to this boundary integral equations formulation. First, the boundary integral equations reside solely on the boundary $\partial \mathcal{V}$ (i.e. on a manifold of dimension one fewer than the original wave equation). From the computational standpoint it is economical to formulate a problem in terms of functions that reside on lower-dimensional manifolds (and hence depend on fewer variables). Second, from the theoretical standpoint, existence theorems have been established for broad classes of integral equations, often encompassing the boundary integral equations that emerge from specific examples. (In fact, it was the goal of establishing exitence theorems for solutions to the Laplace equation in various settings - interior or exterior, Dirichlet or Neumann boundary conditions - that inspired Fredholm to develop the boundary intergal equation approach to potential theory, and subsequently to develop the theory of what we now know as Fredholm integral equations.) And third, from the physical standpoint, boundary integral equations and their iterative solution allow one to organize the computation of wave functions in terms of the multiple scattering of waves from interfaces that separate spatially homogeneous regions, along with free propagation between those scattering events. Thus, one is in a position to focus on the boundary scattering events, and thereby to focus on the geometry of the boundary and the implications of its shape for the physical problem at hand. In essence, we are invoking the piecewise homogeneity of the system to "integrate up" our description of it, leaving us with the need to consider one fewer independent variable. As a result of this "integrating up", we depart from a purely local description, in terms of differential equations, and arrive at a nonlocal formulation in terms of integral equations.

## A.3.1 Jump conditions for one-component Helmholtz wave functions

Consider the single-layer parametrization for Helmholtz wave functions in the region $\overline{\mathcal{V}}$, given by Eq. (A.20). Now, it is known from potential theory that this parametrization is singular as $\mathbf{x}$ goes to any value $\boldsymbol{\beta}$ on the boundary (see App. A.2). Specifically, $\varphi(\mathbf{x})$ is continuous whereas $\mathbf{n}_{\beta} \cdot \nabla_{x} \varphi(\mathbf{x})$ is not, but the discontinuity of the latter has a known and useful form:

$$
\begin{align*}
\lim _{\mathbf{x} \in \mathcal{V} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \int_{\partial \mathcal{V}} d \sigma_{\alpha} g^{\mathrm{H}}(\mathbf{x}-\boldsymbol{\alpha}) \nu(\boldsymbol{\alpha}) & =\int_{\partial \mathcal{V}} d \sigma_{\alpha} g^{\mathrm{H}}(\boldsymbol{\beta}-\boldsymbol{\alpha}) \nu(\boldsymbol{\alpha}) ;  \tag{A.24}\\
\lim _{\mathbf{x} \in \mathcal{V} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \mathbf{n}_{\beta} \cdot \nabla_{x} \int_{\partial \mathcal{V}} d \sigma_{\alpha} g^{\mathrm{H}}(\mathbf{x}-\boldsymbol{\alpha}) \nu(\boldsymbol{\alpha}) & =  \tag{A.25}\\
& \frac{1}{2} \nu(\boldsymbol{\beta})+\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{n}_{\beta} \cdot \nabla_{\beta} g^{\mathrm{H}}(\boldsymbol{\beta}-\boldsymbol{\alpha}) \nu(\boldsymbol{\alpha}) .
\end{align*}
$$

Now consider the double-layer parametrization of the Helmholtz wave functions on the region $\mathcal{V}$, given by Eq. (A.23). It is known from potential theory that this parametrization is also singular as $\mathbf{x}$ goes to any value of $\boldsymbol{\beta}$ on the boundary (see App. A.2). However, in this case $\varphi(\mathbf{x})$ itself is discontinuous, and $\mathbf{n}_{\beta} \cdot \nabla_{x} \varphi(\mathbf{x})$ is even more singular, the discontinuity of $\varphi(\mathbf{x})$ being given by

$$
\begin{equation*}
\lim _{\mathbf{x} \in \mathcal{V} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} g^{\mathrm{H}}(\mathbf{x}-\boldsymbol{\alpha}) \mu(\boldsymbol{\alpha})=\frac{1}{2} \mu(\boldsymbol{\beta})+\int_{\partial \mathcal{V}} d \sigma_{\alpha} g^{\mathrm{H}}(\boldsymbol{\beta}-\boldsymbol{\alpha}) \mu(\boldsymbol{\alpha}) \tag{A.26}
\end{equation*}
$$

Although $\mathbf{n}_{\beta} \cdot \nabla_{x} \varphi(\mathbf{x})$ diverges on $\partial \mathcal{V}$, it is yet a further result from potential theory that the values of the limits of this quantitity, as $\mathbf{x}$ approaches any point $\boldsymbol{\beta}$ on $\partial \mathcal{V}$ either from $\mathcal{V}$ or from $\overline{\mathcal{V}}$, are equal to one another. It will, therefore, prove to be convenient to redefine the quantity $\left.\mathbf{n}_{\beta} \cdot \nabla_{x} \varphi(\mathbf{x})\right|_{\mathbf{x}=\beta}$ as its limiting value:

$$
\begin{equation*}
\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\beta}^{+} \partial_{\alpha} g^{\mathrm{H}}(\boldsymbol{\beta}-\boldsymbol{\alpha}) \nu(\boldsymbol{\alpha}) \equiv \lim _{\mathbf{x} \in \mathcal{V} \text { or } \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \mathbf{n}_{\beta} \cdot \nabla_{x} \int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} g^{\mathrm{H}}(\boldsymbol{\beta}-\boldsymbol{\alpha}) \nu(\boldsymbol{\alpha}) . \tag{A.27}
\end{equation*}
$$

Via this definition the normal derivative $\mathbf{n}_{\beta} \cdot \nabla_{x} \varphi(\mathbf{x})$ is continuous across $\partial \mathcal{V}$.

## A.3.2 Parametrizations for two-component <br> Bogoliubov-De Gennes wave functions

We now turn to the issue of parametrizing solutions of the BDG equation (3.6) in terms of single and double layers, bearing in mind the two-component nature of the wave functions. For the sake of concreteness, as well as experimental relevance, we focus on the setting of an Andreev billiard, so that space is partitioned into two regions by the surface $\partial \mathcal{V}$, with $\Delta(\mathrm{x})=0$ in the region $\mathcal{V}$ inside $\partial \mathcal{V}$ (i.e. the normal-metal region) and $\Delta(\mathbf{x})=\Delta_{0}$ in the region $\overline{\mathcal{V}}$ outside $\partial \mathcal{V}$ (i.e., the superconducting region). However, the are no obstacles of principle in applying the present techniques in other settings, provided they comprise regions of space in which $\Delta(\mathrm{x})$ is constant, separated by surfaces on which $\Delta(\mathrm{x})$ varies discontinuously [i.e. $\Delta(\mathrm{x})$ must be piecewise constant]. The present techniques may also be applied in settings in which other physical parameters vary in a piecewise continuous fashion.

First, consider the normal-state interior of an Andreev billiard [i.e. a region $\mathcal{V}$ surrounded by a surface $\partial \mathcal{V}$ in which $\Delta(\mathrm{x})=0$ ] in which the BDG wave functions satisfy

$$
\left(\begin{array}{cc}
\nabla^{2}-\mu+E & 0  \tag{A.28}\\
0 & -\nabla^{2}+\mu+E
\end{array}\right)\binom{u(\mathbf{x})}{v(\mathbf{x})}=\binom{0}{0} .
$$

Solutions of this equation can be parametrized in terms of the two-component double layer $\boldsymbol{\mu}(\boldsymbol{\alpha})$ in the following way:

$$
\begin{equation*}
\boldsymbol{\Phi}^{\mathrm{N}}(\mathbf{x})=\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\mathbf{x}-\boldsymbol{\alpha}) \cdot \boldsymbol{\mu}(\boldsymbol{\alpha}) \tag{A.29}
\end{equation*}
$$

Here, $\mathbf{G}^{\mathrm{N}}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)$ is the fundamental Green function for the BDG wave equation in the absence of a pair potential, and is given explicitly by Eq. (3.21). Such solutions could also be parametrized in terms of the two-component single layer $\boldsymbol{\nu}(\boldsymbol{\alpha})$, as

$$
\begin{equation*}
\boldsymbol{\Phi}^{\mathrm{N}}(\mathbf{x})=\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{G}^{\mathrm{N}}(\mathbf{x}-\boldsymbol{\alpha}) \cdot \boldsymbol{\nu}(\boldsymbol{\alpha}) \tag{A.30}
\end{equation*}
$$

but it will prove convenient to adopt the former parametrization, Eq. (A.29), rather than the latter, Eq. (A.30).

The two-component layers, $\boldsymbol{\mu}(\boldsymbol{\alpha})$ and $\boldsymbol{\nu}(\boldsymbol{\alpha})$, reflect the two-component (i.e. electron and hole) nature of the wave functions. If one were solely concerned with the case of a normal region, this two-component description would be redundant: as no matrix elements of the Hamiltonian would connect upper and lower components there would be no need to adopt a language that embraces wave functions that describe coherent superpositions of electron and hole states. However, in an Andreev billiard the normal region is surrounded by a superconductor, the pair potential of which provides precisely the matrix element connecting electron and hole wave functions. Therefore it is necessary to adopt this two-component language for the normal region.

In the superconducting exterior of the Andreev billiard (i.e. the region $\overline{\mathcal{V}}$ outside the surface $\partial \mathcal{V}$ in which $\left.\Delta(\mathrm{x})=\Delta_{0}\right)$ the BDG wave functions satisfy

$$
\left(\begin{array}{cc}
\nabla^{2}-\mu+E & -i \Delta_{0}  \tag{A.31}\\
i \Delta_{0} & -\nabla^{2}+\mu+E
\end{array}\right)\binom{u(\mathbf{x})}{v(\mathbf{x})}=\binom{0}{0} .
$$

Solutions of this equation can be parametrized in terms of of the two-component single layer $\boldsymbol{\nu}(\boldsymbol{\alpha})$ in the following way:

$$
\begin{equation*}
\Psi^{\mathrm{S}}(\mathbf{x}) \equiv\binom{u(\mathbf{x})}{v(\mathbf{x})}=\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{G}^{\mathrm{S}}(\mathbf{x}-\boldsymbol{\alpha}) \cdot \boldsymbol{\nu}(\boldsymbol{\alpha}) \tag{А.32}
\end{equation*}
$$

where $G^{S}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)$ is the fundamental Green function for the BDG wave equation in the presence of a homogeneous pair potential $\Delta_{0}$, and is given explicitly by Eq. (3.30). Two-component layers are mandatory here, inasmuch as each component of the wave function is determined by both components of a layer, owing to the presence of the pair-potential. Such solutions could also be parametrized in terms of the twocomponent double layer $\boldsymbol{\mu}(\boldsymbol{\alpha})$, as

$$
\begin{equation*}
\Psi^{\mathrm{S}}(\mathbf{x}) \equiv\binom{u(\mathbf{x})}{v(\mathbf{x})}=\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} \mathrm{G}^{\mathrm{S}}(\mathbf{x}-\boldsymbol{\alpha}) \cdot \boldsymbol{\mu}(\boldsymbol{\alpha}) \tag{А.33}
\end{equation*}
$$

except that under certain circumstances the parametrization must be augmented by an additional term; see, e.g., Ref. [47]. However, it will prove adequate for us to stick with the former parametrization.

## A.3.3 Jump conditions for two-component <br> Bogoliubov-De Gennes wave functions

We now echo for the case of BDG wave functions the discussion, given in App. A.3.1, of the behavior of single- and double-layer parametrizations of Helmholtz wave functions in the vicinity of the surface $\partial \mathcal{V}$. Consider the double-layer parametrization of the BDG wave functions on the region $\mathcal{V}$ given by Eq. (A.29). This parametrization is singular, as $\mathbf{x}$ goes to any point $\boldsymbol{\beta}$ on the boundary (see App. A.2), inasmuch as $\boldsymbol{\Phi}^{\mathrm{N}}(\mathbf{x})$ is discontinuous, and $\mathbf{n}_{\beta} \cdot \nabla_{x} \boldsymbol{\Phi}^{\mathrm{N}}(\mathbf{x})$ is even more singular. From the form of $\mathbf{G}^{\mathrm{N}}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)$, Eq. (3.23), it can be shown that the discontinuity of $\boldsymbol{\Phi}^{\mathrm{N}}(\mathbf{x})$ is given by

$$
\begin{align*}
& \lim _{\mathbf{x} \in \mathcal{V} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\mathbf{x}-\boldsymbol{\alpha}) \cdot \boldsymbol{\mu}(\boldsymbol{\alpha}) \\
&=\frac{1}{2} \boldsymbol{\sigma}_{3} \cdot \boldsymbol{\mu}(\boldsymbol{\beta})+\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\boldsymbol{\beta}-\boldsymbol{\alpha}) \cdot \boldsymbol{\mu}(\boldsymbol{\alpha}) . \tag{A.34}
\end{align*}
$$

Although $\mathbf{n}_{\beta} \cdot \nabla_{x} \boldsymbol{\Phi}^{\mathrm{N}}(\mathbf{x})$ diverges on $\partial \mathcal{V}$, it can also be shown that that the values of the limits of this quantitity, as $\mathbf{x}$ approaches any point $\boldsymbol{\beta}$ on $\partial \mathcal{V}$ either from $\mathcal{V}$ or from $\overline{\mathcal{V}}$, are equal to one another. It will, therefore, prove to be convenient to redefine the quantity $\left.\mathbf{n}_{\beta} \cdot \nabla_{x} \boldsymbol{\Phi}^{\mathrm{N}}(\mathbf{x})\right|_{\mathbf{x}=\beta}$ as its limiting value:

$$
\begin{align*}
\int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\beta}^{+} \partial_{\alpha} & \mathbf{G}^{\mathrm{N}}(\boldsymbol{\beta}-\boldsymbol{\alpha}) \cdot \boldsymbol{\mu}(\boldsymbol{\alpha}) \\
& \equiv \lim _{\mathbf{x} \in \mathcal{V} \text { or } \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \mathbf{n}_{\beta} \cdot \nabla_{x} \int_{\partial \mathcal{V}} d \sigma_{\alpha} \partial_{\alpha} \mathbf{G}^{\mathrm{N}}(\mathbf{x}-\boldsymbol{\alpha}) \cdot \boldsymbol{\mu}(\boldsymbol{\alpha}) . \tag{A.35}
\end{align*}
$$

Now consider the single-layer parametrization of the BDG wave functions in the region $\overline{\mathcal{V}}$ given by Eq. (A.30). This parametrization is singular, as $\mathbf{x} \in \overline{\mathcal{V}}$ goes to any value of $\boldsymbol{\beta} \in \partial \mathcal{V}$ (see App. A.2). Specifically, $\boldsymbol{\Phi}^{\mathrm{S}}(\mathbf{x})$ is continuous whereas $\mathbf{n}_{\beta} \cdot \nabla_{x} \boldsymbol{\Phi}^{\mathrm{S}}(\mathbf{x})$
is not, but the discontinuity has a known and useful form. Indeed, from the form of $\mathbf{G}^{S}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)$, Eq. (3.30), it can be shown that

$$
\begin{align*}
& \lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathrm{G}^{\mathrm{S}}(\mathbf{x}-\boldsymbol{\alpha}) \cdot \boldsymbol{\nu}(\boldsymbol{\alpha})=\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{G}^{\mathrm{S}}(\boldsymbol{\beta}-\boldsymbol{\alpha}) \cdot \boldsymbol{\nu}(\boldsymbol{\alpha}) ;  \tag{A.36}\\
& \lim _{\mathbf{x} \in \overline{\mathcal{V}} \rightarrow \boldsymbol{\beta} \in \partial \mathcal{V}} \mathbf{n}_{\beta} \cdot \nabla_{x} \int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{G}^{\mathrm{S}}(\mathbf{x}-\boldsymbol{\alpha}) \cdot \boldsymbol{\nu}(\boldsymbol{\alpha})=\frac{1}{2} \boldsymbol{\sigma}_{3} \cdot \boldsymbol{\nu}(\boldsymbol{\beta}) \\
& \quad+\int_{\partial \mathcal{V}} d \sigma_{\alpha} \mathbf{n}_{\beta} \cdot \nabla_{\beta} \mathrm{G}^{\mathrm{S}}(\boldsymbol{\beta}-\boldsymbol{\alpha}) \cdot \boldsymbol{\nu}(\boldsymbol{\alpha}) . \tag{A.37}
\end{align*}
$$

The important point here is that all discontinuities of these parametrizations are generated solely by components of $\mathbf{G}^{\mathrm{N}, \mathrm{S}}$ proportional to $\boldsymbol{\sigma}_{3}$. Other components of $\mathbf{G}^{\mathrm{N}, \mathrm{S}}$ are proportional to the scalar Green function composition $\left(g_{+}^{\mathrm{N}, \mathrm{S}}-g_{-}^{\mathrm{N}, \mathrm{S}}\right)$, and whatever discontinuity might be generated by the + term is cancelled by a corresponding one generated by the - term. Therefore the discontinuities generated by the parametrizations have the forms given in Eqs. (A.34) and (A.37).

## Appendix B

## (d - 1)-dimensional Fourier

## transforms

In this appendix we introduce the $d-1$ dimensional Fourier transform for functions whose arguments are $d$ dimensional vectors, which we use extensively throughout the text. Mainly, our functions of interest will be the Green functions $g_{ \pm}^{\mathrm{N}, \mathrm{S}}$ and the functions related to them, such as $\partial g_{ \pm}^{\mathrm{N}, \mathrm{S}}$.

The $d-1$ dimensional Fourier transform (and its inverse) of a function $f(\mathbf{x})$ of a $d$ dimensional vector $\mathbf{x}$, are defined by

$$
\begin{align*}
\tilde{f}(\boldsymbol{\kappa}, z) & =\int_{\mathcal{P}} d \mathbf{x} / / \mathrm{e}^{i \boldsymbol{\kappa} \cdot \mathbf{x}_{/ /}} f(\mathbf{x})  \tag{B.1}\\
f(\mathbf{x}) & =(2 \pi)^{-d+1} \int d \boldsymbol{\kappa} \mathrm{e}^{-i \boldsymbol{\kappa} \cdot \mathbf{x}_{/ /}} \tilde{f}(\boldsymbol{\kappa}, z) \tag{B.2}
\end{align*}
$$

where $\left(\boldsymbol{x}_{/ /}, z\right) \equiv \boldsymbol{x}$, and $\boldsymbol{x}_{/ /}$and $\boldsymbol{\kappa}$ are vectors on the $d-1$ dimensional hyperplane $\mathcal{P}$ perpendicular to the $z$ axis. We now evaluate the $(d-1)$-dimensional Fourier transforms of the functions that are used throughout this Paper. We begin with the Helmholtz Green function $g\left(\boldsymbol{x} ; k^{2}\right)$ :

$$
\begin{aligned}
\tilde{g}\left(\boldsymbol{\kappa}, z ; k^{2}\right) & =\int_{\mathcal{P}} d \mathbf{x}_{/ /} \mathrm{e}^{i \boldsymbol{\kappa} \cdot \mathbf{x}_{/ /}} g\left(\mathbf{x} ; k^{2}\right)=\int_{\mathcal{P}} d \mathbf{x}_{/ /} \mathrm{e}^{i \boldsymbol{\kappa} \cdot \mathbf{x}_{/ /}} \int \frac{d \mathbf{p}}{(2 \pi)^{d}} \frac{\mathrm{e}^{i \mathbf{p} \cdot \mathbf{x}}}{p^{2}-k^{2}} \\
& =\int \frac{d \mathbf{p}}{(2 \pi)^{d}} \int_{\mathcal{P}} d \mathbf{x}_{/ /} \mathrm{e}^{i(\boldsymbol{\kappa}+\mathbf{p} / /) \cdot \mathbf{x}_{/ /}} \frac{\mathrm{e}^{i p_{z} z}}{p^{2}-k^{2}}
\end{aligned}
$$

$$
\begin{align*}
& =\int \frac{d \mathbf{p}}{(2 \pi)^{d}}(2 \pi)^{(d-1)} \delta\left(\boldsymbol{\kappa}+\mathbf{p}_{/ /}\right) \frac{\mathrm{e}^{i p_{z} z}}{p^{2}-k^{2}} \\
& =\int_{-\infty}^{\infty} \frac{d p_{z}}{2 \pi} \frac{\mathrm{e}^{i p_{z} z}}{p_{z}^{2}+\kappa^{2}-k^{2}}=\frac{\mathrm{e}^{-a(\kappa)|z|}}{2 a(\kappa)}, \tag{B.3}
\end{align*}
$$

where $a(\kappa) \equiv \sqrt{\kappa^{2}-k^{2}}$ such that $\operatorname{Re}(a(\kappa))>0$. The evaluation of the $(d-1)$ dimensional Fourier transform of $\partial \delta g\left(\boldsymbol{x} ; k^{2}\right)$ is very similar, except that now $\mathbf{x}=$ $\left(\mathrm{x}_{/ /}, 0\right)$ (i.e. both arguments of the Green function lie on $\mathcal{P}$ and the normal direction is $z$ ):

$$
\begin{align*}
\partial \delta \tilde{g}\left(\boldsymbol{\kappa} ; k^{2}\right) & =\int_{\mathcal{P}} d \mathbf{x}_{/ /} \mathrm{e}^{i \boldsymbol{\kappa} \cdot \mathbf{x}_{/ /}} \partial \delta g\left(\left(\mathbf{x}_{/ /}, 0\right) ; k^{2}\right) \\
& =\left.\int_{\mathcal{P}} d \mathbf{x}_{/ /} \mathrm{e}^{i \boldsymbol{\kappa} \cdot \mathbf{x}_{/ /}}\left(-\frac{d^{2}}{d z^{2}} \int \frac{d \mathbf{p}}{(2 \pi)^{d}} \frac{\mathrm{e}^{i \mathbf{p} / / \cdot \mathbf{x}_{/ /}+i p_{z} z}}{p^{2}-k^{2}}\right)\right|_{z=0} \\
& =\int_{-\infty}^{\infty} \frac{d p_{z}}{2 \pi} \frac{p_{z}^{2} \mathrm{e}^{i p_{z} z}}{p_{z}^{2}+a(\kappa)^{2}}=-\frac{a(\kappa)}{2} \tag{B.4}
\end{align*}
$$

## Appendix C

## Propagation outside the normal region and nonconvex shapes: cancellations

In the present section we study cancellations between terms in the MSE, which occur, in the large $k_{\mathrm{F}} L$ limit, when one (or more) of the homogeneous region(s) constituting the billiard are nonconvex. In the sample billiard shown in Fig. 3.1 the S region is nonconvex, whereas the N region is convex. Another example in which such cancellations arise is provided by antidot-billiard geometries, in which $S$ regions are embedded in an N region. For such cases, the N region is certainly nonconvex, and so may be the S region. In the former case (i.e. convex N ; nonconvex S ), these cancellations eliminate terms that include finite-outside-propagation, validating the claim made in Sec. 5.4 that periodic orbits that include such paths do not contribute at leading order to the semiclassical DOS oscillations.

In order to bring to the fore the physics underlying these cancellations, consider a much simpler case: the MRE for the Helmholtz Green function satisfying homogeneous Dirichlet boundary conditions [49]. By applying the methods described here,


Figure C.1: Example of a nonconvex inside and outside regions. Also shown is a propagation directly from $\boldsymbol{\alpha}$ to $\boldsymbol{\beta}$ and two-reflection correction.
one can express this Green function in terms of the following MRE:

$$
\begin{align*}
g^{\mathrm{D}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=g^{\mathrm{H}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)- & 2 \int_{\partial \mathcal{\nu}} \partial g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\alpha}) g^{\mathrm{H}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) \\
& +4 \int_{\partial \mathcal{V}} \partial g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\beta}) \partial g^{\mathrm{H}}(\boldsymbol{\beta}, \boldsymbol{\alpha}) g^{\mathrm{H}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right)+\cdots, \tag{C.1}
\end{align*}
$$

where $g^{\mathrm{H}}$ is the homogeneous Helmholtz Green function. If the domain $\mathcal{V}$ over which $g^{\mathrm{D}}$ is defined is nonconvex then one may be concerned by the presence of amplitudes involving propagations between pairs of boundary points for which all or part of this propagation lies outside of $\mathcal{V}$ (see Fig. C.1). On physical grounds, one expects that in the large $k L$ limit the obstacle between these points would suppress such amplitudes. Balian and Bloch showed that this is indeed the case. To see this, we follow Balian and Bloch and consider $g^{\mathrm{D}}$ for a planar boundary. Then the MRE (C.1) terminates at the second term:

$$
\begin{equation*}
g^{\mathrm{D}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=g^{\mathrm{H}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)-2 \int_{\partial \mathcal{V}} \partial g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\alpha}) g^{\mathrm{H}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) . \tag{C.2}
\end{equation*}
$$

On the other hand, by using the method of images one has

$$
\begin{equation*}
g^{\mathrm{D}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=g^{\mathrm{H}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)-g^{\mathrm{H}}\left(\mathbf{x}, \tilde{\mathbf{x}}^{\prime}\right) \tag{C.3}
\end{equation*}
$$

where $\tilde{\mathbf{x}}^{\prime}$ is the mirror image of $\mathbf{x}^{\prime}$ with respect to the planar boundary. Comparison of Eq. (C.3) with Eq. (C.2) produces the relation

$$
\begin{equation*}
g^{\mathrm{H}}\left(\mathbf{x}, \tilde{\mathbf{x}}^{\prime}\right)=2 \int_{\partial \mathcal{V}} \partial g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\alpha}) g^{\mathrm{H}}\left(\boldsymbol{\alpha}, \mathbf{x}^{\prime}\right) \tag{C.4}
\end{equation*}
$$

As the length $\left|\boldsymbol{\alpha}-\mathbf{x}^{\prime}\right|$ equals the length $\left|\boldsymbol{\alpha}-\tilde{\mathbf{x}}^{\prime}\right|$, one can replace $\mathbf{x}^{\prime}$ by $\tilde{\mathbf{x}}^{\prime}$ in relation (C.4) and obtain a second relation

$$
\begin{equation*}
g^{\mathrm{H}}\left(\mathbf{x}, \tilde{\mathbf{x}}^{\prime}\right)=2 \int_{\partial \mathcal{V}} \partial g^{\mathrm{H}}(\mathbf{x}, \boldsymbol{\alpha}) g^{\mathrm{H}}\left(\boldsymbol{\alpha}, \tilde{\mathbf{x}}^{\prime}\right) \tag{C.5}
\end{equation*}
$$

Now suppose $\partial \mathcal{V}$ is not planar. Although relations (C.4) and (C.5) no longer hold, in the large $k L$ limit the dominant contribution comes from the stationary-phase point and, thus, Eq. (C.5) becomes exact as $k L$ tends to infinity. That Eq. (C.4) is not exact in this limit is due to the fact that the fluctuation determinant in this case (viz. when $\mathbf{x}$ and $\mathbf{x}^{\prime}$ are on the same side of the surface) depends on the curvature of the surface at the stationary-phase point.

Now let us return to what happens for nonconvex surfaces. In particular, consider a term in MRE in which part of the amplitude has propagation between boundary points $\boldsymbol{\alpha}$ and $\boldsymbol{\beta}$, where part of the line joining $\boldsymbol{\alpha}$ and $\boldsymbol{\beta}$ lies outside $\mathcal{V}$ and intersects $\partial \mathcal{V}$ at the points $\gamma_{0}$ and $\boldsymbol{\delta}_{0}$ (see Fig. C.1). In relation to this term, consider three related terms with higher numbers of reflections: (i) the term with one more reflection near $\gamma_{0}$, (ii) the term with one more reflection near $\boldsymbol{\delta}_{0}$, and (iii) the term with two more reflections near $\gamma_{0}$ and $\boldsymbol{\delta}_{0}$. In the MRE, all four are in the sum. The sum of the part of the amplitude involving direct propagation from $\boldsymbol{\alpha}$ to $\boldsymbol{\beta}$ and the one with one more reflection near $\gamma_{0}$ is

$$
\begin{equation*}
\partial g^{\mathrm{H}}(\boldsymbol{\alpha}, \boldsymbol{\beta})-2 \int_{\partial \mathcal{V}} \partial g^{\mathrm{H}}(\boldsymbol{\alpha}, \boldsymbol{\gamma}) \partial g^{\mathrm{H}}(\boldsymbol{\gamma}, \boldsymbol{\beta}) \tag{C.6}
\end{equation*}
$$

By virtue of Eq. (C.5) this sum vanishes in the $k L \rightarrow \infty$ limit. The same holds for the sum of remaining two terms. Thus in the limit $k L \rightarrow \infty$ sum of all terms involving paths that lie partially outside of $\mathcal{V}$ vanishes.

These considerations extend readily to Andreev billiards. In order to see this, we must extend the identity (C.5). This can be achieved as follows:

$$
\begin{align*}
2 \int_{\partial \mathcal{V}} \partial \mathbf{G}^{\mathrm{N}}(\mathbf{x}, \boldsymbol{\alpha}) \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{N}}(\boldsymbol{\alpha}, \tilde{\mathbf{x}}) & =2 \int_{\partial \mathcal{L}}\left(\begin{array}{cc}
\partial g_{+}(\mathbf{x}, \boldsymbol{\alpha}) g_{+}(\boldsymbol{\alpha}, \tilde{\mathbf{x}}) & 0 \\
0 & -\partial g_{-}(\mathbf{x}, \boldsymbol{\alpha}) g_{-}(\boldsymbol{\alpha}, \tilde{\mathbf{x}})
\end{array}\right) \\
& =\left(\begin{array}{cc}
g_{+}(\mathbf{x}, \tilde{\mathbf{x}}) & 0 \\
0 & -g_{-}(\mathbf{x}, \tilde{\mathbf{x}})
\end{array}\right)=\mathbf{G}^{\mathrm{N}}(\mathbf{x}, \tilde{\mathbf{x}}), \tag{C.7}
\end{align*}
$$

where, as in the case of Helmholtz Green functions, $\partial \mathcal{V}$ is a planar surface and $\mathbf{x}$ and $\tilde{\mathbf{x}}$ lie on different sides of $\partial \mathcal{V}$, and in going to the second line we have made use of Eq. (C.5). A similar identity holds for $G^{S}$

$$
\begin{equation*}
\mathbf{G}^{\mathrm{S}}(\mathbf{x}, \tilde{\mathbf{x}})=2 \int_{\partial \mathcal{V}} \partial \mathbf{G}^{\mathrm{S}}(\mathbf{x}, \boldsymbol{\alpha}) \boldsymbol{\sigma}_{3} \mathbf{G}^{\mathrm{S}}(\boldsymbol{\alpha}, \tilde{\mathbf{x}})=2 \int_{\partial \mathcal{V}} \mathrm{G}^{\mathrm{S}}(\tilde{\mathbf{x}}, \boldsymbol{\alpha}) \boldsymbol{\sigma}_{3} \delta \mathbf{G}^{\mathrm{S}}(\boldsymbol{\alpha}, \mathbf{x}) \tag{C.8}
\end{equation*}
$$

This identity can be obtained by performing the matrix multiplication and using Eq. (C.5) in each matrix element. As in the case of Helmholtz Green functions, identities (C.7) and (C.8) hold in the large $k_{\mathrm{F}} L$ limit, even if $\partial \mathcal{V}$ is not planar. As we did for the Helmholtz case, consider a term in MSE in which part of the amplitude has propagation between boundary points $\boldsymbol{\alpha}$ and $\boldsymbol{\beta}$, where part of the line joining $\boldsymbol{\alpha}$ and $\boldsymbol{\beta}$ lies outside $\mathcal{V}$ and intersects $\partial \mathcal{V}$ at the points $\boldsymbol{\gamma}_{0}$ and $\boldsymbol{\delta}_{0}$. However, now we have additional terms owing to the possibility of propagation between any two boundary points involving either $\mathbf{G}^{\mathrm{N}}$ or $\mathbf{G}^{\mathrm{S}}$. By using the identities (C.7) and (C.8) it is not hard to see that, in the large $k_{\mathrm{F}} L$ limit, a propagation from $\boldsymbol{\alpha}$ to $\boldsymbol{\beta}$ is not cancelled only if the line joining these points lies totally in either $\mathcal{V}$ or $\overline{\mathcal{V}}$, with the propagation involving the homogeneous Green function appropriate to the corresponding region. Thus, in the semiclassical limit the terms that survive in the MSE consist of pure reflection or pure transmission/tunneling. For the example at hand, the surviving term would be the one with normal propagation from $\boldsymbol{\alpha}$ to $\gamma$, superconducting propagation from $\gamma$ to $\boldsymbol{\delta}$ (i.e. tunneling), and then normal propagation from $\boldsymbol{\delta}$ to $\boldsymbol{\beta}$

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[38] They are the leading order terms in the Weyl (or Thomas-Fermi) expansion for the Andreev billiard.
[39] Magnetic fields introduce similar chaos; see [5].
[40] Note that in settings that mix superconducting confinement with conventional confinement due to a single-particle potential, it is possible (such as when a chaotic conventional billiard is in proximity of a superconductor) that the lengthscale of the primitive periodic orbits be much larger than the linear size of the billiard. In such cases, Scheme A breaks down and Scheme B should be used.
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[102] Note that we are ignoring broader classes of fluctuations, such as those that would take $\bar{\Delta}(\mathbf{k}, \mathbf{R})$ out of the space of $\mathrm{d}_{x^{2}-y^{2}}$ states. For example, one can envision the situation in which phase-fluctuating d-wave order is augmented by fluctuating local s-wave order. Additional classes of fluctuations, such as this, can readily be incorporated by the scheme discussed in the present Paper, and would lead to further broadening of the single-particle spectral function.
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a doubly-scattering trajectory would provide two states, both very close indeed to zero energy.
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